



Non-linear 3D Hybrid Kinetic-MHD Simulations of ELMs in the ASDEX Upgrade Tokamak with MEGA

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M. Garcia-Munoz, M. Toscano-Jimenez, E. Viezzer, Y. Suzuki,
Y. Todo, J. Galdon-Quiroga, M. Hoelzl, P. Oyola, J. Rivero-Rodriguez,
C. Soria-Hoyo, ASDEX Upgrade and EUROfusion MST1 Teams



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Outline

- Motivation
- 3D non-linear hybrid kinetic-MHD MEGA
- Non-linear MHD simulations of ELMs
- Fast-ion effects on ELM stability
- Summary and outlook



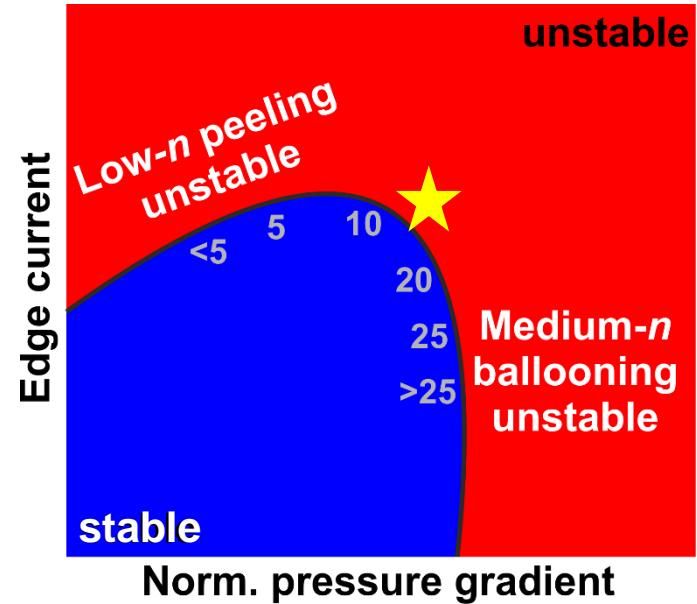
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Motivation

- ELMs appear in H-mode
→ Peeling-balloonning unstable

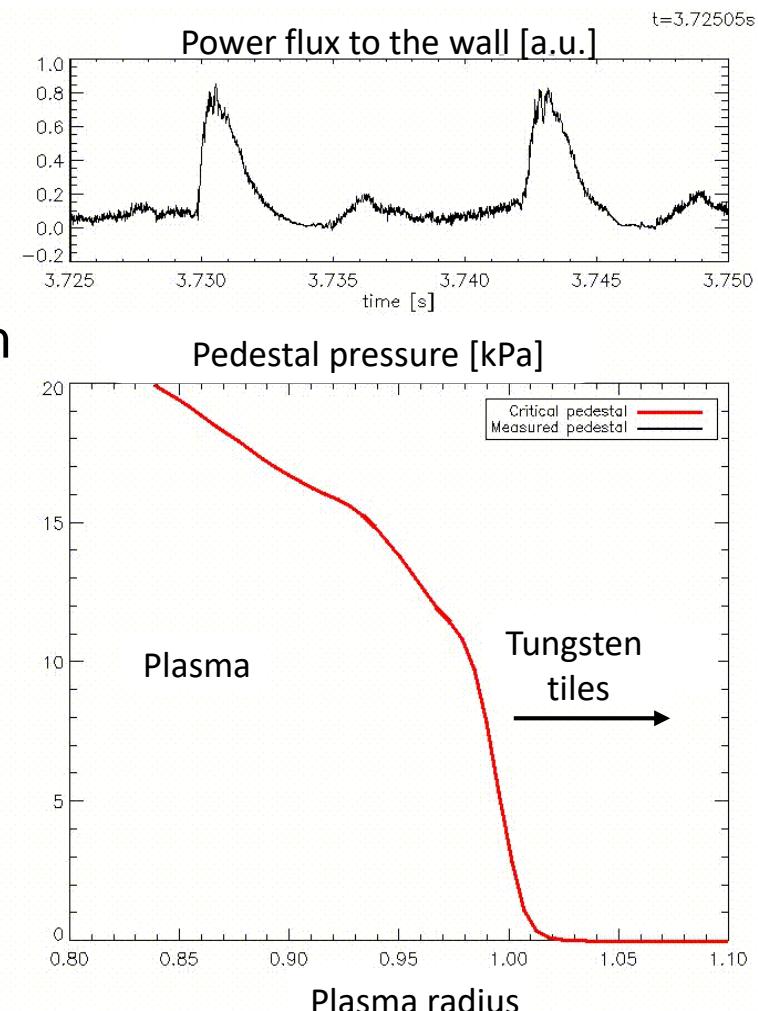


*E. Viezzer



Motivation

- ELMs appear in H-mode
→ Peeling-balloonning unstable
- ELMs expel particles and energy from plasma
→ Degradation of pedestal
→ Untolerable for future fusion devices

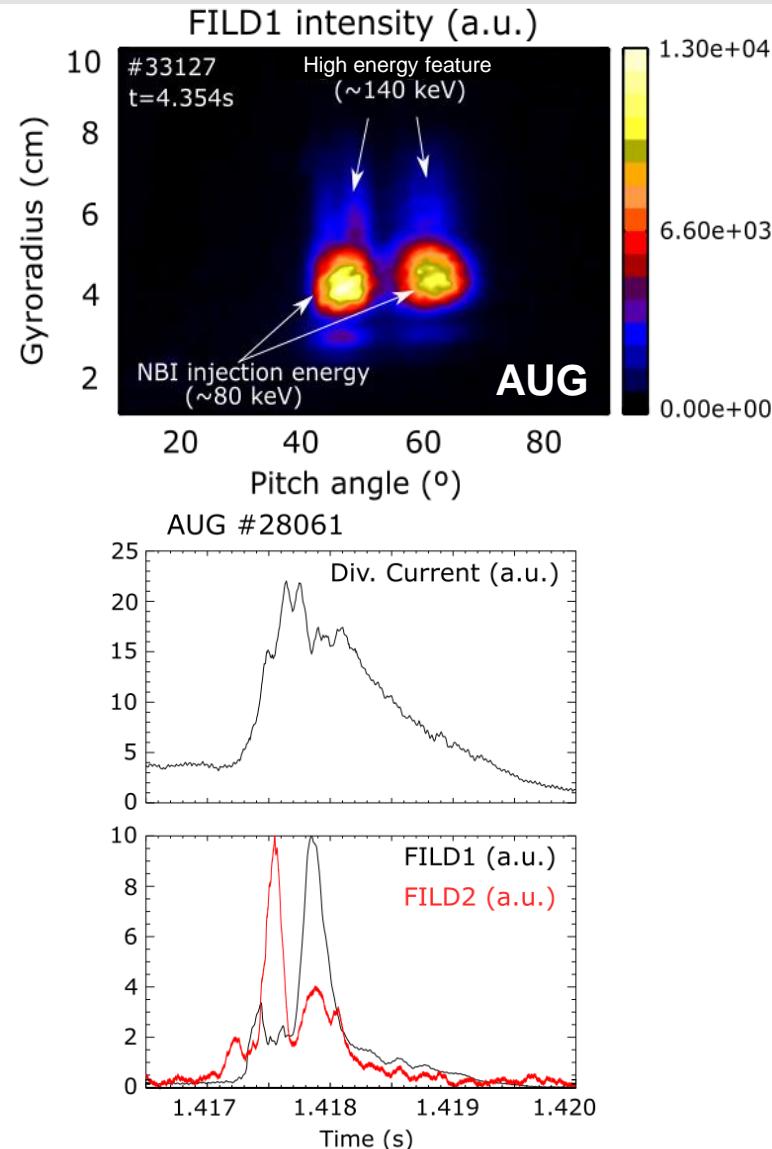


*A. Burckhart



Motivation

- ELMs appear in H-mode
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- Recent observations indicate strong interaction between ELMs and fast-ions



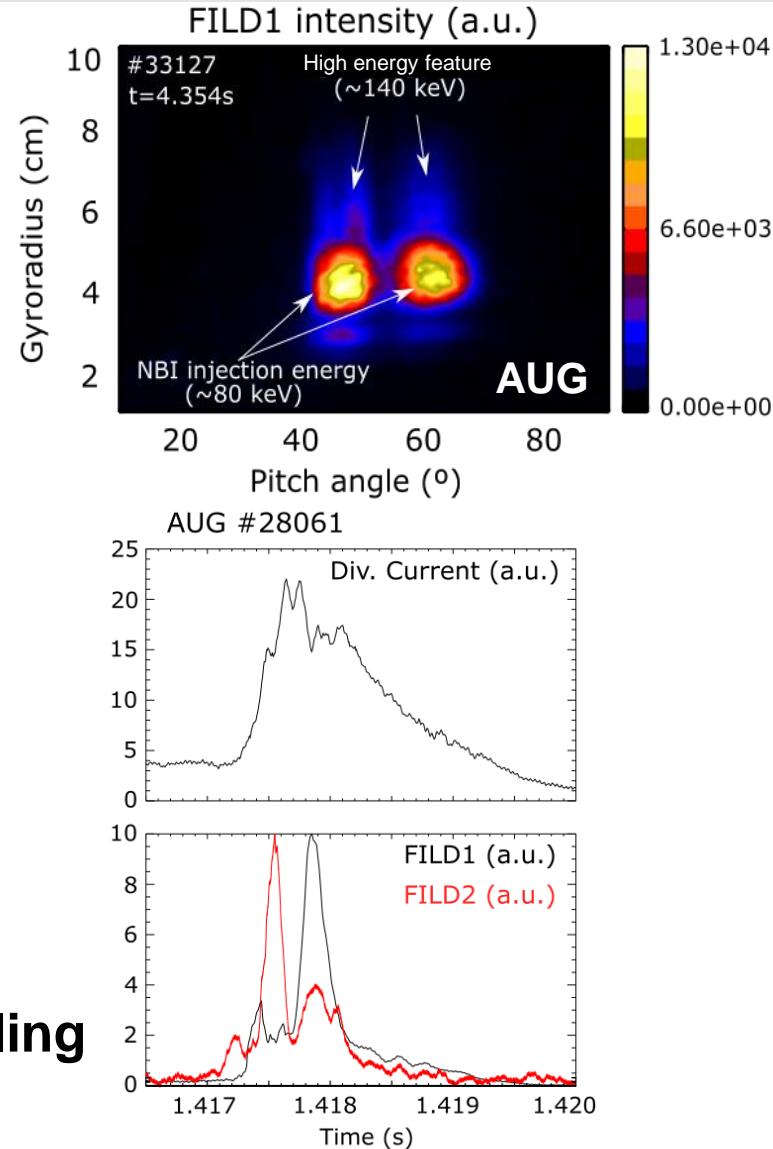
**See J.F. Rivero-Rodriguez (I-8, Wed 11:00)

*J. Galdon-Quiroga et al., 2018, Phys. Rev. Lett. **121**, 025002



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→ Peeling-balloonning unstable
- ELMs expel particles and energy from plasma
→ Degradation of pedestal
→ Untolerable for future fusion devices
- Recent observations indicate strong interaction between ELMs and fast-ions
- **Kinetic effects needed in ELM modelling**



**See J.F. Rivero-Rodriguez (I-8, Wed 11:00)

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3D Non-Linear Hybrid Kinetic-MHD Modelling of ELMs with MEGA



- Non-linear hybrid kinetic-MHD code^[1]: EP and MHD dynamics coupled through EP current density as follows

$$\begin{aligned}\rho \frac{\partial \vec{v}}{\partial t} &= -\rho(\vec{v} \cdot \nabla)\vec{v} - \nabla p + (\vec{j} - \vec{j}_{h'}) \times \vec{B} - \nabla \times (\nu \rho \nabla \times \vec{v}) + \frac{4}{3} \nabla(\nu \rho \nabla \cdot \vec{v}) \\ \frac{\partial p}{\partial t} &= -\nabla \cdot (p \vec{v}) - (\gamma - 1)p \nabla \cdot \vec{v} + \nabla \cdot [\chi_{\perp} \nabla_{\perp}(p - p_{\text{eq}}) + \chi_{\parallel} \nabla_{\parallel}(p - p_{\text{eq}})] \\ &\quad + (\gamma - 1) [\nu \rho (\nabla \times \vec{v})^2 + \frac{4}{3} \nu \rho (\nabla \cdot \vec{v})^2 + \eta (\vec{j} - \vec{j}_{h'}) \cdot (\vec{j} - \vec{j}_{\text{eq}})]\end{aligned}$$

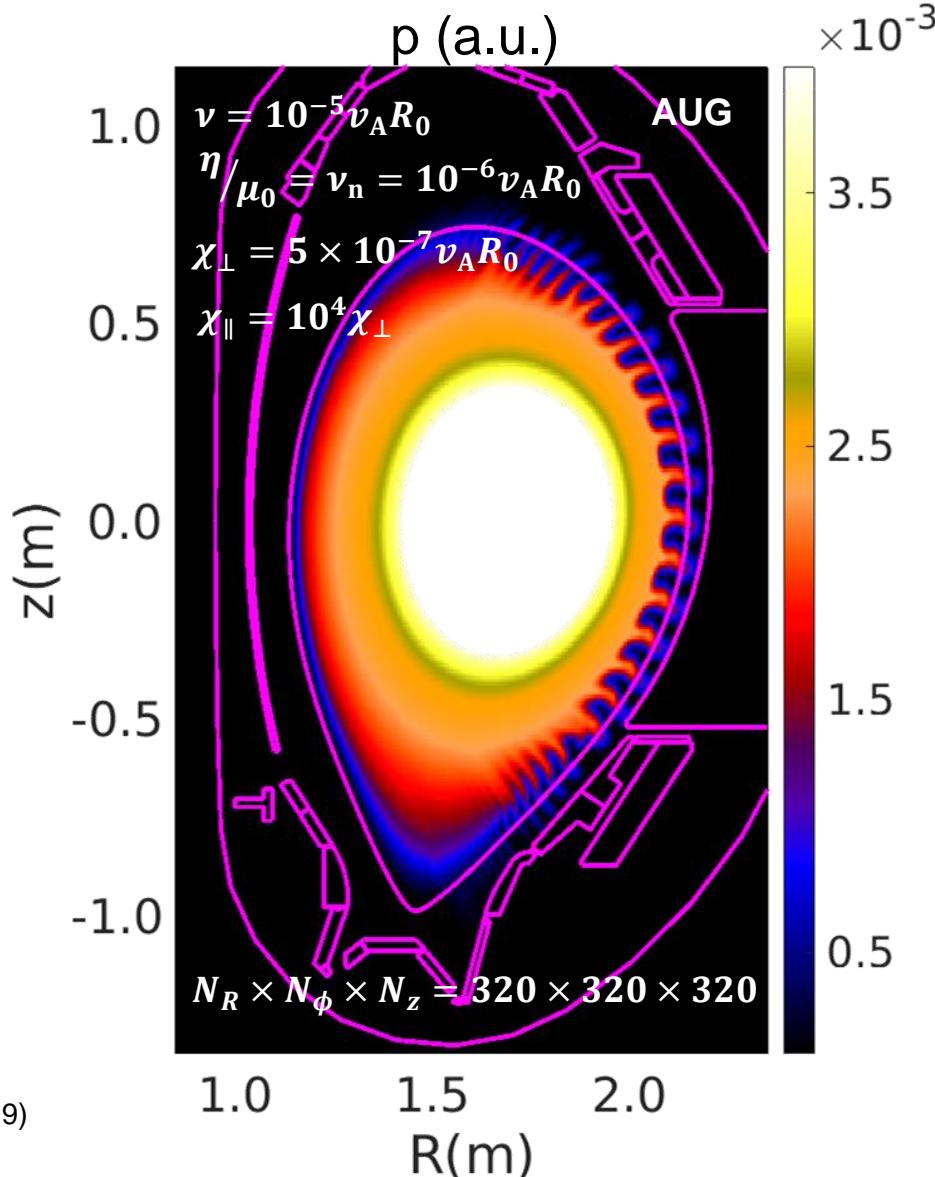
- What do we need to consider for ELM simulations?

[1] Y. Todo *et al.*, Phys. Plasmas 5, 1321 (1998)

3D Non-Linear Hybrid Kinetic-MHD Modelling of ELMs with MEGA



- Cylindrical coordinates (R, ϕ, z)
- Fully 3D rectangular geometry
- SOL and Private Flux Region below X-Point included in simulation domain^[1]
 - ELM relevant area
 - Important to study the interaction with fast-ions

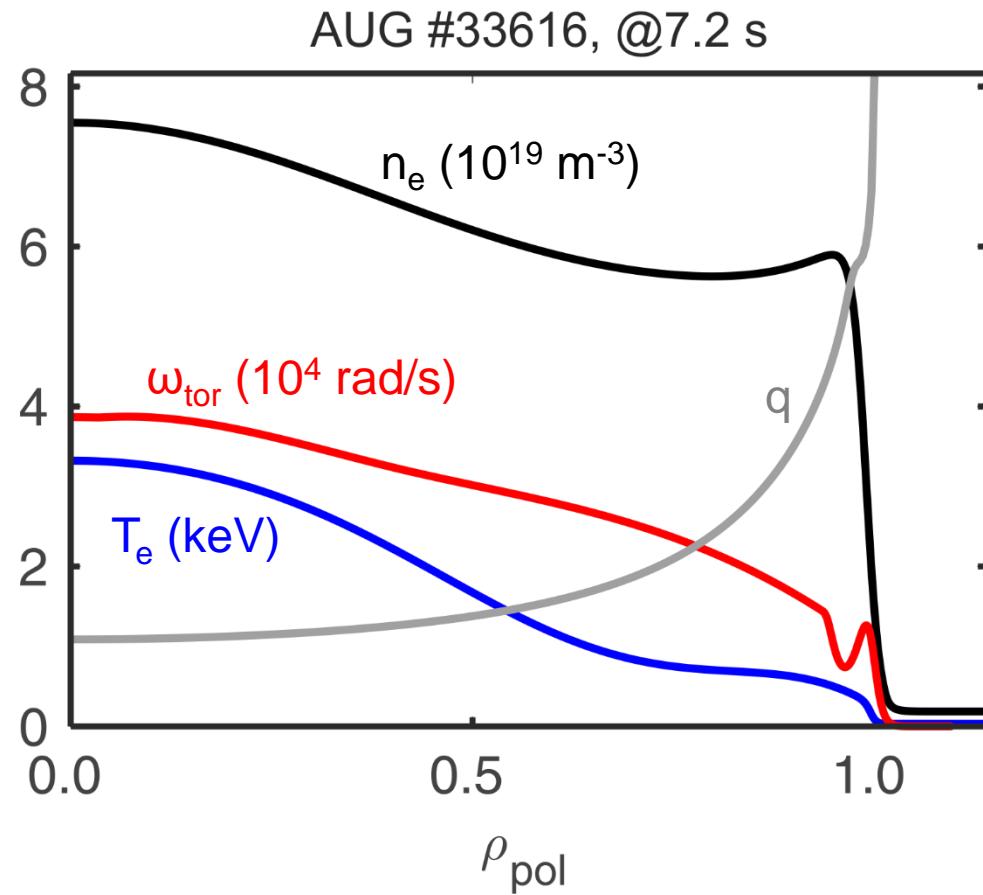


[1] S. Futatani *et al.*, Plasma Phys. Control. Fusion 61, 095014 (2019)

3D Non-Linear Hybrid Kinetic-MHD Modelling of ELMs with MEGA



- Kinetic and rotation profiles of ASDEX Upgrade shot #33616 used as initial conditions
- Simulations include $n < 20$ (experimentally, dominant mode $n \sim 2 - 5$ during ELM crash^[1])
- Single n simulations to calculate linear growth rates, multi n for non-linear phase
- Standard MHD model excluding diamagnetic, toroidal and neoclassical flows



[1] A.F. Mink *et al.*, Nucl. Fusion 58, 026011 (2018)



Outline

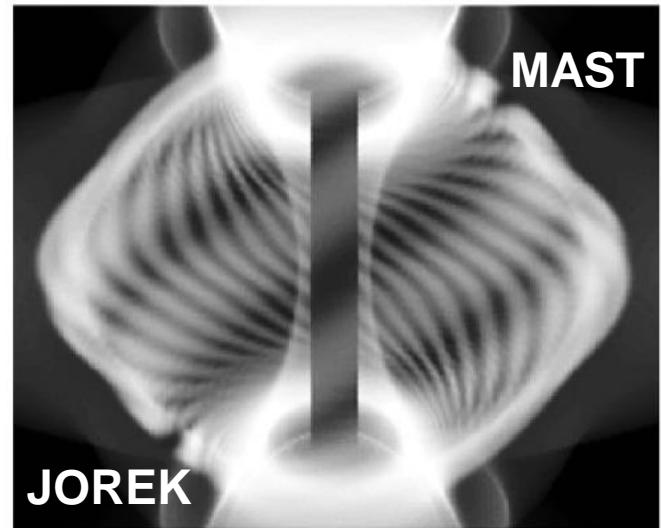
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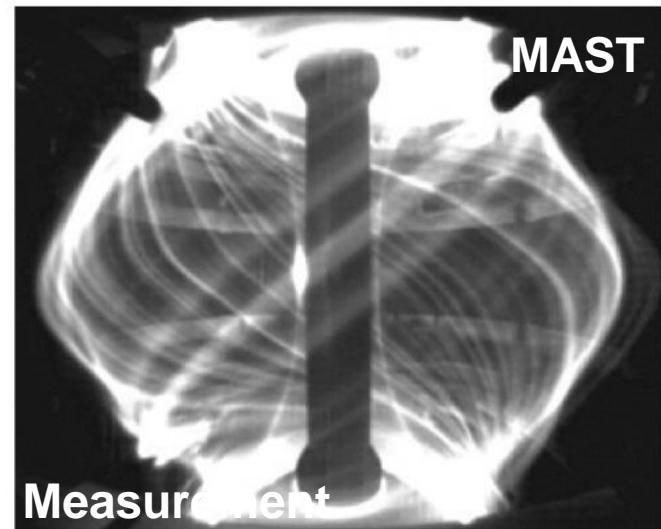
How to Identify an ELM in MEGA

- In standard MHD simulations, ELMs are characterized by^[1]:
 - High n ballooning modes when using standard MHD
 - Non-linear growth of low n harmonics
 - Filamentary structure
 - Relaxation of profiles

D _{α} image of an ELM in MAST^[2]



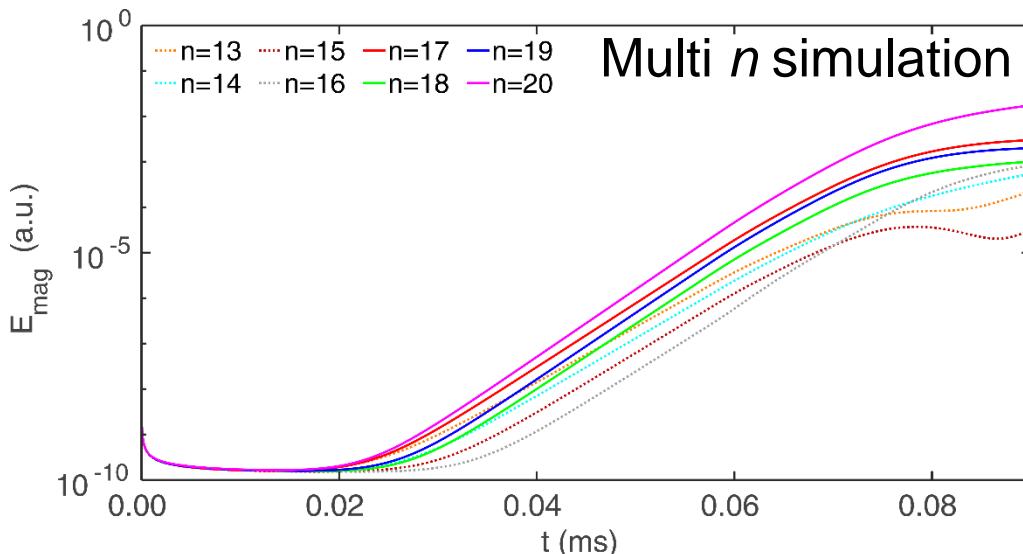
JOREK



[1] G.T.A. Huijsmans *et al.*, Phys. Plasmas 22, 021805 (2015)

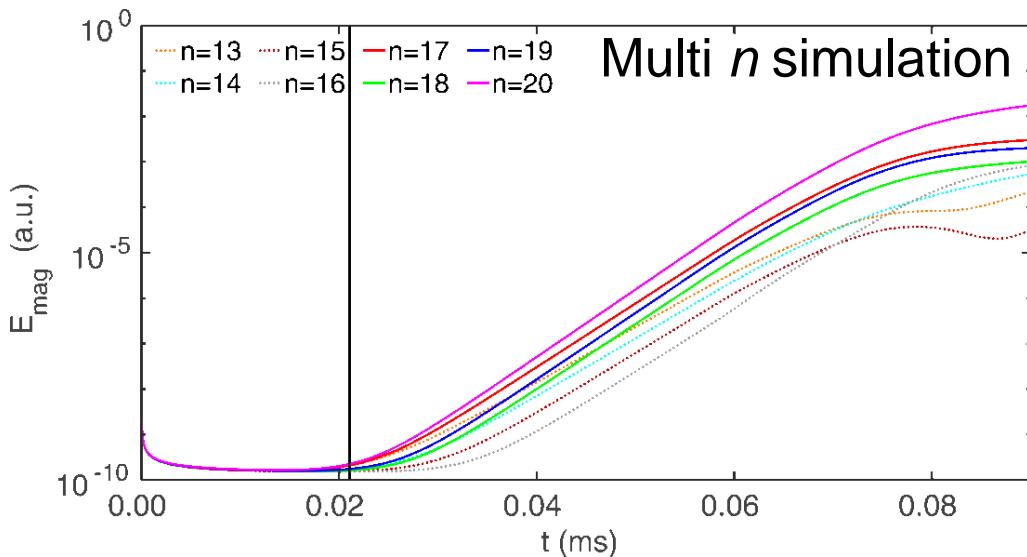
[2] S.J.P. Pamela *et al.*, Plasma Phys. Control. Fusion 55, 095001 (2013)

Ballooning Mode Structure Observed in Linear Phase

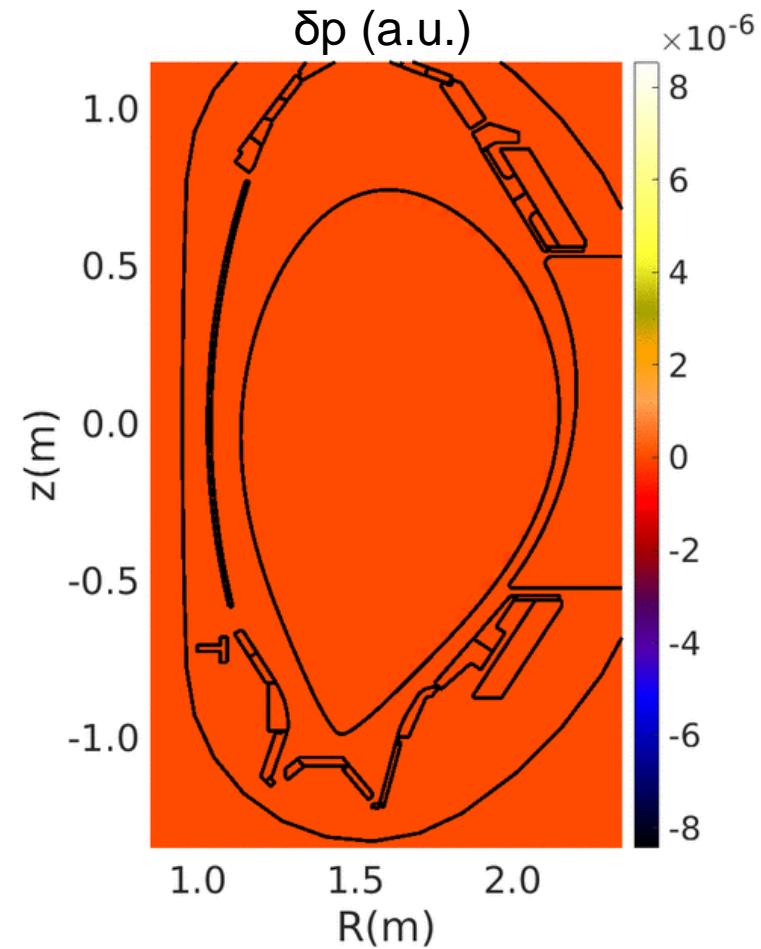


- High n ($n>13$) modes dominate linear phase

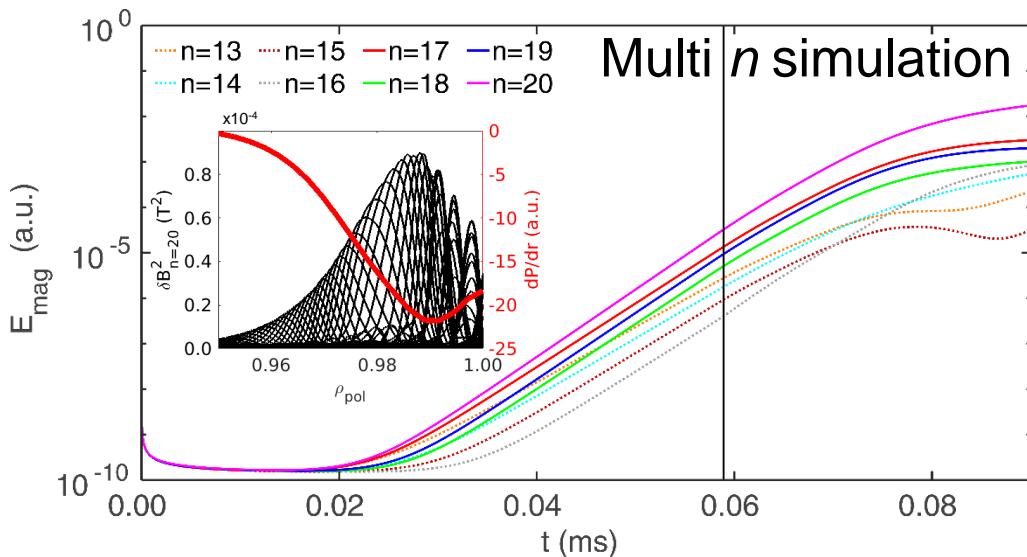
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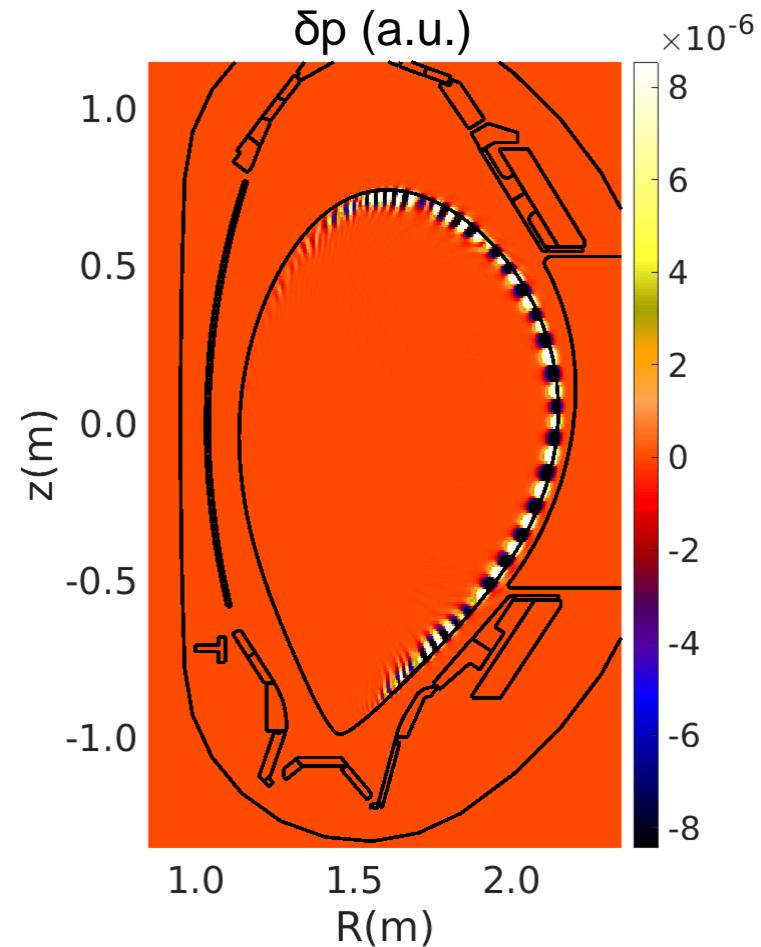
- High n ($n > 13$) modes dominate linear phase
- Perturbation located at LFS (maximum pressure gradient)



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- Perturbation located at LFS (maximum pressure gradient)



➤ Ballooning modes

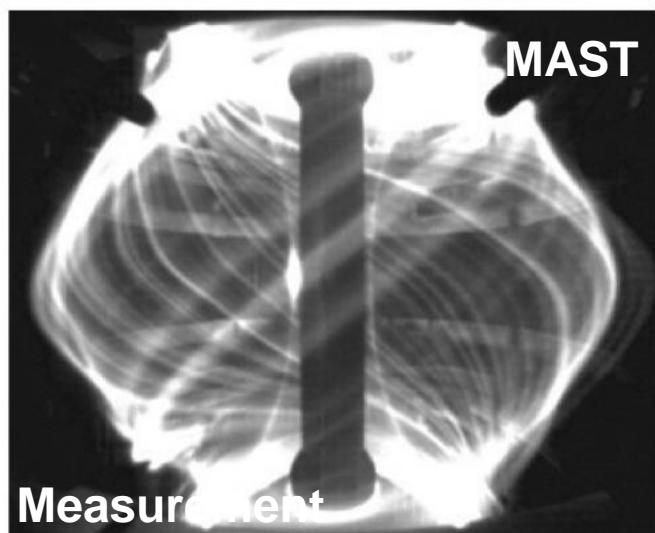
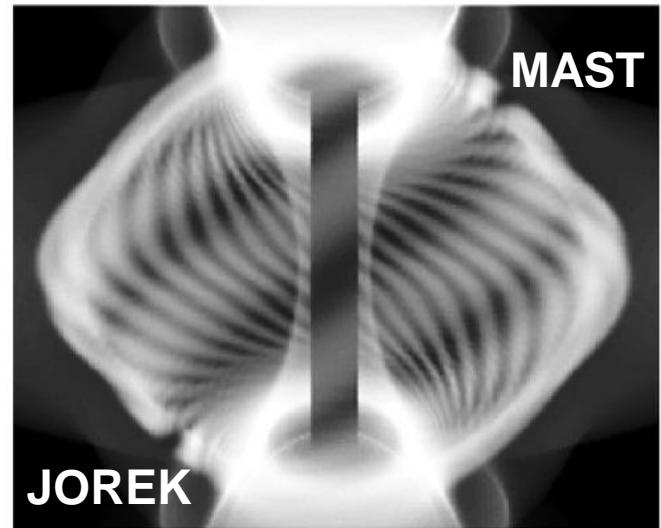


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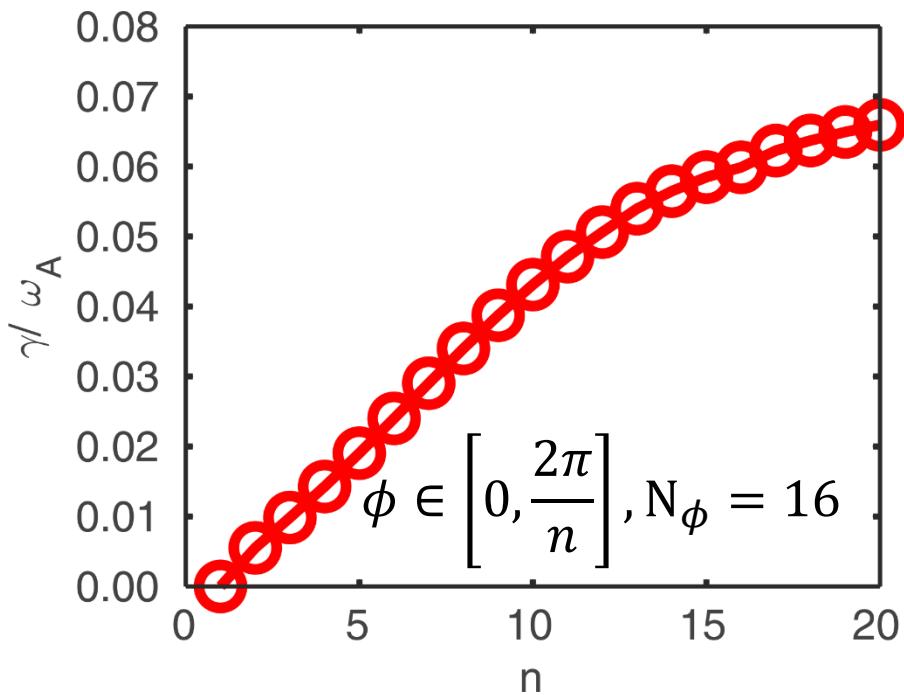
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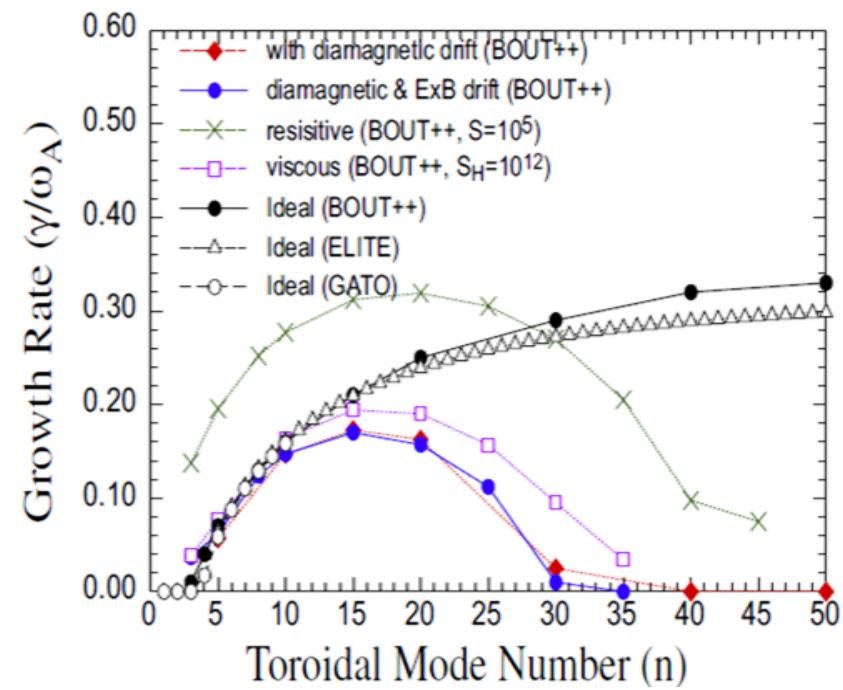
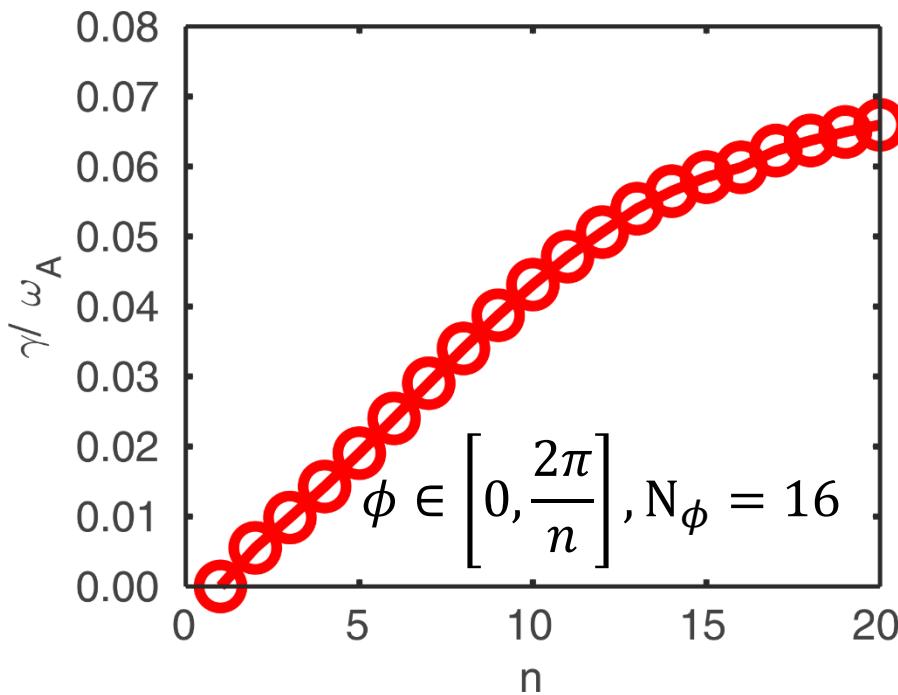
[2] S.J.P. Pamela *et al.*, Plasma Phys. Control. Fusion 55, 095001 (2013)

High n Instabilities Dominate Linear Phase in Standard MHD



- Higher linear growth rates for higher toroidal mode numbers

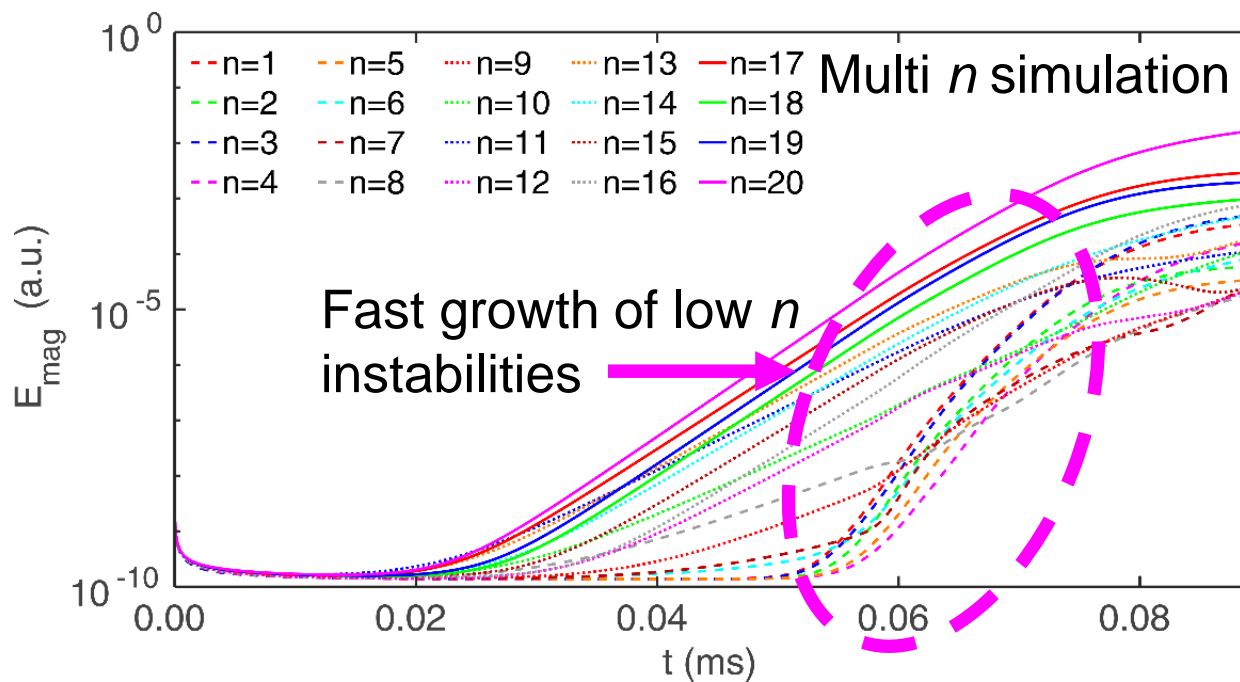
High n Instabilities Dominate Linear Phase in Standard MHD



- Higher linear growth rates for higher toroidal mode numbers
- Results are in agreement with infinite n ballooning theory^[1]

[1] X.Q. Xu *et al.*, Nucl. Fusion 51, 103040 (2011)

Low n Modes Are Growing Faster Due to Non-Linear Coupling



- Linear phase in multi n simulation dominated by high n harmonics
- Fast non-linear growth of low n modes due to non-linear coupling^[1]

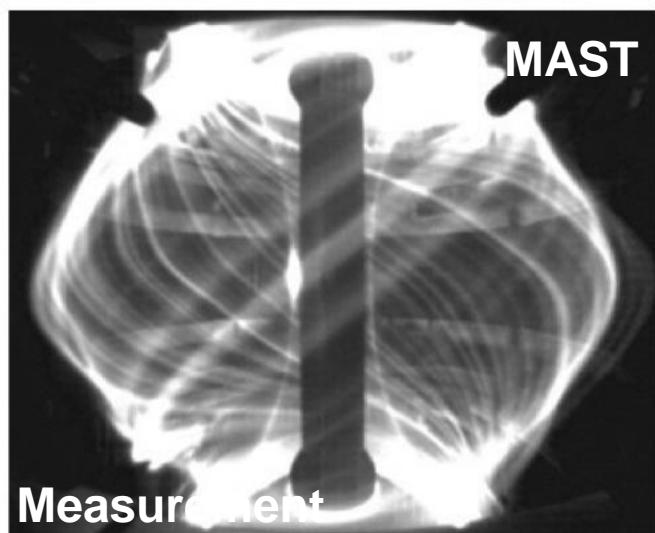
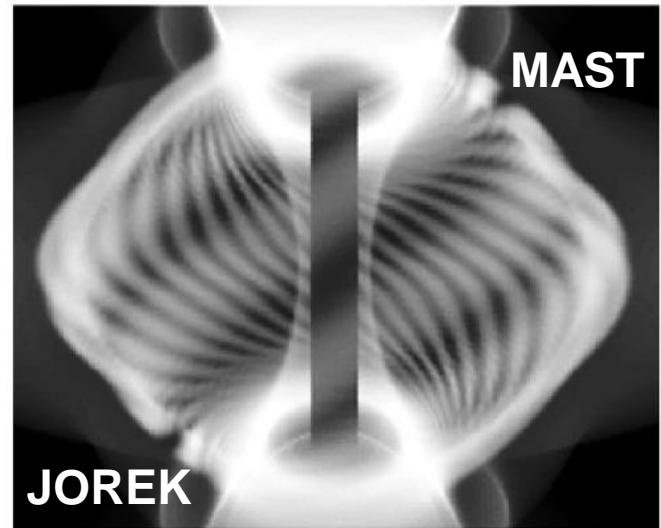
[1] I. Krebs *et al.*, Phys. Plasmas 20, 082506 (2013)



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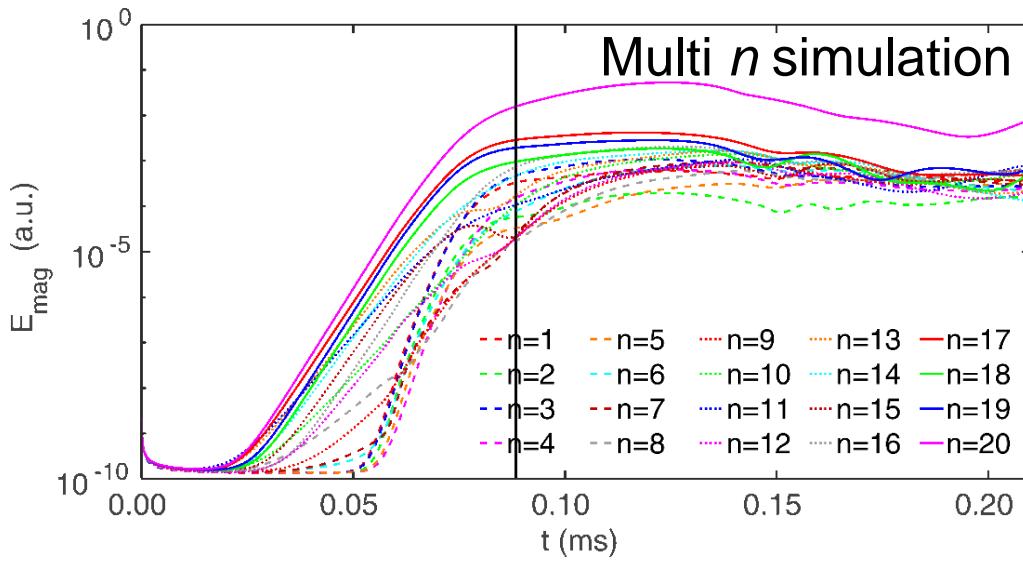
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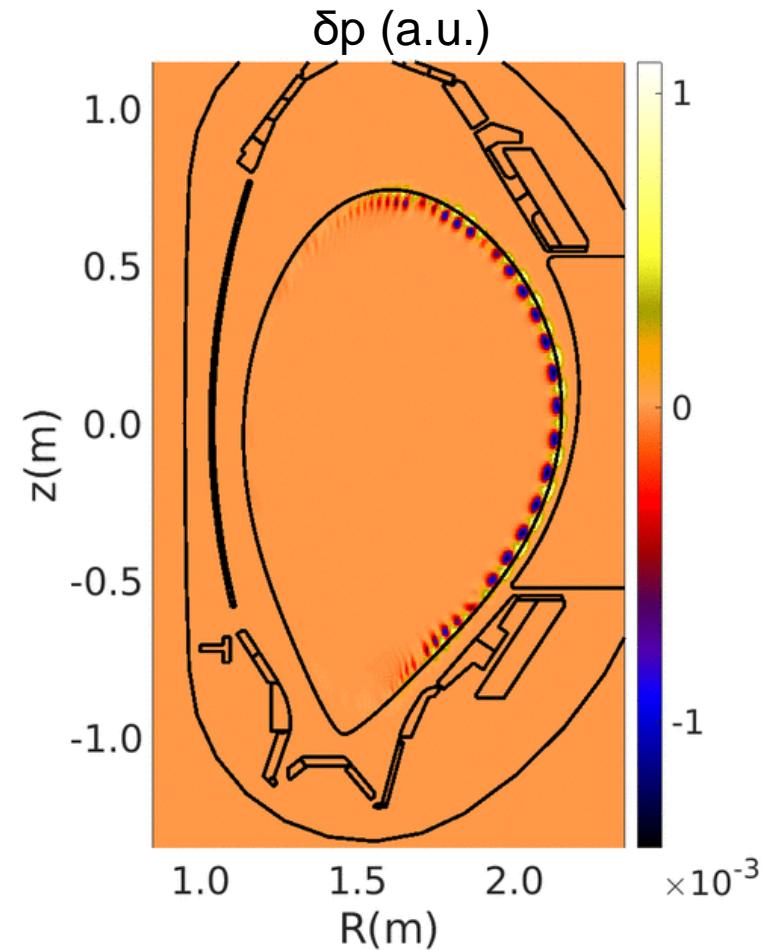
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ELM Crash Successfully Simulated With MEGA



- Modes saturate and filaments extend into SOL

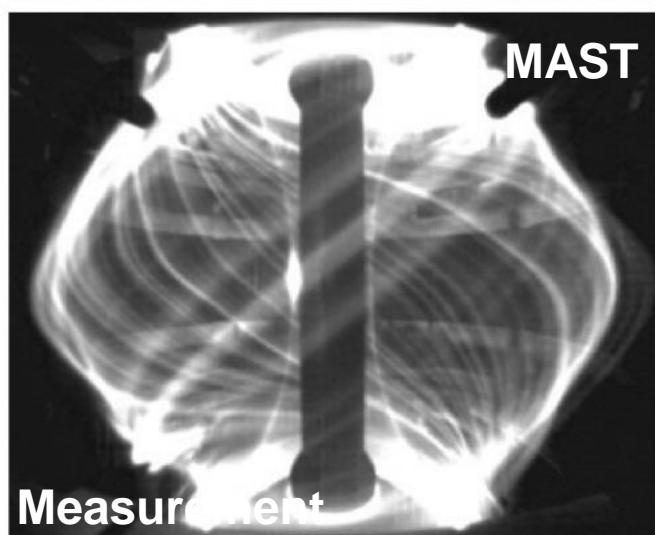
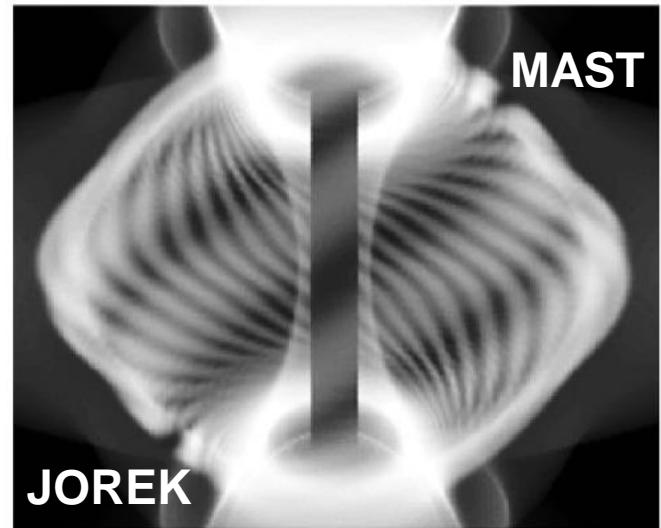




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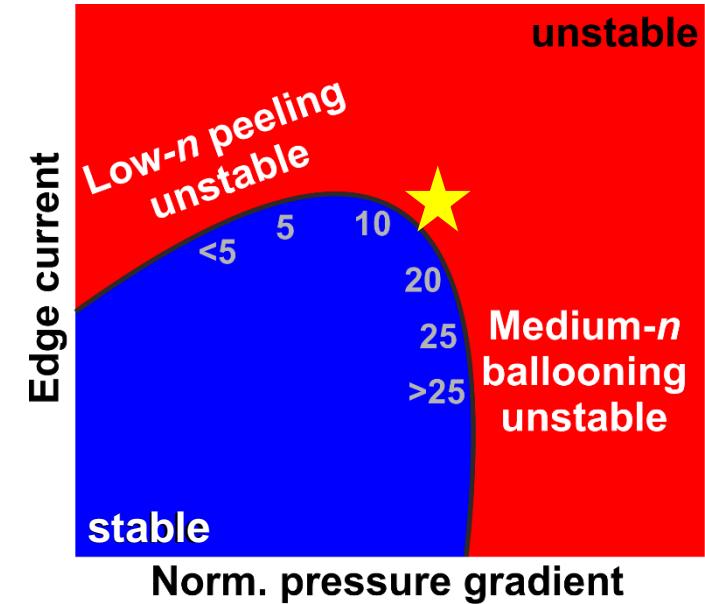
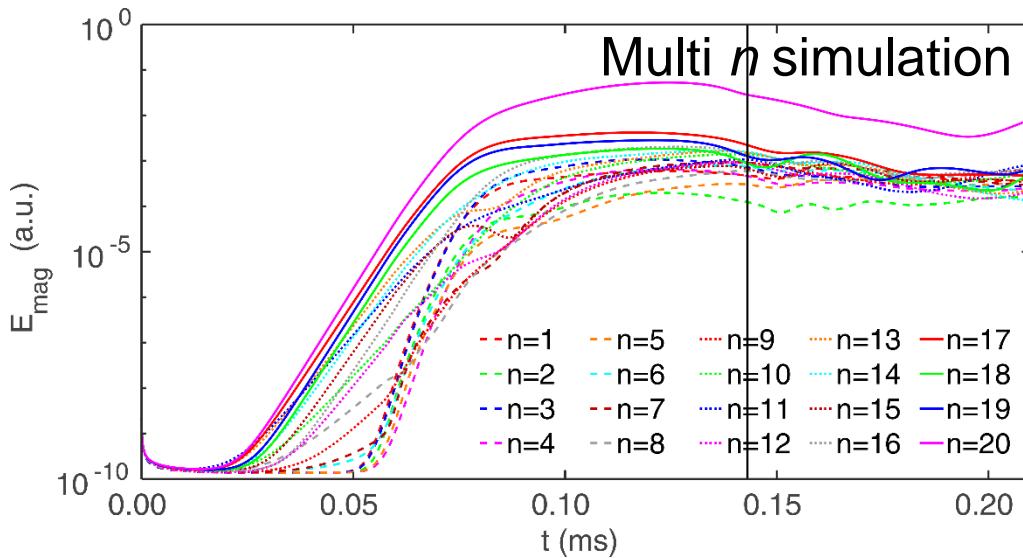
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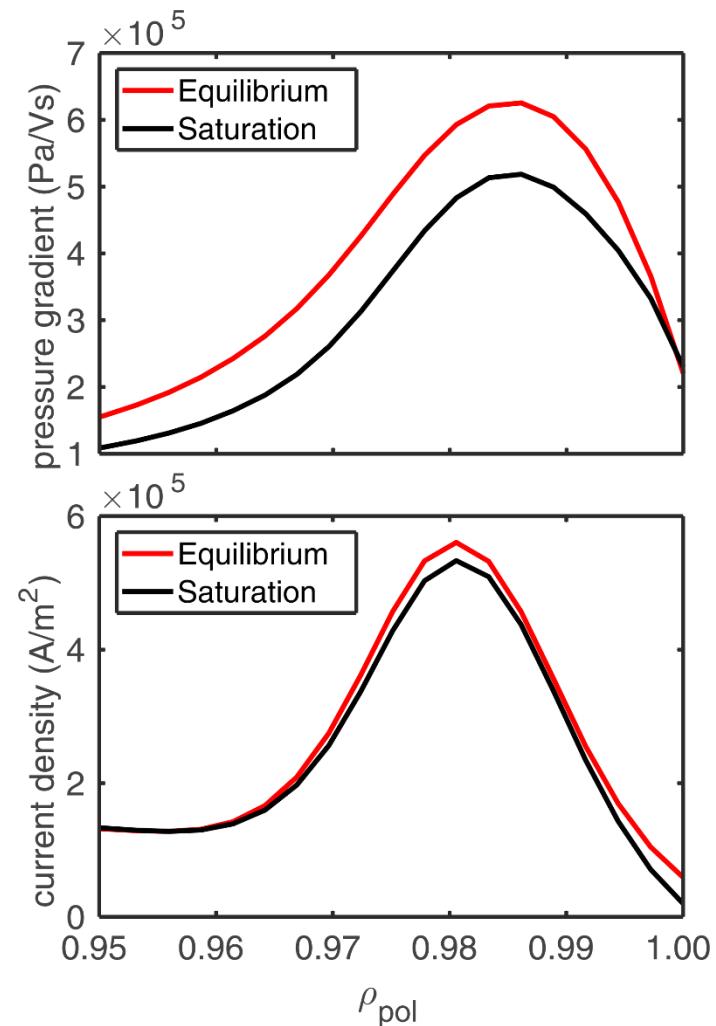
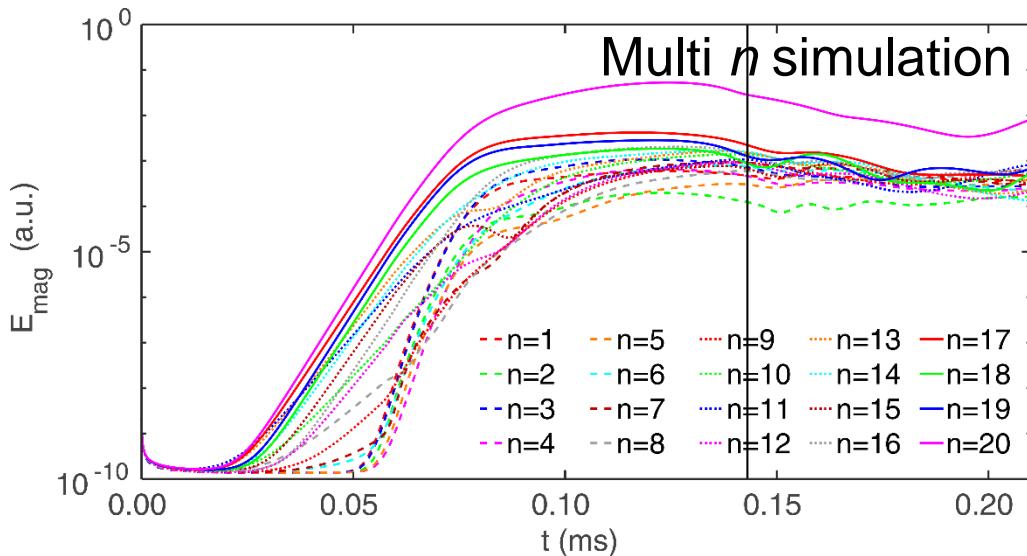
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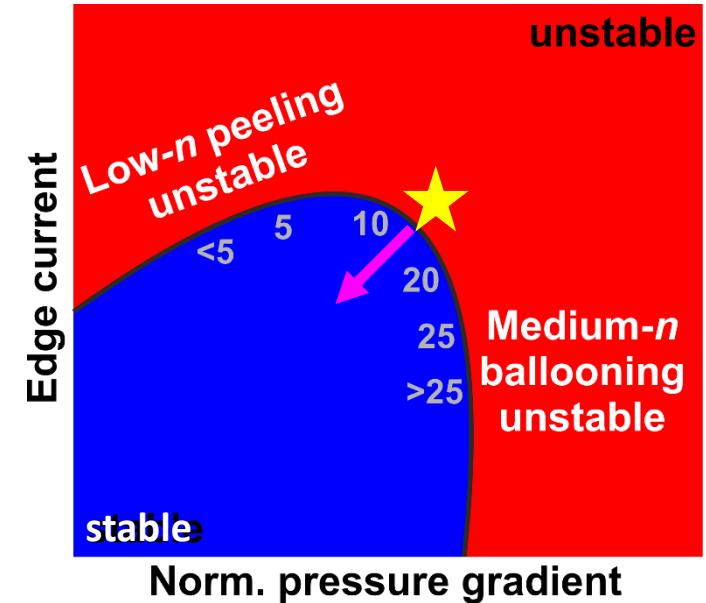
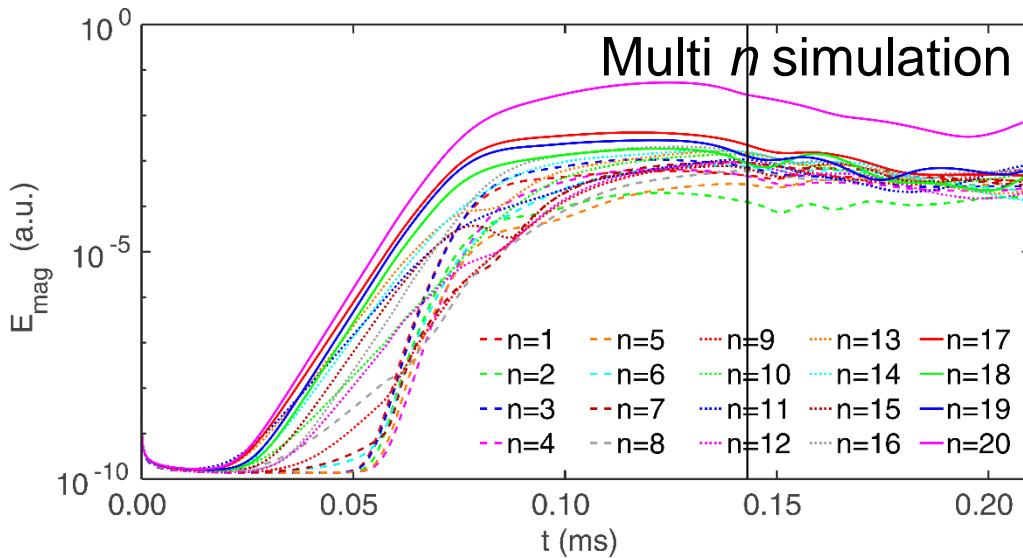
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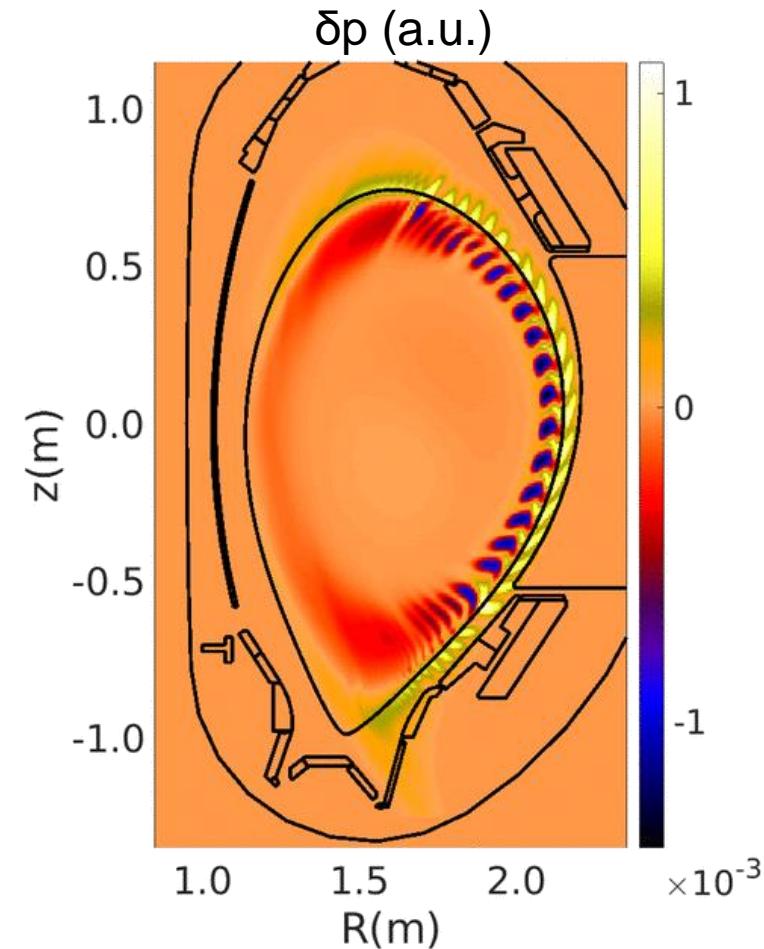
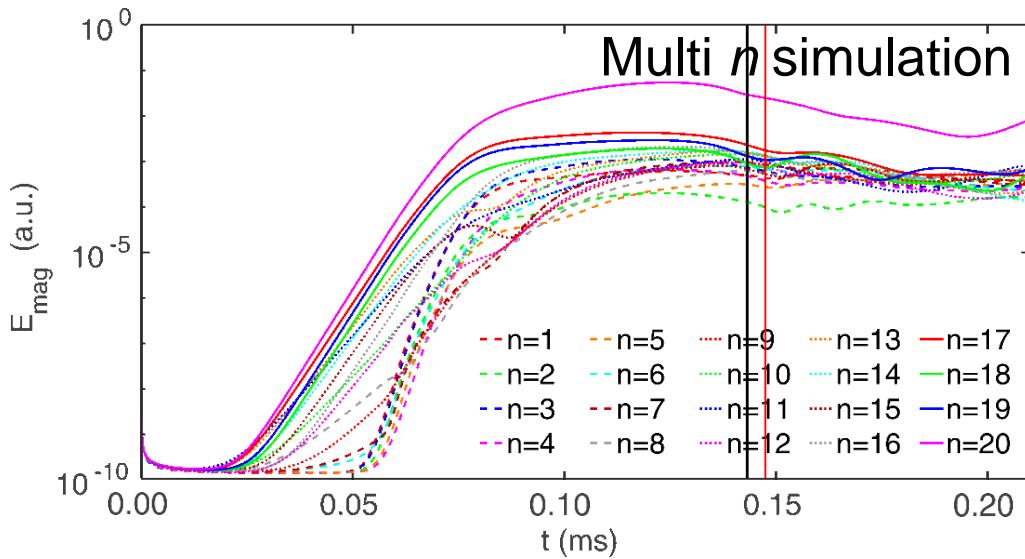
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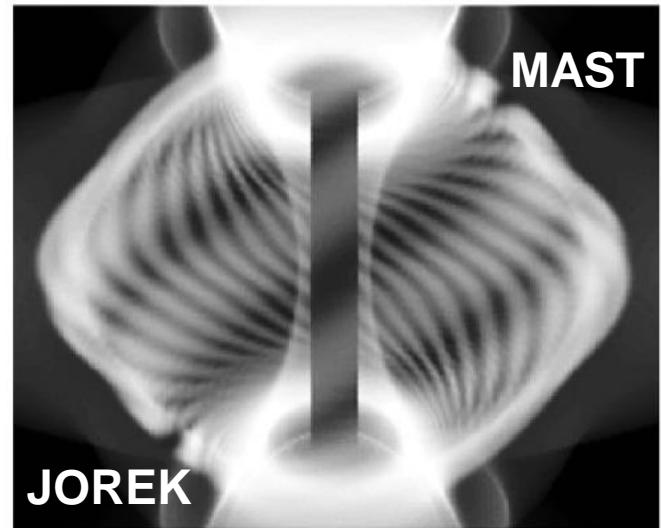
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- Ballooning structure relaxes → ELM signature



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JOREK

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Measurement



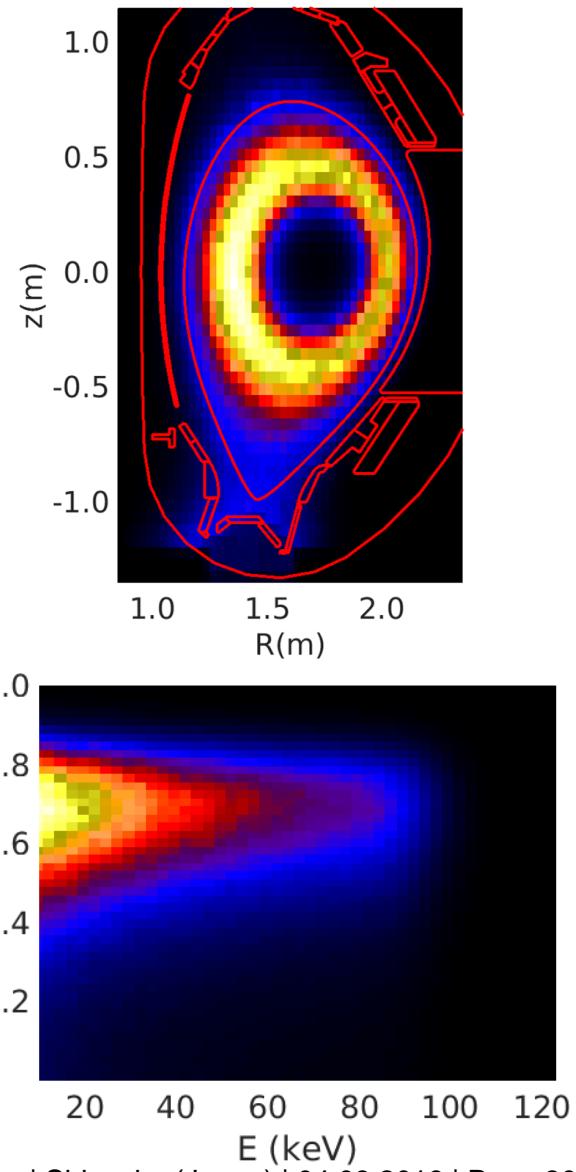
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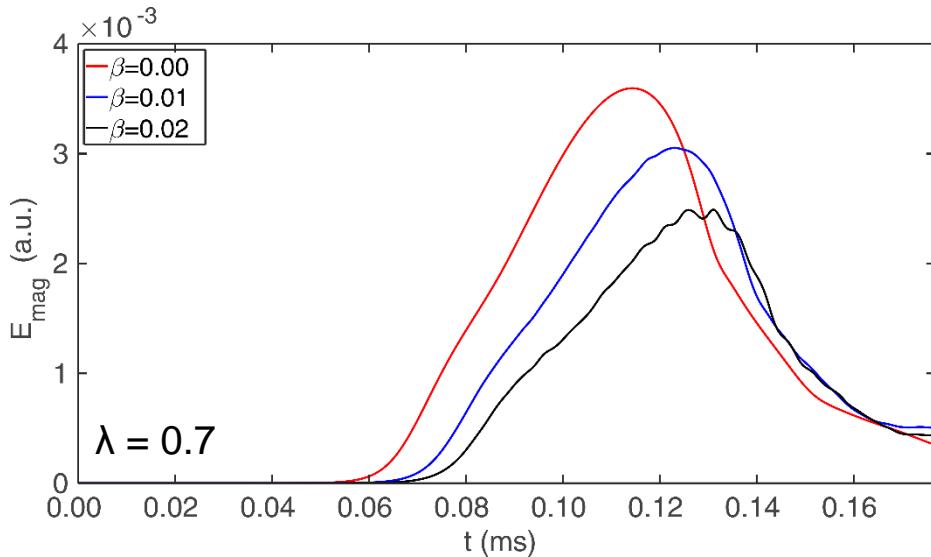
Off-axis Anisotropic Slowing Down Distribution in MEGA



- δf PIC method for gyrokinetic markers
- Collisions and pitch angle scattering not considered
- β_{EP} and pitch angle scan to study impact on $n = 20$

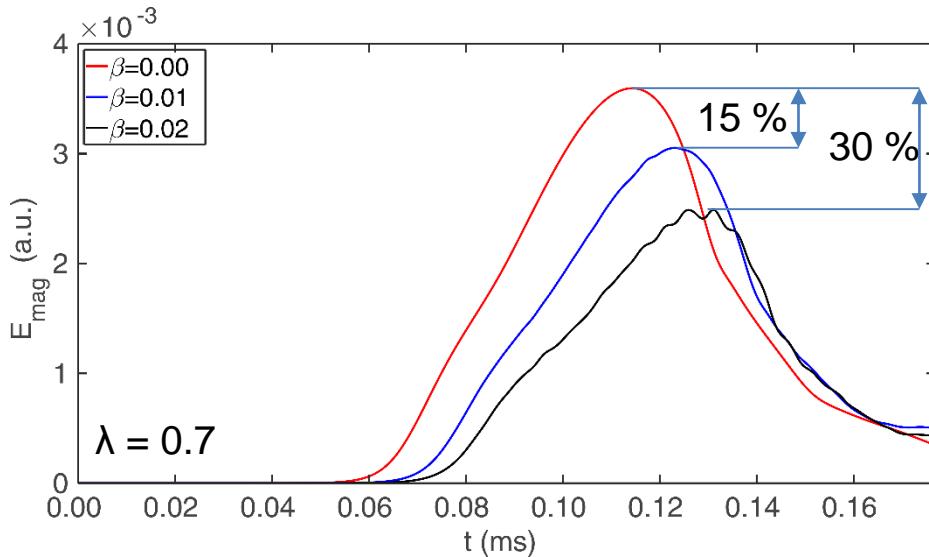


Fast-ions Stabilize $n=20$ Mode in Single- n Simulation



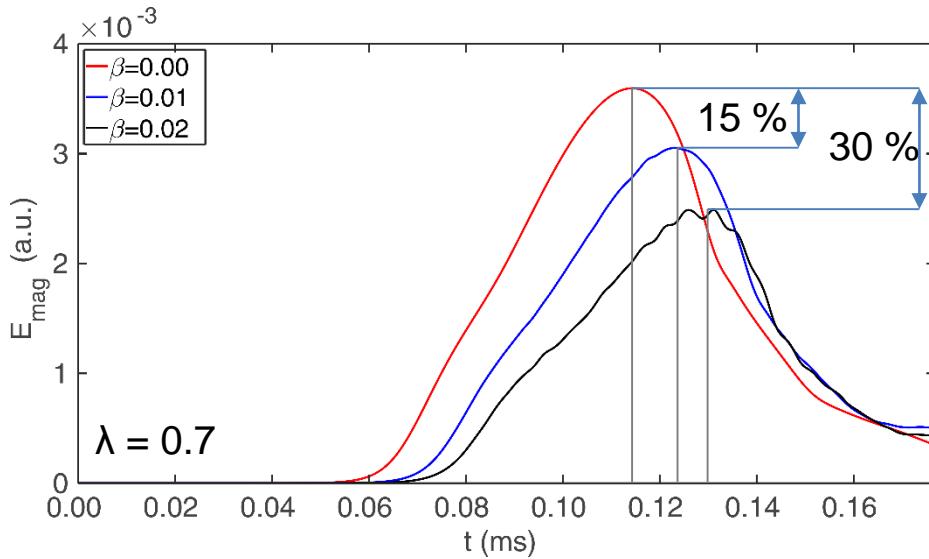
- Mode maximum amplitude depends on β_{EP}

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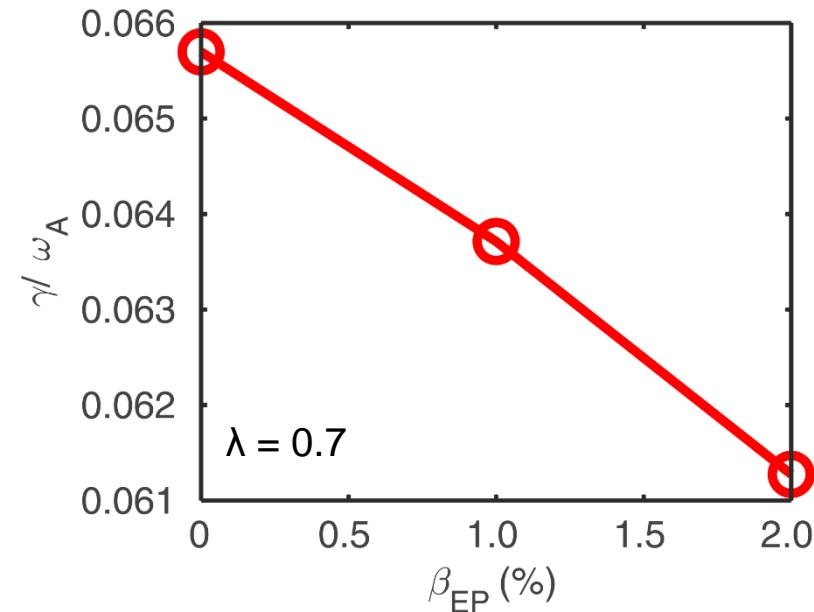
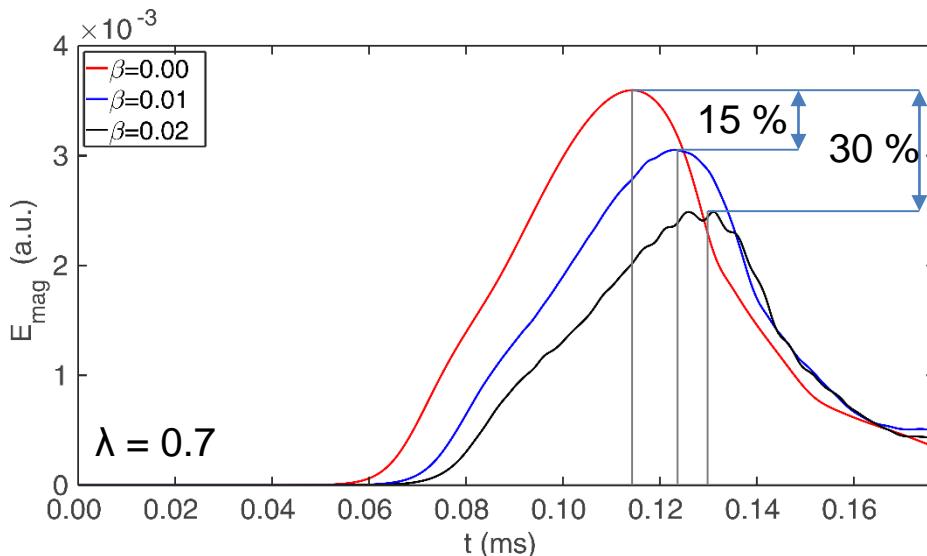
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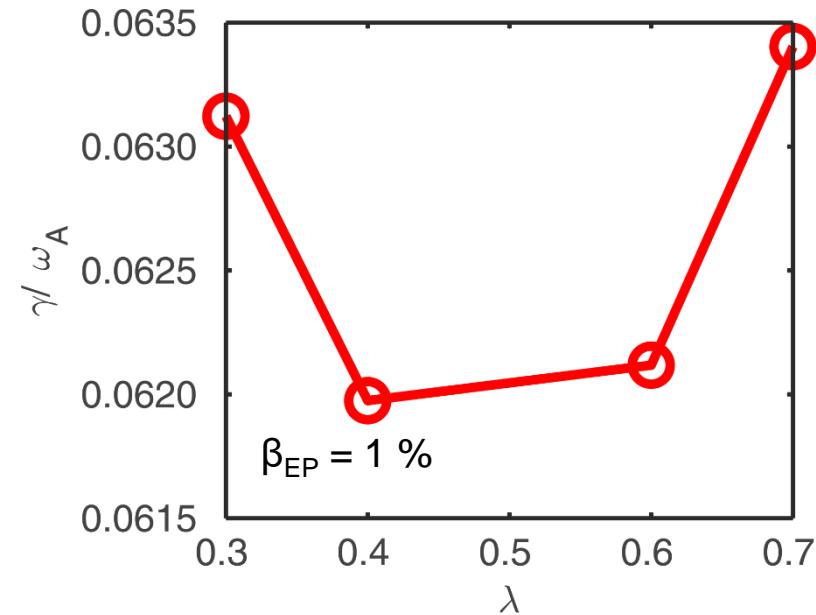
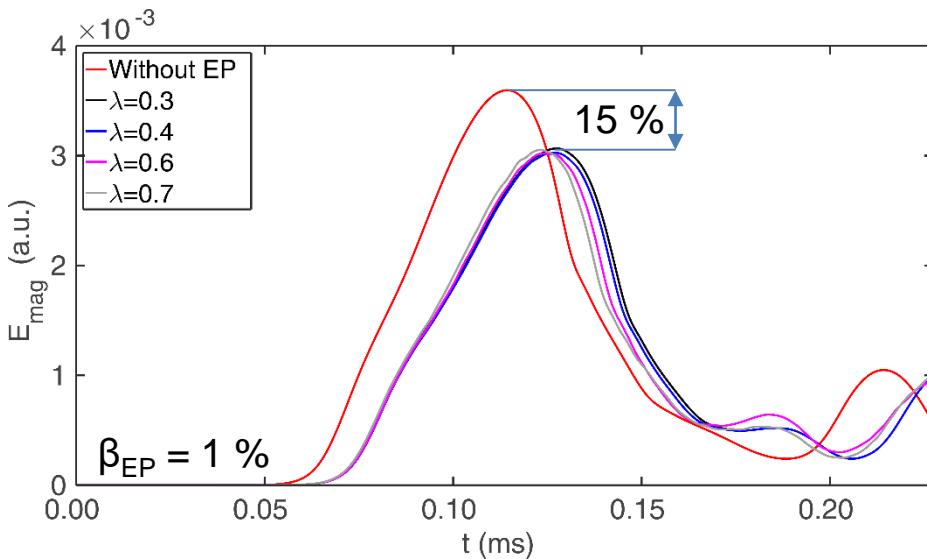
- Mode maximum amplitude depends on β_{EP}
- Saturated maximum amplitude is achieved slightly later

Fast-ions Stabilize $n=20$ Mode in Single- n Simulation



- Mode maximum amplitude depends on β_{EP}
- Saturated maximum amplitude is achieved slightly later
 - Linear growth rate slightly decreases with β_{EP}

Fast-ions Stabilize $n=20$ Mode in Single- n Simulation



- Mode maximum amplitude depends on β_{EP}
- Saturated maximum amplitude is achieved slightly later
 - Linear growth rate slightly decreases with β_{EP}
- For different pitch angles, $n = 20$ saturates at similar energies



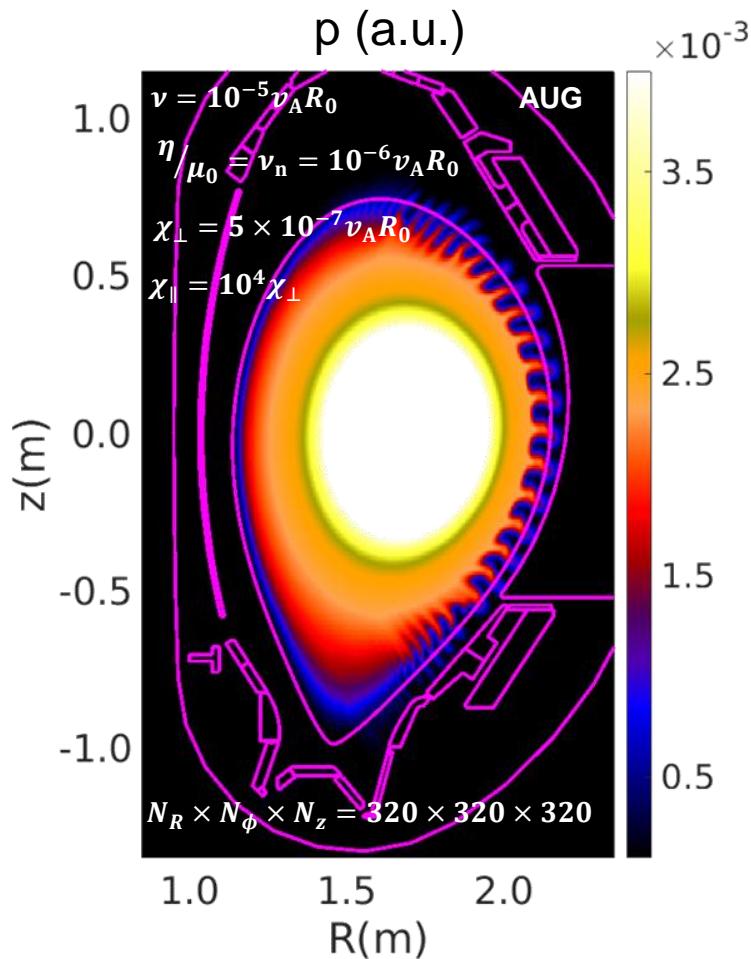
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Summary

- ELMs have been successfully simulated with MEGA
- Ballooning mode growth has been observed
- Filaments extend into SOL and perturbation relaxes
- Fast-ions stabilize high n ballooning modes





Outlook

- Impact of diamagnetic, toroidal and neoclassical flows on ELM stability
- Benchmark results against other codes, such as JOREK^[1], ELITE, GATO
- Study ELM mitigation/suppression by RMPs* including fast-ion kinetic effects
- Study energetic particle transport due to ELMs**

[1] M. Hoelzl *et al.*, Contrib. Plasma Phys. 58, 512-28 (2018)

*See J. Gonzalez-Martin Poster (P 1-4, Wed 13:30)

**See J.F. Rivero-Rodriguez (I-8, Wed 11:00)



Back Up

Non-Linear MHD Equations Solved in MEGA (Standard MHD Model)



$$\frac{\partial \rho}{\partial t} = -\nabla \cdot (\rho \vec{v}) + \nu_n \Delta (\rho - \rho_{\text{eq}})$$

$$\rho \frac{\partial \vec{v}}{\partial t} = -\rho(\vec{v} \cdot \nabla)\vec{v} - \nabla p + (\vec{j} - \vec{j}_h') \times \vec{B} - \nabla \times (\nu \rho \nabla \times \vec{v}) + \frac{4}{3} \nabla(\nu \rho \nabla \cdot \vec{v})$$

$$\begin{aligned} \frac{\partial p}{\partial t} = & -\nabla \cdot (p \vec{v}) - (\gamma - 1)p \nabla \cdot \vec{v} + \nabla \cdot [\chi_{\perp} \nabla_{\perp} (p - p_{\text{eq}}) + \chi_{\parallel} \nabla_{\parallel} (p - p_{\text{eq}})] \\ & + (\gamma - 1) \left[\nu \rho (\nabla \times \vec{v})^2 + \frac{4}{3} \nu \rho (\nabla \cdot \vec{v})^2 + \eta (\vec{j} - \vec{j}_h') \cdot (\vec{j} - \vec{j}_{\text{eq}}) \right] \end{aligned}$$

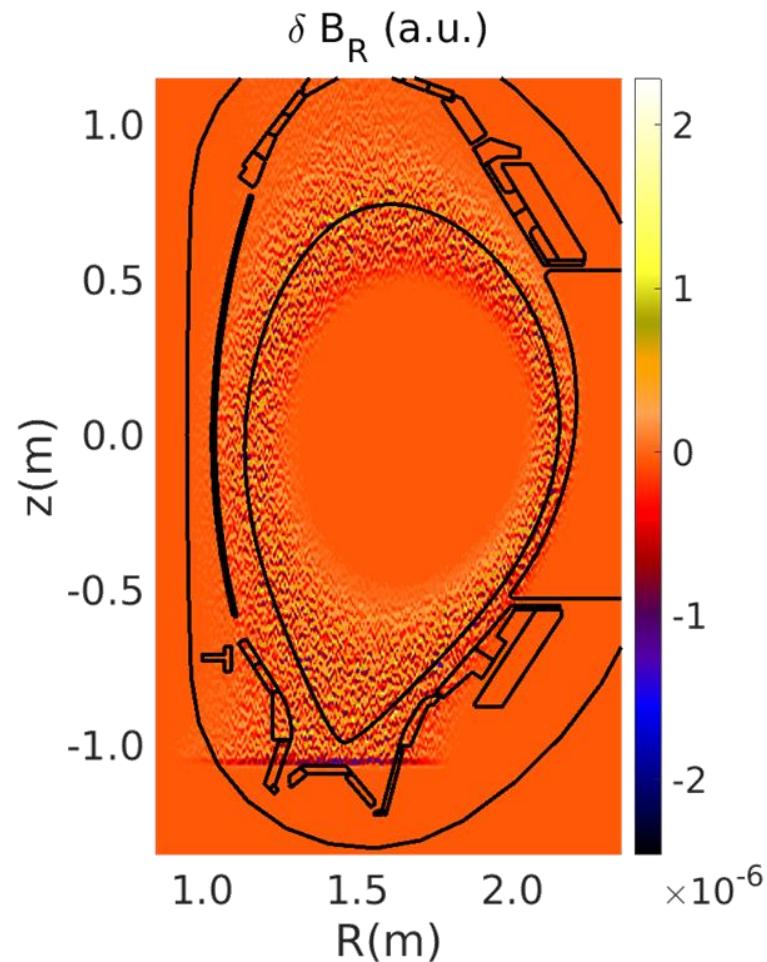
$$\frac{\partial \vec{B}}{\partial t} = -\nabla \times \vec{E}, \vec{j} = \frac{1}{\mu_0} \nabla \times \vec{B}$$

$$\vec{E} = -\vec{v} \times \vec{B} + \eta (\vec{j} - \vec{j}_{\text{eq}})$$

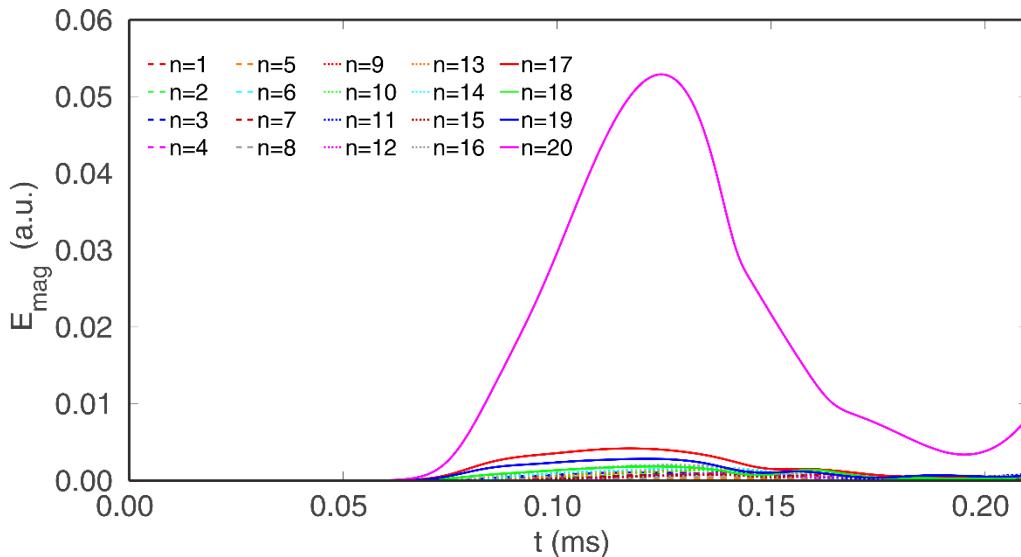
Initializing Simulations in MEGA



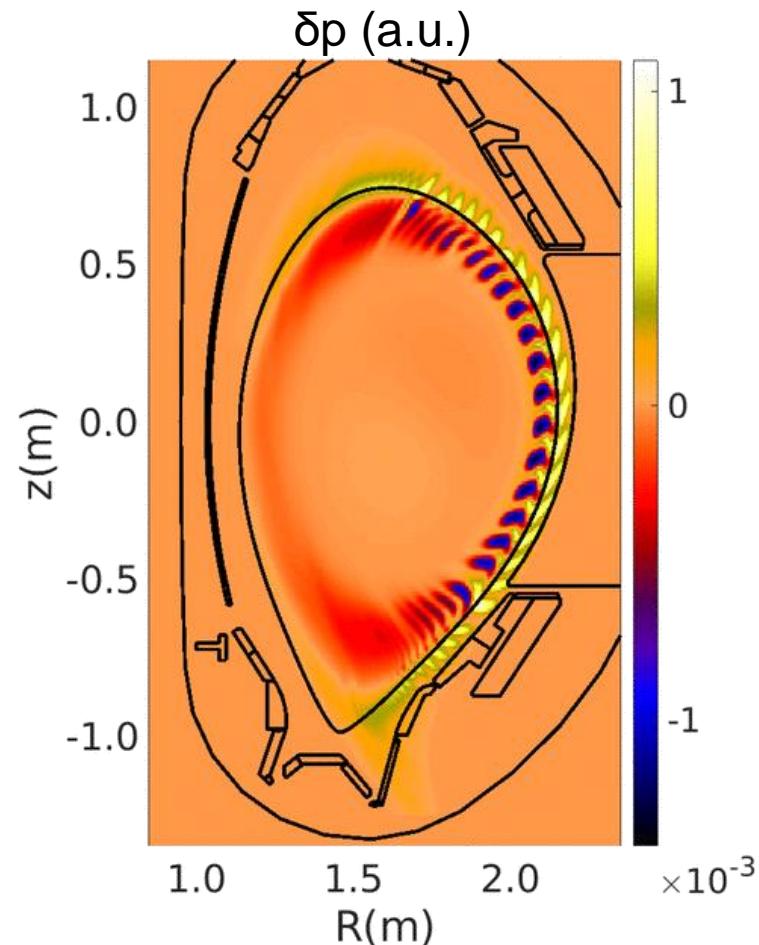
- Initial MHD force balance is calculated before time advances, from which perturbations are calculated
- If fast-ions are included, they are considered in the initial MHD force balance. Then, MHD perturbations are calculated
- Initial weak perturbation is later applied at the edge, $0.90 < \rho_{\text{pol}} < 1.05$
- Serves as seed of instability and helps modes to be excited



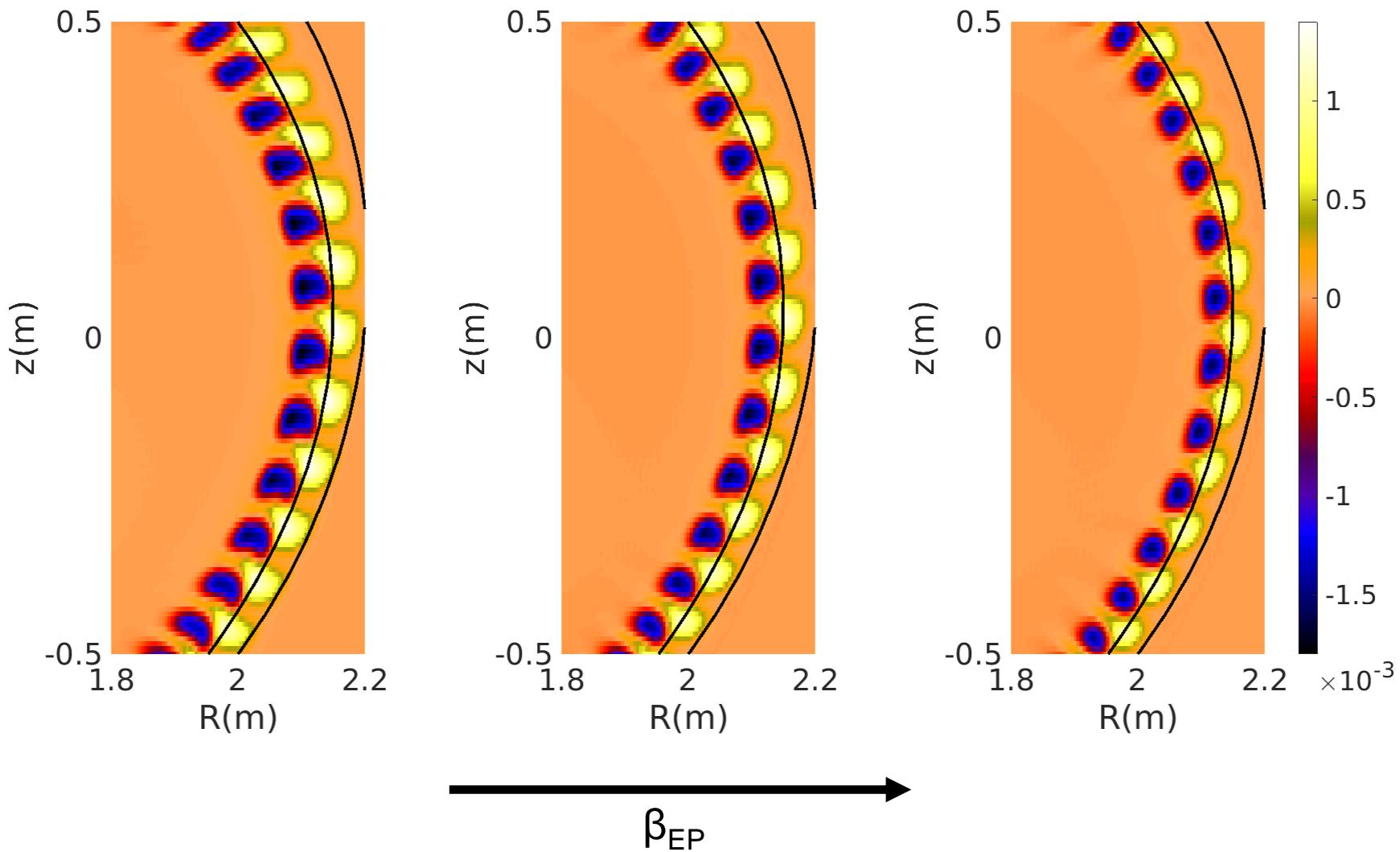
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Poloidal Structure of Perturbation in Hybrid Simulations



Binary Mask Applied in Non-Linear Hybrid Kinetic-MHD Simulations in MEGA



- Binary mask applied in the white coloured region
- Applied to the perturbed density, pressure and velocity evolved by the MHD module of MEGA
- Binary Mask is necessary to avoid numerical instabilities outside the walls

