

driven Alfvén Eigenmodes in ECRH and ECCD Experiments

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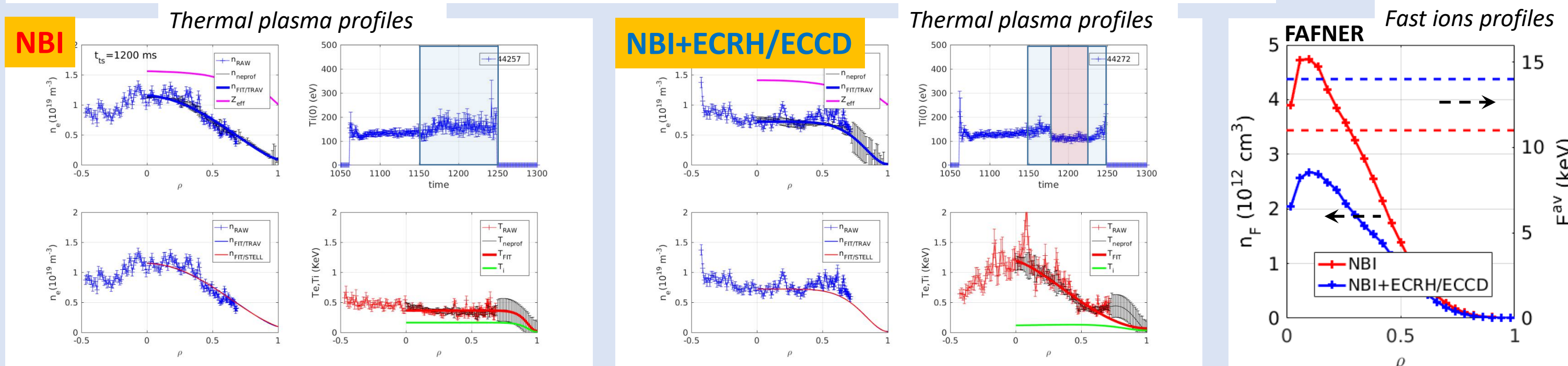
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ABSTRACT

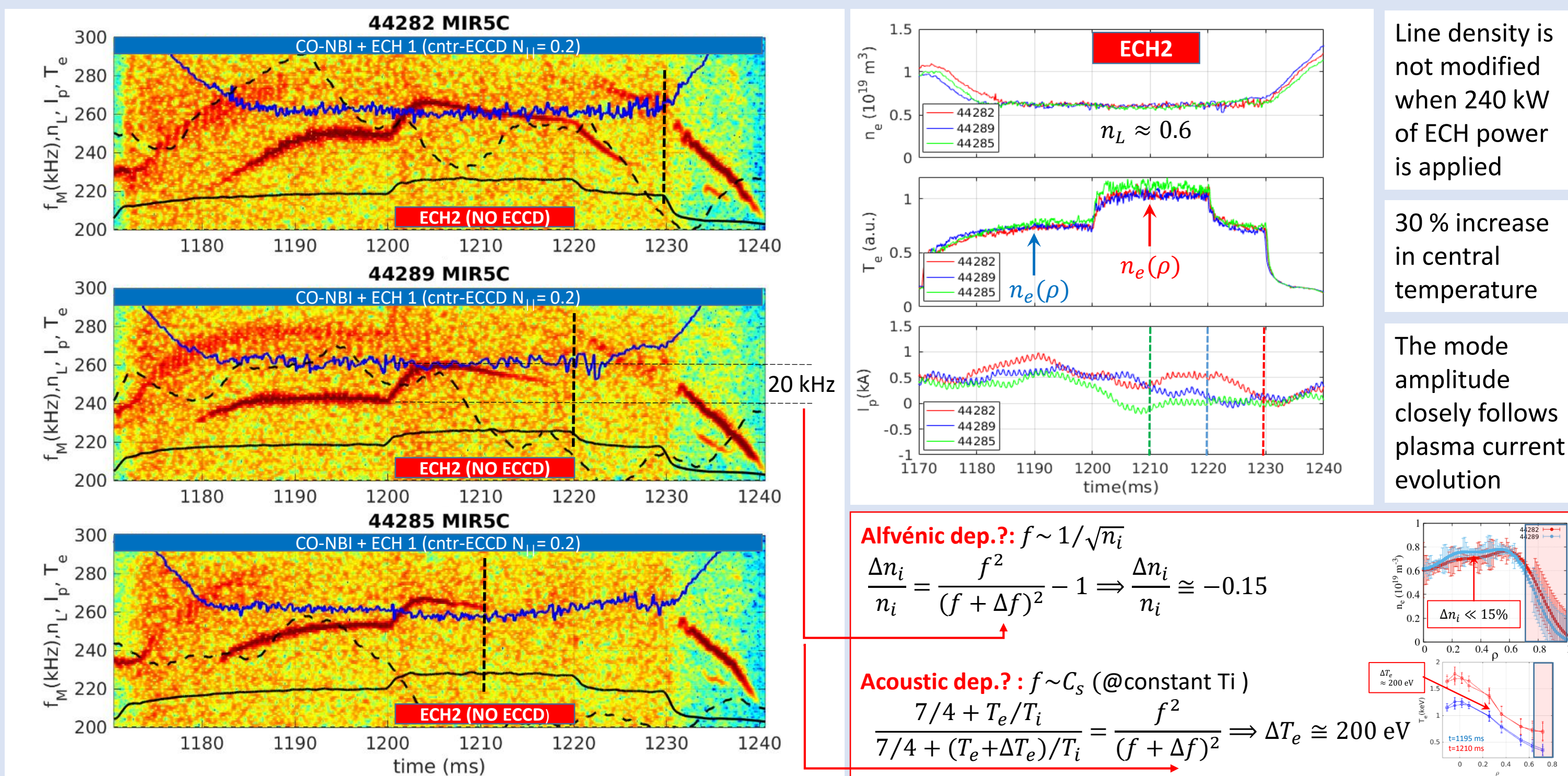
- FAR3D [1] code is used to model the NBI-driven Alfvén Eigenmodes activity in the TJ-II stellarator when ECRH, with and without ECCD, is applied. Preliminary mode identification and growth rates calculation is performed and reasonable agreement is found in most cases.
- Lack of mode number measurements prevents further confirmation and discrimination among different modes with similar frequencies.

BACKGROUND

- ECRH and ECCD are considered as external “actuators” for AEs control [2].
- Experiments in TJ-II show that slight changes in total current have a strong impact on shear Alfvén waves spectrum (SAS) [3].
- $\iota(\rho)$ is calculated with VMEC using theoretical plasma current profiles obtained with FAFNER [4], DKES [5] and TRAVIS [6] \Rightarrow STELLGAP [7] and AE3D [8] results shows changes in SAS.
- No $\iota(\rho)$ or n/m measurement were available in these shots



- Moreover, mode frequency observed in ECCD experiments is modified by adding ECRH power ($n_e(\rho) \sim \text{constant}$, $T_e(\rho) \nearrow$)



Numerical tools

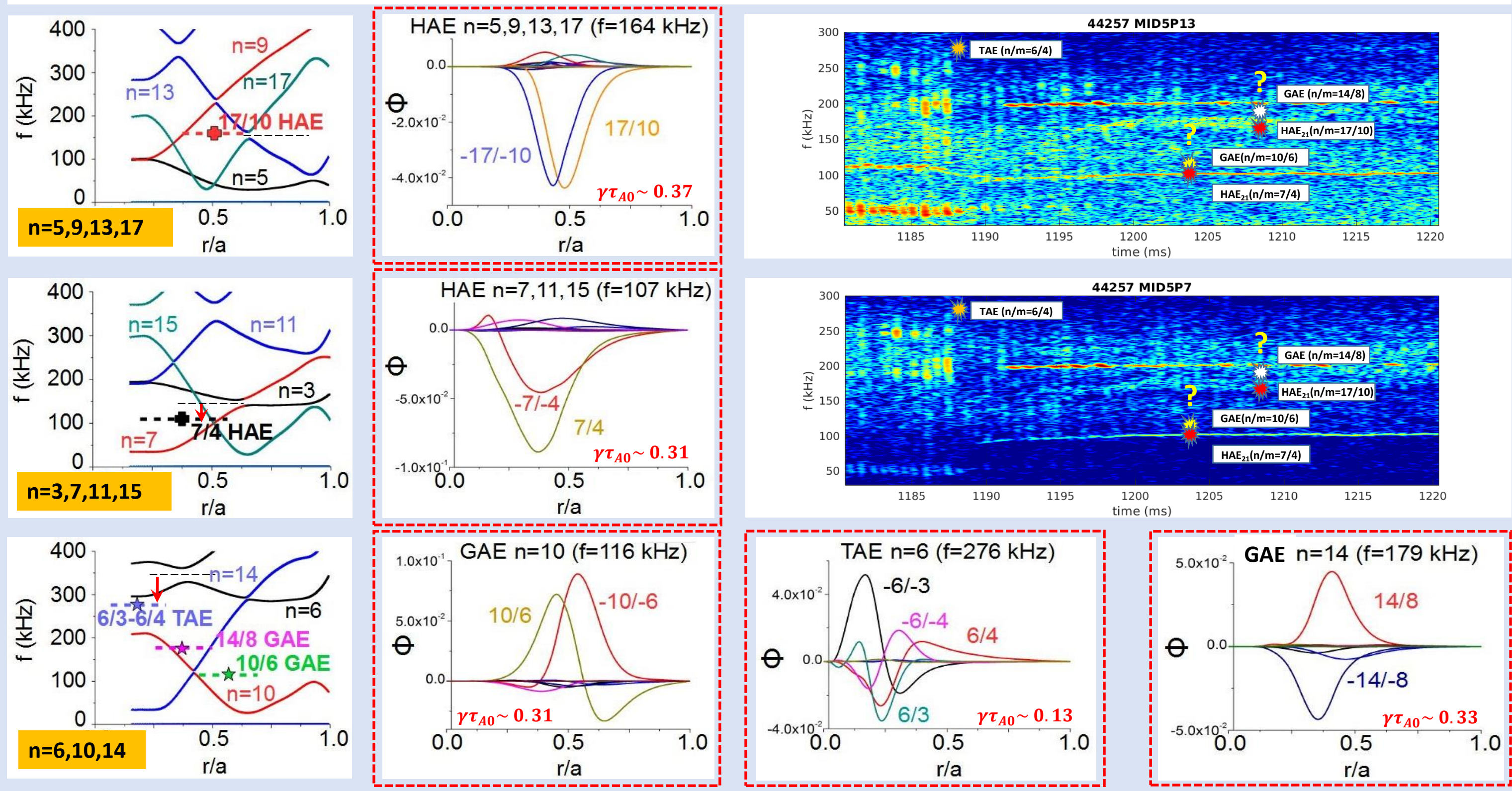
FAR3D gyro fluid simulations [1 and refs. therein]
The code solves the reduced linear resistive MHD equations and the moment equations governing the evolution of the energetic ions density and their parallel velocities, including resonance effects to account for the growth of the MHD perturbations. Stabilizing effects due to electron Landau damping have not been included.

STELLGAP
Calculates 3D shear Alfvén continua

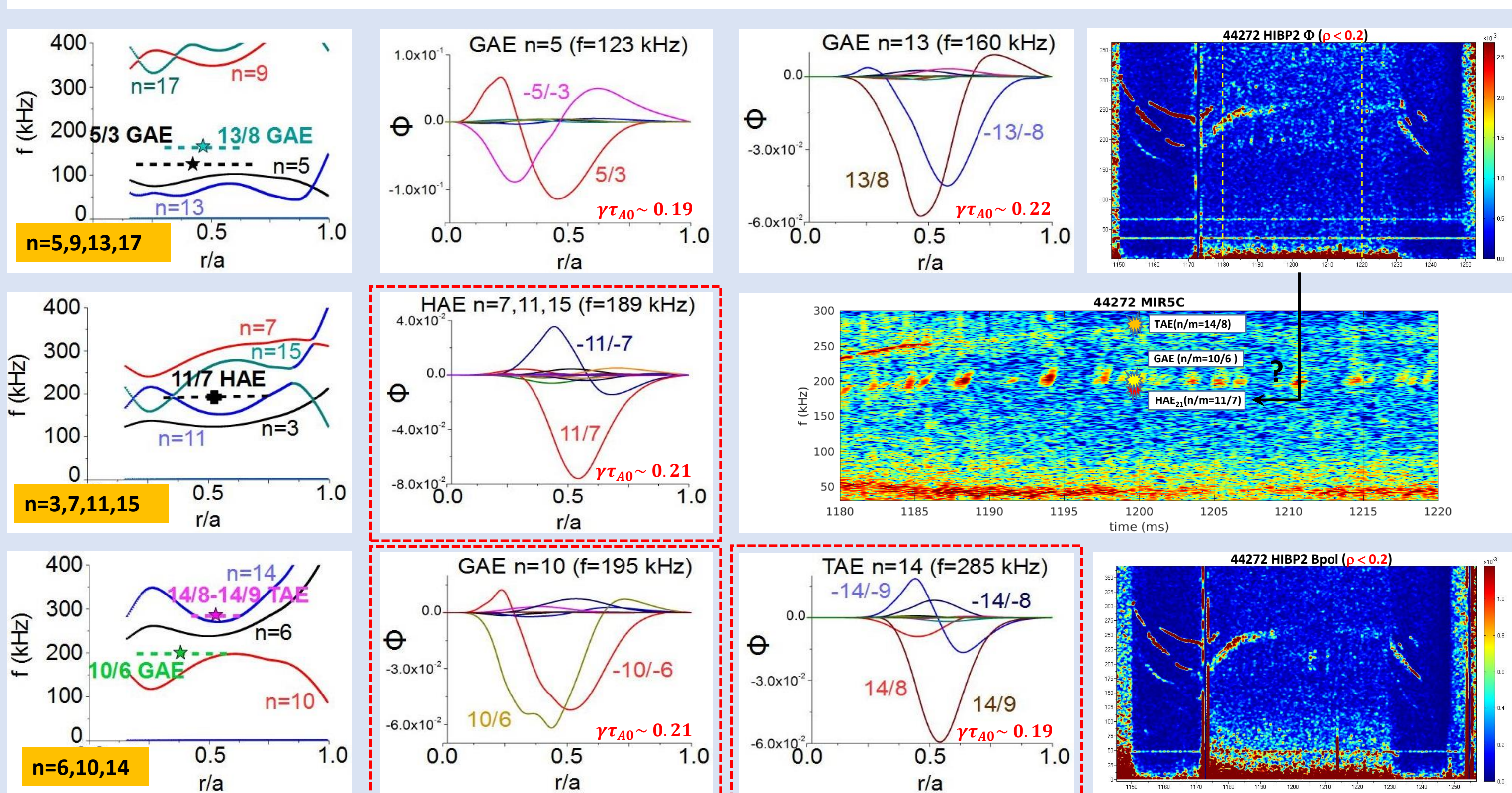
Modes number considered	
toroidal (n)	poloidal (m)
n=[5,9,13,17]	m=[2,13]
n=[6,10,14]	m=[3,11]
n=[7,11,15]	m=[3,11]
n=[0,4,8,12]	m=[0,10]

RESULTS

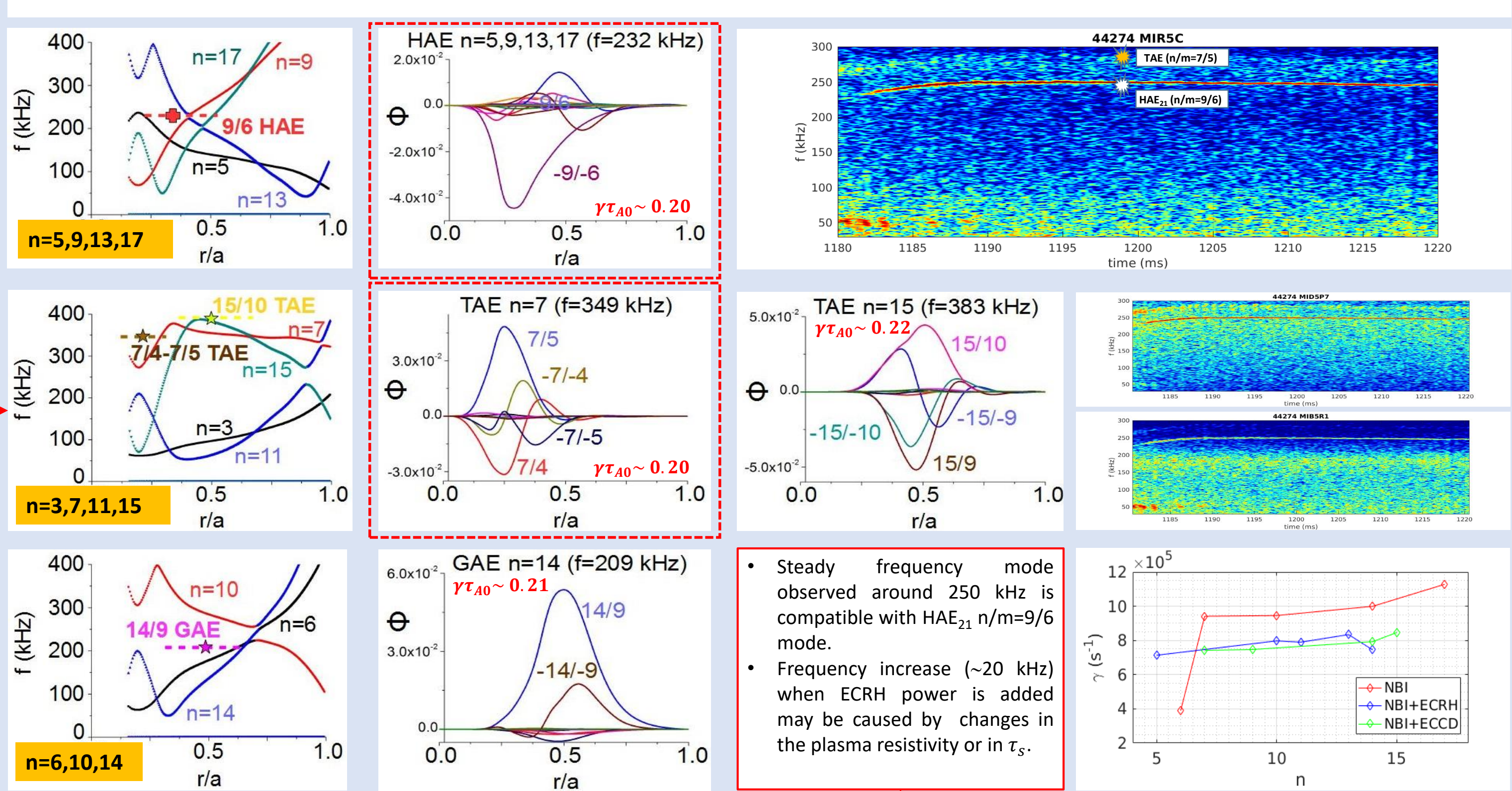
NBI: instabilities compatible with TAE, HAE and GAE $\tau_{A0} = \frac{R_0}{v_{A0}} = R_0 \frac{\sqrt{4\pi\rho m}}{B_0}$ $\tau_{A0}^{NBI} = 0.33 \mu\text{s}$



NBI+ECRH: instabilities compatible with HAE, TAE and GAE $\tau_{A0}^{NBI+ECRH} = 0.26 \mu\text{s}$



NBI+ECCD: instabilities compatible with HAE and TAE $\tau_{A0}^{NBI+ECCD} = 0.26 \mu\text{s}$



CONCLUSION

- Several instabilities with frequencies compatible with the different modes observed in each case are found.
- Modes with the highest observed frequency are consistent with faint TAEs.
- Rest of modes observed at intermediate/high frequency are consistent with HAEs and GAEs.
- Mismatch between gap frequency determined by STELLGAP and FAR3D probably to different implementation of compressibility effects and overestimated electron temperature in STELLGAP calculations.
- Uncertainties related to rotational transform profile or the absence of electron-ion Landau damping processes (not yet included) may explain why several predicted instabilities are not observed in the experiments.

REFERENCES
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