

# **Overview of ASDEX Upgrade results**



**3-strap ICRF antenna pair** 

reduces W production



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# Introduction

The ASDEX Upgrade (AUG) programme is directed towards physics input to critical elements of the ITER design and the preparation of ITER operation, as well as addressing physics issues for a future DEMO design. This overview summarizes the progress in the last 2 years on the machine hardware, non-Inductive operation, ELM suppression with magnetic perturbations, pedestal stability and confinement, power exhaust and scenario integration. Challenges caused by the presence of tungsten PFCs are addressed and trends for extrapolation to large devices are given.

# **Scenario development**

**Non-inductive operation** 

# **Tungsten related hardware upgrades**

#### Massive tungsten divertor III performs without failure

#### **Cracks formed – through tile and at the surface**

- cracks not observed in high heat flux tests in GLADIS before
- forces during disruption in addition to thermo-mechanical stresses







[2]

• about 40 % of I<sub>p</sub> driven by NBCD, 50 % boostrap, 10 % ECCD • ECCD used to tailor current profile for optimum stability and  $q_{min} > 1.5$ 



#### **ELM suppression with RMPs**

• full ELM suppression obtained in AUG. trick:  $\delta$ -dependence of threshold as found in DIII-D (minor radius  $a \times 1.19$ ) no accumulation of W at pedestal top



quite high W conc. due to hot divertor

[3]

FEM calculations suggest vertical tile splitting for crack avoidance (done in current AUG vent) <sup>1</sup>/<sub>4</sub> of tiles will be made from more ductile material 97 % W, 2 % Ni, 1 % Fe (magnetic)

## **Power exhaust**

#### **Detachment studies with X-point radiator**

- X-point radiator at pronunced detachment with N and Ar seeding
- favourable properties: cold divertor with low power load, no W sputtering, smaller ELMs
- expected to be feedback-controllable



#### **Favourable tungsten behaviour expected** in a large device

neoclassical W transport for pedestal scaling n<sub>W,ped</sub>/n<sub>W,sep</sub>

#### Integration of Tdiv feedback control and pellet density control

• Fb control of heat flux (N) and density (pellets) are compatible



High nitrogen divertor enrichment achieved **ELM frequency is important for enrichment, but not exclusive** 

impurity divertor	enrichment	F -
······	]	$E = \overline{Z}$



• N seeding effects fueling via reduction of high field side high density

27100 t = 2.2 s



#### other effects:

impurity ionization length neoclassical pedestal inward drift

higher E for N compared to Ne seen in modelling [8]

# Summary

• new 3-strap ICRF antenna reduces W release as predicted

- massive W divertor performs well despite cracking
- overcoming challenges of W PFC environment
  - full ELM suppression achieved à là DIII-D
  - non-inductive operation up to Ip= 0.8 MA
- simultaneous heat flux (N) and density control (pellets) - high impurity enrichment obtained, essential for clean core
- relative position of pedestal density and temperature profiles
  - important for edge stability
  - reduction of HFSHD responsible for improved confinement with nitrogen
- good prospects for W in large devices with hot pedestal

#### **References**

200

300

f<sub>ELM</sub> / Hz

400

[1] A. Herrmann, NF 55 (2015), PSI 2016

# **Pedestal stability and confinement**

### **Pedestal- and global energy confinement depend** on edge density profile





on pedestal top pressure





leading to inward shift of density profile [5, 8]

**Effect of N seeding via HFSHD** 

as heating power reduction

### Magnetic fluctuations @ few 100 kHz in later phase of ELM cycle measured both at HFS and LFS side

fluctuations are associated with modes responsible for pressure

 pedestal top pressure closely linked to density at separatrix • further effects, like amplification of  $\beta$  increase via shear

Stability calculations confirm effect of n<sub>e</sub> profile location



profile clamping (KBM in EPEd model) ExB velocity in steep gradient region, ELM – synchronized estimated as compensating diamag. drift

magnetic fluctuations  $\partial B_{r}/\partial t$ 

see [5]



1.8

[2] J.M. Noterdaeme, EX/P6-26 this conference [3] J. Stober, Post Deadline, this conference [4] M. Willensdorfer, this conference [5] M. Dunne, this conference, EX/3-5, Wed [6] F. Laggner, PPCF 58 (2016) 06505 [7] R. Dux, PSI 2016, subm. Nucl. Mat. and Energy [8] F. Reimold, EXS/P6-191, Thu [9] H. Meyer, MST1 Overview, OV/P-12

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