





## Synergetic effects of collisions, turbulence and sawtooth crashes on impurity transport.

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### **Motivation : impurity transport**

- Tungsten plasma facing components → impurity accumulation in the core ?
- Neoclassical and turbulence transport processes compete Casson 13, Angioni 14
- Interplay with MHD events : tearing modes, ELMs, sawtooth crashes Hender 16, Sertoli 15



Asdex Upgrade – tungsten density- Sertoli 15







- 1) Interaction between turbulent and neoclassical transport.
- 2) Interaction between sawtooth cycles and neoclassical transport.

Punchline : interplay between different contributions to impurity transport are mediated by large scale flows



- Gyrokinetic description (GYSELA code) : d<sub>t</sub>F=C(F) + Poisson equation→ neoclassical and turbulent transport Grandgirard 16
- MHD equations (XTOR code) + impurity density and momentum equations Lütjens 10

$$\frac{\partial N}{\partial t} + \nabla \cdot (N\mathbf{V}) = \nabla \cdot (\mathcal{D}\nabla N - \mathcal{V}N) \qquad \begin{array}{l} \text{collisional friction} \\ \text{force} \rightarrow \\ \text{neoclassical flux} \end{array}$$

$$Nm\left(\frac{\partial}{\partial t} + \mathbf{V} \cdot \nabla\right) \mathbf{V} = Ne\left(\mathbf{E} + \mathbf{V} \times \mathbf{B}\right) - \nabla \cdot \mathbf{\Pi} + \mathbf{R}$$

 $\rightarrow$  Pfirsch-Schlüter transport included in the fluid dynamics  $\rightarrow$  neoclassical and MHD transport







- Pfirsch-Schlüter convection cell due to perpendicular compressibility Hinton & Hazeltine 76
- controls  $\Gamma^{\Psi}$  at high collisionality

 $v_{7}^{*}>1$ 



R



General form of the impurity flux Hirshman & Sigmar 81



- Impurity collisional, ions weakly collisional  $\rightarrow$  H = -1/2 Hirshman 76
- GYSELA benchmarked against theory and NEO code Belli 08







- Neoclassical and turbulent contributions isolated by playing with collisionality and symmetries
- Total flux ≠ neoclassical + turbulent



CCCC

#### Interplay is mediated by poloidal convective cells



- Turbulent Reynolds stress → poloidal convective cells
- Poloidal asymmetries  $\rightarrow$  change neoclassical impurity flux
- // momentum transport, turbulence self-regulation Diamond 05







- Anti-correlation due to poloidal convective cells
- Thermal screening factor H>-1/2 : consequence of static density poloidal asymmetries? Romanelli 98, Fülöp 99, Angioni 14, Breton 16





- $\nabla N_i = 0, \ \nabla T_i \neq 0 \rightarrow \text{screening}$
- Crash time << collision time → neoclassical transport processes inefficient during crash
- Post-crash profile
   consistent with Kadomtsev
   model Kadomtsev 75, Porcelli 96,
   Nicolas 15









# ExB drift is the main cause of impurity transport during a sawtooth crash





Impurity density and stream function

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#### **ExB drift is the main cause of impurity transport during a sawtooth crash (cont.)**





- ExB impurity flux ~10 flux due to magnetic flutter
- Consistent with SXR measurements on TFTR Nagayama 91



- Neoclassical transport dominant during recovery phase, but ion temperature gradient is lower
- → weaker thermal screening effect
- Overall temperature profile flatter with sawteeth

1.02 Initial profile <N<sub>z</sub>> (XTOR unit) 96'0 86'0 after 9 sawtooth crashes no sawteeth 0.94 0.92<sup>L</sup> 0.2 0.4 0.6 0.8 Normalized minor radius

Impurity density profile - Ahn 16





- Interplay between turbulent and neoclassical transport processes:
- poloidal convective cells generated by turbulence → poloidal asymmetries – total flux ≠ turbulent + neoclassical calculated separately. Should play at low rotation speed (e.g. EAST, WEST, ITER)
- thermal screening gets weaker
- Sawteeth cycles affect neoclassical transport
- crashes flatten impurity density profile + lower main ion temperature gradient → thermal screening less efficient