Penetration and amplification of resonant perturbations in 3D ideal-MHD equilibria

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- 1. Macroscopic force-balance, $\nabla p = \mathbf{j} \times \mathbf{B}$, is fundamental for magnetic confinement. defines the equilibrium state, linear stability, neoclassical transport, experimental design & reconstruction
- 2. Conventional solutions with rational rotational-transform flux-surfaces are non-analytic and non-physical.
- 3. We compute <u>physically acceptable</u>, <u>mathematically self-consistent</u> solutions to $\nabla p = \mathbf{j} \times \mathbf{B}$ with nested flux-surfaces in arbitrary (three-dimensional) geometry.
- 4. Our new class of solutions have <u>sheet-currents</u> and <u>discontinuous rotational-transform</u>.
- 5. In our solutions, RMPs penetrate into the core, and are amplified by pressure.

1. Consider e.g. *dynamical*, resistive, single-fluid equations with flow.

0

$$\frac{\partial p}{\partial t} = (\gamma - 1) \left(Q + \nabla \cdot (\kappa_{\parallel} \nabla_{\parallel} p + \kappa_{\perp} \nabla_{\perp} p) + \eta j^{2} - \Pi : \nabla \mathbf{v} \right) - (\mathbf{v} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{v})$$

$$\frac{\partial n}{\partial t} = -\nabla \cdot (n \mathbf{v})$$

$$\frac{\partial \mathbf{v}}{\partial t} = \frac{1}{nm_{i}} \left(\mathbf{j} \times \mathbf{B} - \nabla p - \nabla \cdot \mathbf{\Pi} \right) - \mathbf{v} \cdot \nabla \mathbf{v}$$

$$\frac{\partial \mathbf{B}}{\partial t} = \nabla \times (\eta \nabla \times \mathbf{B} - \mathbf{v} \times \mathbf{B})$$

2. Numerically too complicated, too slow for stellarator design optimization.

(i) let $\partial_t = 0$, (ii) ignore non-ideal terms, η , κ_{\perp} , Π , (iii) ignore velocity, v_{τ} , (iv) preserves p, to obtain $\nabla p = j \times B$, ..., calculation is still too slow ..., (v) restrict attention to integrable magnetic fields.

- The most widely used equilibrium code for stellarators, VMBC, and the IPEC code for perturbed tokamaks solve $\nabla p = j \times B$ with nested flux surfaces [.
- Foundation of equilibrium, linear stability, and neo-classical transport.

1. Consider resistive, single-fluid equations with flow.

$$0 = (\gamma - 1) \left(Q + \nabla \cdot (\kappa_{\parallel} \nabla_{\parallel} p + \kappa_{\perp} \nabla_{\perp} p) + \eta j^{2} - \mathbf{\Pi} : \nabla \mathbf{v} \right) - (\mathbf{v} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{v})$$

$$0 = -\nabla \cdot (n\mathbf{v})$$

$$0 = \frac{1}{nm_{i}} \left(\mathbf{j} \times \mathbf{B} - \nabla p - \nabla \cdot \mathbf{\Pi} \right) - \mathbf{v} \cdot \nabla \mathbf{v}$$

$$0 = \nabla \times (\eta \nabla \times \mathbf{B} - \mathbf{v} \times \mathbf{B})$$

2. Numerically too complicated, too slow for stellarator design optimization.

(i) let $\partial_t = 0$,

- The most widely used equilibrium code for stellarators, VMEC, and the IPEC code for perturbed tokamaks solve $\nabla p = j \times B$ with rested flux surfaces.
- I. Foundation of equilibrium, linear stability, and neo classical transport.

1. Consider single-fluid equations with flow.

$$0 = (\gamma - 1) \left(Q + \nabla \cdot (\kappa_{\parallel} \nabla_{\parallel} p \qquad) \qquad \right) - (\mathbf{v} \cdot \nabla p + \gamma p \nabla \cdot \mathbf{v})$$

$$0 = -\nabla \cdot (n \mathbf{v})$$

$$0 = \frac{1}{n m_{i}} \left(\mathbf{j} \times \mathbf{B} - \nabla p \qquad \right) - \mathbf{v} \cdot \nabla \mathbf{v}$$

$$0 = \nabla \times (\qquad -\mathbf{v} \times \mathbf{B})$$

2. Numerically too complicated, too slow for stellarator design optimization.

(i) let $\partial_t = 0$, (ii) ignore non-ideal terms, $\eta, \kappa_{\perp}, \mathbf{\Pi}$,

The most widely used equilibrium code for stellarators, VMBC, and the IPBC code for perturbed tokamaks solve $\nabla p = j \times B$ with nested flux surfaces [.

Foundation of equilibrium, linear stability, and neo-classical transport.

1. Consider single-fluid equations

$$0 = (\gamma - 1) \left(Q + \nabla \cdot (\kappa_{\parallel} \nabla_{\parallel} p) \right)$$
$$0 = \frac{1}{nm_{i}} \left(\mathbf{j} \times \mathbf{B} - \nabla p \right)$$

2. Numerically too complicated, too slow for stellarator design optimization.

(i) let $\partial_t = 0$, (ii) ignore non-ideal terms, $\eta, \kappa_{\perp}, \mathbf{\Pi}$, (iii) ignore velocity, \mathbf{v} ,

iv) prescribe p, to obtain $\nabla p = j \times B$, calculation is still too slow.

(v) restrict attention to integrable magnetic fields.

- The most widely used equilibrium code for stellarators, VMEO, and the IPEO code for perturbed tokamaks solve $\forall p = j \times B$ with rested flux surfaces.
- 4. Foundation of equilibrium, linear stability, and neo classical transport.

1. Consider single-fluid equations

 $p(\psi)$ is given (i.e. equilibrium model, not transport model)

$$0 = \frac{1}{nm_i} \left(\mathbf{j} \times \mathbf{B} - \nabla p \right)$$

2. Numerically too complicated, too slow for stellarator design optimization.

(i) let $\partial_t = 0$, (ii) ignore non-ideal terms, $\eta, \kappa_{\perp}, \mathbf{\Pi}$, (iii) ignore velocity, \mathbf{v} , (iv) prescribe p, to obtain $\nabla p = \mathbf{j} \times \mathbf{B}, \ldots$ calculation is still too slow . . .

The most widely used equilibrium code for stellarators, VMBRC, and the EPRC code for perturbed tokamaks solve $\nabla p = j \times B$ with rested flux surfaces.

4. Foundation of equilibrium, linear stability, and neo classical transport.

1. Consider single-fluid equations

 $p(\psi)$ is given (i.e. equilibrium model, not transport model)

$$0 = \frac{1}{nm_i} \left(\mathbf{j} \times \mathbf{B} - \nabla p \right) \quad \text{with } \mathbf{B} \cdot \nabla \psi = 0$$

2. Numerically too complicated, too slow for stellarator design optimization.

(i) let $\partial_t = 0$, (ii) ignore non-ideal terms, $\eta, \kappa_{\perp}, \mathbf{\Pi}$, (iii) ignore velocity, \mathbf{v} ,

(iv) prescribe p, to obtain $\nabla p = \mathbf{j} \times \mathbf{B}$, . . . calculation is still too slow . . .

- (v) restrict attention to integrable magnetic fields.
- 3. The most widely used equilibrium code for stellarators, VMEC, and the IPEC code for perturbed tokamaks solve $\nabla p = \mathbf{j} \times \mathbf{B}$ with nested flux surfaces.
- 4. Foundation of equilibrium, linear stability, and neo-classical transport.

Problem: ideal-MHD equilibria with rational flux-surfaces have a <u>non-analytic</u> dependence on the 3D boundary.

1. Breakdown of perturbation theory:

Following [Rosenbluth, Dagazian & Rutherford, Phys. Fluids 16, 1894 (1973)]

".. the standard perturbation theory approach .. is not applicable here due to the singular nature of the lowest order step function solution for $\boldsymbol{\xi}$ "

- 2. The mathematics:
 - i. Equilibrium: $\nabla p \mathbf{j} \times \mathbf{B} = 0$
 - ii. Displacement: $\boldsymbol{\xi} = \epsilon \, \boldsymbol{\xi}_1 + \epsilon^2 \, \boldsymbol{\xi}_2 + ..$
 - iii. Ideal Variations: $\delta \mathbf{B} \equiv \nabla \times (\boldsymbol{\xi} \times \mathbf{B}), \, \delta p \equiv (\gamma 1)\boldsymbol{\xi} \cdot \nabla p \gamma \nabla \cdot (p\boldsymbol{\xi})$
 - iv. Linearized Equation: $\mathcal{L}_0[\boldsymbol{\xi}_1] = \nabla \delta p \delta \mathbf{j} \times \mathbf{B} \mathbf{j} \times \delta \mathbf{B} = 0$

3. " \mathcal{L}_0 is a singular operator .. we must abandon the perturbation theory approach."

- i. The linear displacement is <u>discontinuous</u>.
- \implies ii. The linearly perturbed flux surfaces overlap!
- 4. The singularity in $\mathcal{L}_0 \equiv \nabla_{\mathbf{x}} \mathbf{F}$, where $\mathbf{F} \equiv \nabla p \mathbf{j} \times \mathbf{B}$, also affects:
 - i. iterative solvers, e.g. $\mathbf{x}_{i+1} \equiv \mathbf{x}_i \nabla_{\mathbf{x}} \mathbf{F}^{-1} \cdot \mathbf{F}[\mathbf{x}_i]$,
 - ii. and descent algorithms, e.g. $\frac{\partial \mathbf{x}}{\partial \tau} = -\nabla_{\mathbf{x}} \mathbf{F}$.





Problem: solutions to force balance with nested surfaces have non-integrable singularities in the current-density.

1. $\nabla p = \mathbf{j} \times \mathbf{B}$ yields $\mathbf{j}_{\perp} = \mathbf{B} \times \nabla p/B^2$. Note: \mathbf{j} is current-<u>density</u>, current through \mathcal{S} is $\int_{\mathcal{S}} \mathbf{j} \cdot ds$.

2. Write $\mathbf{j} = \sigma \mathbf{B} + \mathbf{j}_{\perp}$, then $\nabla \cdot \mathbf{j} = 0$ yields $\mathbf{B} \cdot \nabla \sigma = -\nabla \cdot \mathbf{j}_{\perp}$.

3. Nested flux surfaces allows $\mathbf{B} = \nabla \psi \times \nabla \theta + \iota(\psi) \nabla \zeta \times \nabla \psi$, $\sigma \equiv \sum_{m,n} \sigma_{m,n}(\psi) \exp[i(m\theta - n\zeta)]$;



Solution: introduce new class of solutions to $\nabla p = \mathbf{j} \times \mathbf{B}$, with sheet currents and discontinuous transform



 $\mathcal{L}_0[\boldsymbol{\xi}] \equiv \nabla \delta p - \delta \mathbf{j} \times \mathbf{B} - \mathbf{j} \times \delta \mathbf{B} = 0$ is *not* singular if no resonances

In cylindrical geometry, reduces to Newcomb's equation,

$$\frac{d}{dr}\left(f\frac{d\xi}{dr}\right) - g\,\xi = 0$$

Results: perturbation penetrates past "resonant" surface, perturbation amplified by pressure.



RMP penetrates into the core. [Loizu, Hudson *et al.*, Phys. Plasmas **22**, 090704 (2015)]

RMP amplified by pressure near ideal stability limit [Loizu, Hudson *et al.*, Phys. Plasmas **23**, 055703 (2016)]

- 1. Even in ideal-MHD, there are no resonant "shielding" currents at the resonant rational surface.
- 2. The perturbation penetrates, even in cylindrically symmetric equilibria (where there is no poloidal or toroidal "coupling" to lowest order).
- 3. Similar penetration and amplification of RMPs are expected for tokamak equilibria, with consequences for transport.

New solutions match smoothly to "tearing" solution: amplification and penetration of the RMP is still present.

"IDEAL" REGIO

AL" REGIO

"RELAXED" REGION

- 1. Now, include a "relaxed" region,
 - i. $\Delta \psi_t \equiv$ toroidal flux in relaxed region
 - ii. $\Delta t \equiv \text{jump in transform across relaxed region}$
 - iii. "partially" relaxed equilibrium computed with Stepped Pressure Equilibrium Code (SPEC) [Hudson, Dewar et al., Phys. Plasmas 19, 112502 (2012)]
- 2. a magnetic island (tearing mode) arises.



Conclusion: We have solved a fundamental theoretical problem. The classes of general, tractable 3D MHD equilibria are

1. Stepped-pressure equilibria,

- i. [Bruno & Laurence, Commun. Pure Appl. Math. 49, 717 (1996)]
- ii. transform constrained discretely
- iii. pressure discontinuity at ι = irrational
- iv. allows for islands, magnetic fieldline chaos

2. Stepped-transform equilibria,

- i. [Loizu, Hudson et al., Phys. Plasmas 22, 090704 (2015)]
- ii. transform almost everywhere irrational
- iii. arbitrary, smooth pressure
- iv. continuously-nested flux surfaces

3. Or, a combination of the above.

- i. each can be computed using SPEC
- ii. suggests VMEC, NSTAB, IPEC, etc. should be modified to allow for discontinuous transform
- Q. How does a state with continuous transform "ideally evolve" into a 3D state with discontinuous transform? implications for ideal stability if no accessible 3D state exists?



Back up slides

Multi-Region Relaxed MHD, (MRxMHD), and the Stepped Pressure Equilibrium Code, (SPEC)

$$\mathcal{F} \equiv \sum_{i=1}^{N_V} \left[\int_{\mathcal{R}_i} \left(\frac{p}{\gamma - 1} + \frac{B^2}{2} \right) dv - \frac{\mu}{2} \left(\int_{\mathcal{R}_i} \mathbf{A} \cdot \mathbf{B} \, dv - H_i \right) \right]$$

Theoretical developments

- i. Including flow, $\rho v^2/2$, into energy functional, [Dennis, Hudson *et al.*, Phys. Plasmas **21**, 042501 (2014)]
- ii. Including flow & pressure-anisotrophy,
 [Dennis, Hudson *et al.*, Phys. Plasmas **21**, 072512 (2014)]
- iii. Including flow & Hall-effect,
 [Lingam, Abdelhamid & Hudson,
 Phys. Plasmas 23, 082103 (2016)]
- iv. Derivation from Lagrangian variational,
 [Dewar, Yoshida et al., J. Plasma Phys. 81, 515810604 (2015)] free-boundary

 $\mathbf{B} \cdot \mathbf{n} \neq 0$

on ∂ Domain

Numerical developments

- i. finite-elements replaced by Chebshev polynomials
- ii. arbitrary (non-stellarator-symmetric) geometry,
- iii. Cartesian, cylindrical or toroidal,
- iv. linearized equations,
- v. free-boundary equilibria.



Problem: the pressure-driven 1/x current density gives infinite parallel currents through certain surfaces.

Parallel current-density

Parallel current through cross-section $\int_{S} \mathbf{j}_{\parallel} \cdot d\mathbf{s} = \int d\psi \int d\theta \ \sqrt{g} \, \mathbf{j}_{\parallel} \cdot \nabla \zeta$ $\begin{aligned} & = \int_{\epsilon}^{\delta} dx \int_{0}^{\pi/m} d\theta \ \frac{g_{m,n} p'}{x} e^{i(m\theta - n\zeta)} \sqrt{g} \ \mathbf{B} \cdot \nabla\zeta \\ & = g_{m,n,0} p'_{0} \frac{2}{m} \int_{0}^{\delta} dx \frac{1}{x} \end{aligned}$

 $\mathbf{j}_{\parallel} = \sum_{m,n} \left| \frac{g_{m,n} p'}{x} + \Delta_{m,n} \delta_{m,n}(x) \right| e^{(im\theta - in\zeta)} \mathbf{B}.$

$$= g_{m,n,0} p'_0 \frac{2}{m} (\ln \delta - \ln \epsilon) \to \infty \text{ as } \epsilon \to 0.$$

The problem is NOT a lack of numerical resolution.

Is a dense collection of alternating infinite currents physical?

 θ

Shown below is the total *current* through elemental transverse area, for different (m,n) perturbations



If there are rational surfaces, then we must choose:

- 1. flatten pressure near rationals, smooth pressure; ×
- 2. flatten pressure near rationals, fractal pressure; ×
- 3. flatten pressure near rationals, discontinuous pressure; ✓
- 4. restrict attention to "healed" configurations
- [Weitzner, PoP 21, 022515 (2014); Zakharov, JPP 81, 515810609, (2015)]
- 1. Locally-flattened, smooth pressure:

if (*i*.)
$$p'(x) = 0$$
 if $|x - n/m| < \epsilon_{m,n}, \ \forall (n,m),$

and (*ii.*) p'(x) is continuous, then p'(x) = 0, $\forall x$. No pressure!

2. "Diophantine" pressure profile: e.g. from KAM theory

$$p'(x) = \begin{cases} 1, & \text{if } |x - n/m| > r/m^k, \quad \forall (n,m), \text{ e.g. } r = 0.2, \, k = 2, \\ 0, & \text{if } |x - n/m| < r/m^k, \quad \exists (n,m), \end{cases}$$

p'(x) is discontinuous on an uncountable infinity of points,

"The function p is continuous but its derivative is pathological." Grad, Phys. Fluids 10, 137 (1967)]

3. "Stepped" pressure profile: \checkmark

Existence of Three-Dimensional Toroidal MHD Equilibria with Nonconstant Pressure [Bruno & Laurence, Commun. Pure Appl. Math. 49, 717 (1996)]



Culmination of long history of "waterbag" or "sharp-boundary" equilibria: [Potter, Methods Comp. Phys., 16, 43 (1976); Berk et al., Phys. Fluids, 29, 3281 (1986); Kaiser & Salat Phys. Plasmas 1, 281 (1994)]



Relaxed MHD \leftarrow Multi-Region relaxed MHD \rightarrow Ideal MHD

[Taylor, Phys. Rev. Lett. 33, 1139 (1974)]

[Dewar, Hole, Hudson, et al., circa 2006] [Kruskal & Kulsrud, Phys. Fluids 1, 265 (1958)]

$$N_V = 1$$
 Relaxed MHD

$$\mathcal{F} \equiv \underbrace{\int_{\mathcal{R}} \left(\frac{p}{\gamma - 1} + \frac{B^2}{2} \right) dv}_{energy} - \frac{\mu}{2} \underbrace{\int_{\mathcal{R}} \mathbf{A} \cdot \mathbf{B} \, dv}_{helicity},$$

$$\delta \mathbf{B} \equiv \nabla \times \delta \mathbf{A} \text{ is arbitrary in } \mathcal{R}$$
$$(\delta \mathbf{B} = \nabla \times (\boldsymbol{\xi} \times \mathbf{B}) \text{ on } \partial \mathcal{R})$$

+ constrained flux

 $\delta \mathcal{F} = 0, \quad p = p_0, \quad \nabla \times \mathbf{B} = \mu \mathbf{B} \text{ in } \mathcal{R};$

$$\frac{N_V = \infty \quad \text{Ideal MHD}}{\mathcal{F}} \equiv \int_{\mathcal{R}} \left(\frac{p}{\gamma - 1} + \frac{B^2}{2} \right) dv,$$

 $\delta \mathcal{F} = 0, \quad p = p(\psi), \quad \nabla p = \mathbf{j} \times \mathbf{B} \text{ in } \mathcal{R}.$

 $\delta \mathbf{B} = \nabla \times (\boldsymbol{\xi} \times \mathbf{B}) \text{ in } \mathcal{R}$ (fluxes & helicity conserved)

Relaxed MHD
$$\leftarrow$$
 Multi-Region relaxed MHD \rightarrow **Ideal MHD**
[Taylor, Phys. Rev. Lett. 33, 1139 (1974)] [Dewar, Hole, Hudson, et al., circa 2006] [Kruskal & Kulsrud, Phys. Fluids 1, 265 (1959)]
 $N_V = 1$ **Relaxed MHD**
 $\mathcal{F} \equiv \int_{\mathcal{R}} \left(\frac{p}{\gamma-1} + \frac{B^2}{2}\right) dv - \frac{\mu}{2} \int_{\mathcal{R}} \mathbf{A} \cdot \mathbf{B} dv, helicity$
 $\delta \mathcal{F} = 0, \quad p = p_0, \quad \nabla \times \mathbf{B} = \mu \mathbf{B} \text{ in } \mathcal{R};$
 $N_V < \infty$ **MRx MHD**
 $\mathcal{F} \equiv \sum_{i=1}^{N_V} \left\{ \int_{\mathcal{R}_i} \left(\frac{p}{\gamma-1} + \frac{B^2}{2}\right) dv - \frac{\mu_i}{2} \int_{\mathcal{R}_i} \mathbf{A} \cdot \mathbf{B} dv \right\}, \quad \begin{cases} \delta \mathbf{B}_i \equiv \nabla \times \delta \mathbf{A}_i \text{ is arbitrary in } \mathcal{R}_i \\ \delta \mathbf{B}_i \equiv \nabla \times (\boldsymbol{\xi} \times \mathbf{B}_i) \text{ on } \partial \mathcal{R}_i + \text{ constrained flux} \end{cases}$
 $\delta \mathcal{F} = 0, \quad p = p_0, \quad \nabla \times \mathbf{B} = \mu \mathbf{B} \text{ in } \mathcal{R};$
 $\delta \mathcal{F} = 0, \quad p = p_i, \quad \nabla \times \mathbf{B} = \mu_i \mathbf{B} \text{ in } \mathcal{R}_i; \quad \left[\left[p + \frac{B^2}{2} \right] \right] = 0 \text{ across } \partial \mathcal{R}_i; + \text{ constrained fluxes in } \mathcal{R}_i \right]$
 $\delta \mathcal{F} = 0, \quad p = p_i, \quad \nabla \times \mathbf{B} = \mu_i \mathbf{B} \text{ in } \mathcal{R}_i; \quad \left[\left[p + \frac{B^2}{2} \right] \right] = 0 \text{ across } \partial \mathcal{R}_i; - \text{ constrained fluxes in } \mathcal{R}_i \right]$
 $\delta \mathcal{F} = 0, \quad p = p_i, \quad \nabla \times \mathbf{B} = \mu_i \mathbf{B} \text{ in } \mathcal{R}_i; \quad \left[\left[p + \frac{B^2}{2} \right] \right] = 0 \text{ across } \partial \mathcal{R}_i; - \text{ constrained fluxes in } \mathcal{R}_i \right]$
 $\delta \mathcal{F} = \int_{\mathcal{R}} \frac{\mathbf{Ideal MHD}}{\mathcal{F} \equiv \int_{\mathcal{R}} \left(\frac{p}{\gamma-1} + \frac{B^2}{2}\right) dv, \quad \delta \mathbf{B} = \nabla \times (\boldsymbol{\xi} \times \mathbf{B}) \text{ in } \mathcal{R}_i \right]$
 $\delta \mathcal{F} = 0, \quad p = p(\psi), \quad \nabla p = \mathbf{j} \times \mathbf{B} \text{ in } \mathcal{R}.$

Nv=2, "double-Taylor" state with transport barrier; MRxMHD explains self-organization of RFP into helix.

EXPERIMENTAL RESULTS Overview of RFX-mod results

P. Martin et al., *Nucl. Fusion*, **49** 104019 (2009)

Fig.6. Magnetic flux surfaces in the transition from a QSH state . . to a fully developed SHAx state . . The Poincaré plots are obtained considering <u>only the axisymmetric field and dominant perturbation</u>"



NUMERICAL CALCULATION USING STEPPED PRESSURE EQUILIBRIUM CODE Minimally Constrained Model of Self-Organized Helical States in Reversed-Field Pinches G. Dennis, S. Hudson et al. Phys. Rev. Lett. **111**, 055003 (2013)



Excellent qualitative agreement between numerical calculation and experiment.

nature August, 2009 physics

Reversed-field pinch gets self-organized

Compute the 1/x and δ -function current densities in perturbed geometry Self-consistent solutions require **infinite shear**

Cartesian, slab geometry with an (m, n) = (1, 0) resonantly-perturbed boundary

i. $N_V = 3$ MRxMHD calculation, no pressure, $\iota(\psi)$ given discretely,

ii. take limit
$$\Delta \psi \equiv x^{\beta}, \, \epsilon_i = -x^{\alpha}/2, \, \epsilon_{i+1} = +x^{\alpha}/2, \, \text{shear} \equiv \Delta \epsilon/\Delta \psi = x^{\alpha-\beta}, \, \left| \beta > \alpha \right|$$

- iii. island forced to vanish,
- iv. resonant $\delta_{m,n}$ -function current-density appears as tangential discontinuity in **B**.



[Loizu, Hudson et al., Phys. Plasmas 22, 022501 (2015)]

Necessary condition for non-overlapping of perturbed surfaces Existence of non-linear solutions



2. An asymptotic analysis near the rational surface gives the *sine-qua-non* condition (an indispensable condition, element, or factor; something essential)



[Loizu, Hudson et al., Phys. Plasmas 22, 090704 (2015)]

Amplification and penetration as stability boundary is approached

1. Can define a measure of

"Amplification" $A_{rmp} = \xi_s/\epsilon$, where $\epsilon \equiv$ boundary deformation "Penetration" $P_{rmp} = 1 - r_*/r_s$, where $\xi(r_*) \equiv \xi_s/e$

2. A necessary condition for interchange stability in a screw pinch is given by the Suydam criterion, $D_S \equiv -\left(\frac{2p't^2}{rB_z^2t'^2}\right)_{s} < \frac{1}{4}$.



3. Amplification and penetration of RMP **fantastically increased** as stability limit approached.



Discontinuous transform solution cf. "Tearing" solution



SPECallowsdiscontinuousprofiles: exactagreementVMECassumessmoothprofiles: approximateagreement

1. VMEC assumes smooth profiles

and smooth profiles imply discontinuous displacement

2. but, VMEC enforces nested flux surfaces

nested flux surfaces in 3D imply $\frac{\partial \xi}{dr} < 1$ displacement from 2D and this is consistent only with discontinuous transform with $\Delta t > \Delta t_{min}$

3. Empirical study (i.e. radial convergence) shows that

VMEC qualitatively reproduces self-consistent, perturbed solution



Given continuous, non-integrable B, B. $\nabla p = 0$ implies p is fractal. Given fractal p, what is continuous, non-integrable B?

- **Defn.** An equilibrium code computes the magnetic field consistent with a given p and e.g. given t.
- **Theorem.** The topology of **B** is partially dictated by *p*.
 - \hookrightarrow Where $p' \neq 0$, $\mathbf{B} \cdot \nabla p = 0$ implies \mathbf{B} must have flux surfaces.
 - \hookrightarrow Where p' = 0, **B** can have islands, chaos and/or flux surfaces.

TRANSPORT: given \mathbf{B} , solve for p.

- 1. Given general, non-integrable magnetic field, $\mathbf{B} = \nabla \times [\psi \nabla \theta \chi(\psi, \theta, \zeta) \nabla \zeta]$
 - i. fieldline Hamiltonian: $\chi(\psi, \theta, \zeta) = \chi_0(\psi) + \sum_{m,n} \chi_{m,n}(\psi) e^{i(m\theta n\zeta)}$
- 2. KAM theorem: for suff. small perturbation, "sufficiently irrational" flux surfaces survive
 - i. if t satisfies a "Diophantine" condition, $|t n/m| > r/m^k$, $\forall (n, m)$, excluded interval about every rational
 - ii. need e.g. Greene's residue criterion to determine if flux-surface_t exists; lot's of work;

3. With $\mathbf{B} \cdot \nabla p = 0$, i.e. infinite parallel transport, pressure profile must be fractal:

$$p'(t) = \begin{cases} 1, & \text{if } |t - n/m| > r/m^k, \quad \forall (n,m), \text{ e.g. } r = 0.2, \ k = 2, \\ 0, & \text{if } |t - n/m| < r/m^k, \quad \exists (n,m), \end{cases}$$



p'(x) is discontinuous on an uncountable infinity of points; impossible to discretize accurately; EQUILIBRIUM: given p, solve for **B**.

Q. <u>Given</u> a fractal p', how can the topology of **B** be constrained to enforce $\mathbf{B} \cdot \nabla p = 0$?

- i. e.g. if $p(\psi)$ is continuous and smooth, nowhere zero, then **B** must be integrable, i.e. $\chi_{m,n}(\psi) = 0$
- ii. if $p'(\psi)$ is fractal, then what are $\chi_{m,n}(\psi) = ?$

Convergence studies using VMEC

[Lazerson, Loizu et al., Phys. Plasmas 23, 012507 (2016)]



FIG. 2. Profile of the perturbed ρ harmonic (left) and the m = 2 n = 1 component of the toroidal current density (right) showing dependence on radial resolution at fixed shear. Boundary perturbation 1×10^{-4} of minor radius. The q = 2 surface is located at s = 0.5 $(r/a \sim 0.7)$ in this plot. Note that the toroidal current density includes a Jacobian factor.



FIG. 5. Comparison of VMEC response (solid) to Loizu's solution to Newcomb's equation (dotted) (left) and the effective Δi necessary to fit each curve (right). The colors are the same as those in Figure 2, and NS refers to the number of radial grid points.

Published SPEC convergence / verification calculations



FIG. 2. Scaling of components of error, $\delta \mathbf{j} \equiv \mathbf{j} - \mu \mathbf{B}$, with respect to radial resolution. The diamonds are for the n = 3 (cubic) basis functions, the triangles are for the n = 5 (quintic) basis functions. The solid lines have gradient -3, -2, and -2, and the dotted lines have gradient -5, -4, and -4.



FIG. 6. Difference between finite M, N approximation to interface geometry, and a high-resolution reference approximation (with M = 13 and N = 8), plotted against Fourier resolution.



FIG. 2. Convergence of the error between linear and nonlinear SPEC equilibria as ξ_a is decreased, and for different values of Δt , ranging from 10^{-4} (upper curve) to 10^{-1} (lower curve).



FIG. 7. Pressure profile (smooth) from a DIIID reconstruction using STEL-LOPT and stepped-pressure approximation. Also, shown is the inverse rotational transform \equiv safety factor.



FIG. 5. Convergence: the error (Δ) between the continuous pressure (VMEC) and stepped pressure (SPEC) solutions are shown as a function of the number of plasma regions *N* for the *s* = 1/4 SPEC interface. The dotted line shows the zero-beta case ($p_0 = 0$), and the solid line shows the high-beta case ($p_0 = 16$). The grey line has a slope -2, the expected rate of convergence. These simulations were run on a single 3 GHz Intel Xeon 5450 CPU with the longest (the *N* = 128 case) taking 10.1 min using 20 poloidal Fourier harmonics and 768 fifth-order polynomial finite elements in the radial direction.

Early and recent publications

Hole, Hudson & Dewar,	PoP, PoP	2006	(theoretical model)
Huuson, Hole & Dewal,	FUF,	2007	
Dewar, Hole et al.,	Entropy,	2008	(theoretical model)
Hudson Dowar at al	DoD	2012	
Huuson, Dewar et al.,	FUF,	2012	(SPEC)
Dennis, Hudson et al.,	PoP,	2013	(MRxMHD \rightarrow ideal as N _R $\rightarrow \infty$)
Dennis, Hudson et al.,	PRL,	2013	(helical states in RFP = double Taylor state)
Dennis, Hudson et al.,	PoP,	2014	(MRxMHD+flow)
Dennis, Hudson et al.,	PoP,	2014	(MRxMHD+flow+pressure anisotrophy)
Loizu, Hudson et al.,	PoP,	2015	(first ever computation of 1/x & δ current-densities in ideal-MHD)
Loizu, Hudson et al.,	PoP,	2015	(well-defined, 3D MHD with discontinuous transform)
Dewar, Yoshida et al.,	JPP,	2015	(variational formulation of MRxMHD dynamics)
Loizu, Hudson et al.,	PoP,	2016	(pressure amplification of RMPs)

Recent and upcoming invited talks

Hudson, Dewar, et al.,	2012
Dennis, Hudson, et al.,	2013
Dennis, Hudson, et. al.,	2013
Hole, Dewar, et al.,	2014
Loizu, Hudson, et al.,	2015
Loizu, Hudson, et al.,	2015
Hudson, Loizu et al.,	2016
Hudson, Loizu, et al.,	2016
Loizu, Hudson, et al.,	2016

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- International Sherwood Fusion Theory Conference
- 13 International Stellarator Heliotron Workshop
- 14 International Congress on Plasma Physics
- 15 International Sherwood Fusion Theory Conference
- 15 APS-DPP
- 16 International Sherwood Fusion Theory Conference
- 16 Asia Pacific Plasma Theory Conference, 2016
- 16 Varenna Fusion Theory Conference