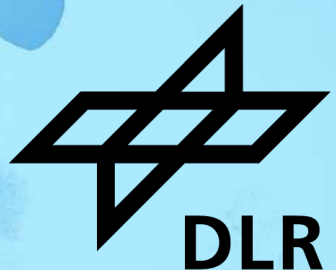


RADIATION EXPOSURE FROM NEUTRONS IN SPACEFLIGHT AND AVIATION

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- Radiation exposure in aviation and space is elevated due to
 - Galactic Cosmic Radiation (GCR)
 - trapped particles in the radiation belts
 - solar energetic particles
- GCR: atomic nuclei with energies $> 100 \text{ MeV/n}$
- Radiation belt: protons (inner belt) and electrons (inner and outer belt)
- Solar energetic particles: electrons and protons (up to hundreds MeV, rarely GeV)
 - **Neutrons are part of the secondary radiation field**
- This talk:
 - aviation, Low Earth Orbit (LEO/ISS), lunar surface
 - GCR, radiation belt

Neutrons in Aviation

- GCR (and rarely high energy solar particles)
- Secondary neutrons from interactions with the atmosphere
- Total dose rates up to $10\mu\text{Sv/h}$
 - 40% - 60% from neutrons

EURADOS, “Cosmic Radiation Exposure of Aircraft Crew”, Final report of the EURADOS WG 5, 2004.

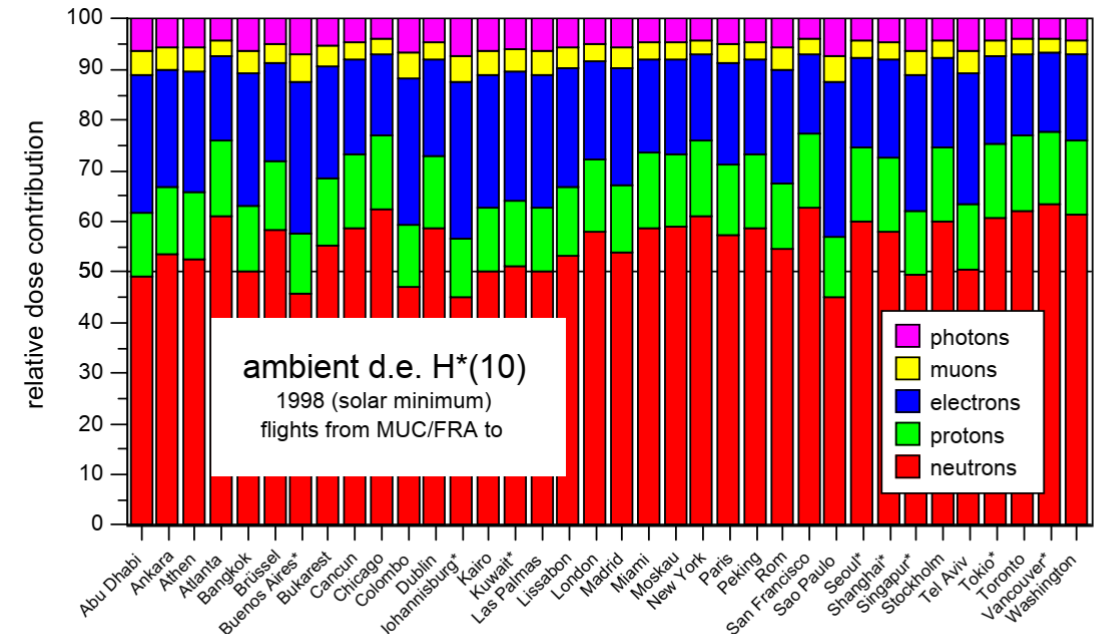
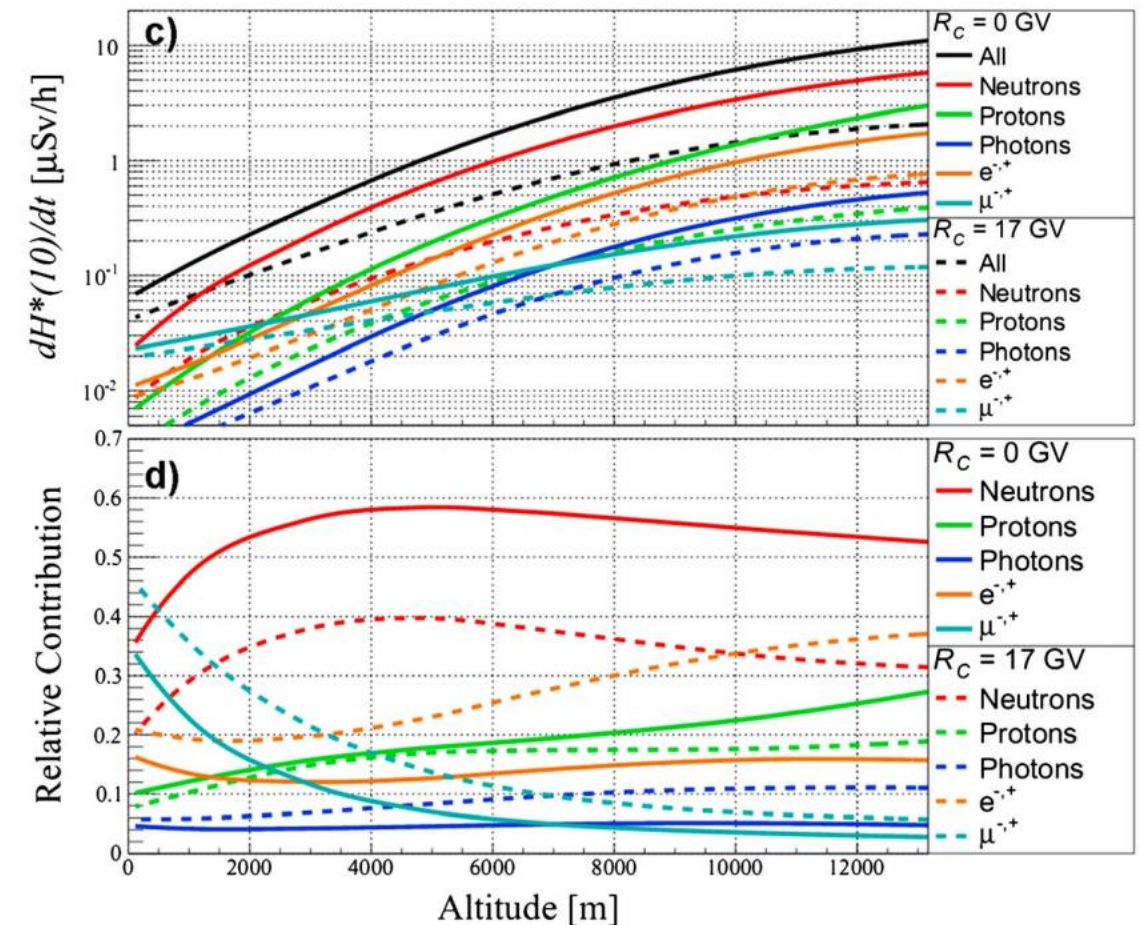


Figure II.6 Relative contribution to ambient dose equivalent for various destinations near solar minimum condition (1998) for 37000 ft flight altitude.

Neutrons in Aviation

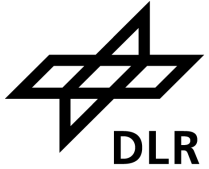
- GCR (and rarely high energy solar particles)
- Secondary neutrons from interactions with the atmosphere
- Total dose rates up to 10 $\mu\text{Sv/h}$
 - 40% - 60% from neutrons
- Neutron contribution decreases
 - with increasing altitude
 - with increasing magnetic shielding

Matthiä et al., Numerical calculation of the radiation exposure from galactic cosmic rays at aviation altitudes with the PANDOCA core model, Space Weather, 2014, [doi: 10.1002/2013SW001022](https://doi.org/10.1002/2013SW001022)

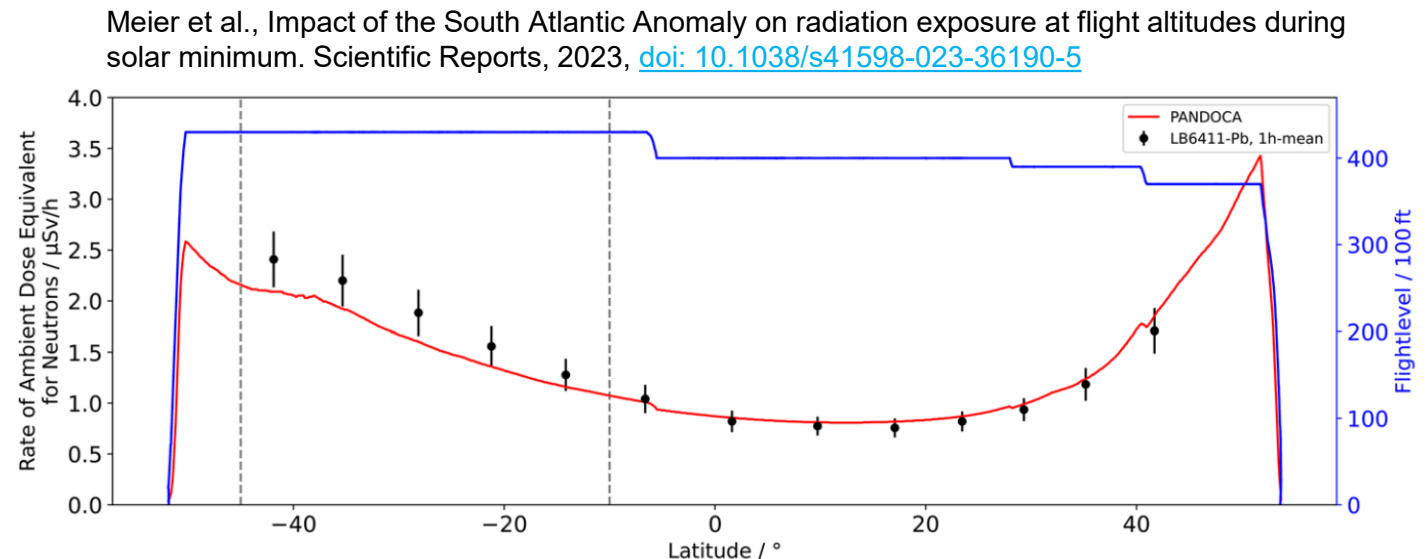


Neutrons in Aviation – Measuring flight to Falkland Islands

Ambient dose equivalent $H^*(10)$



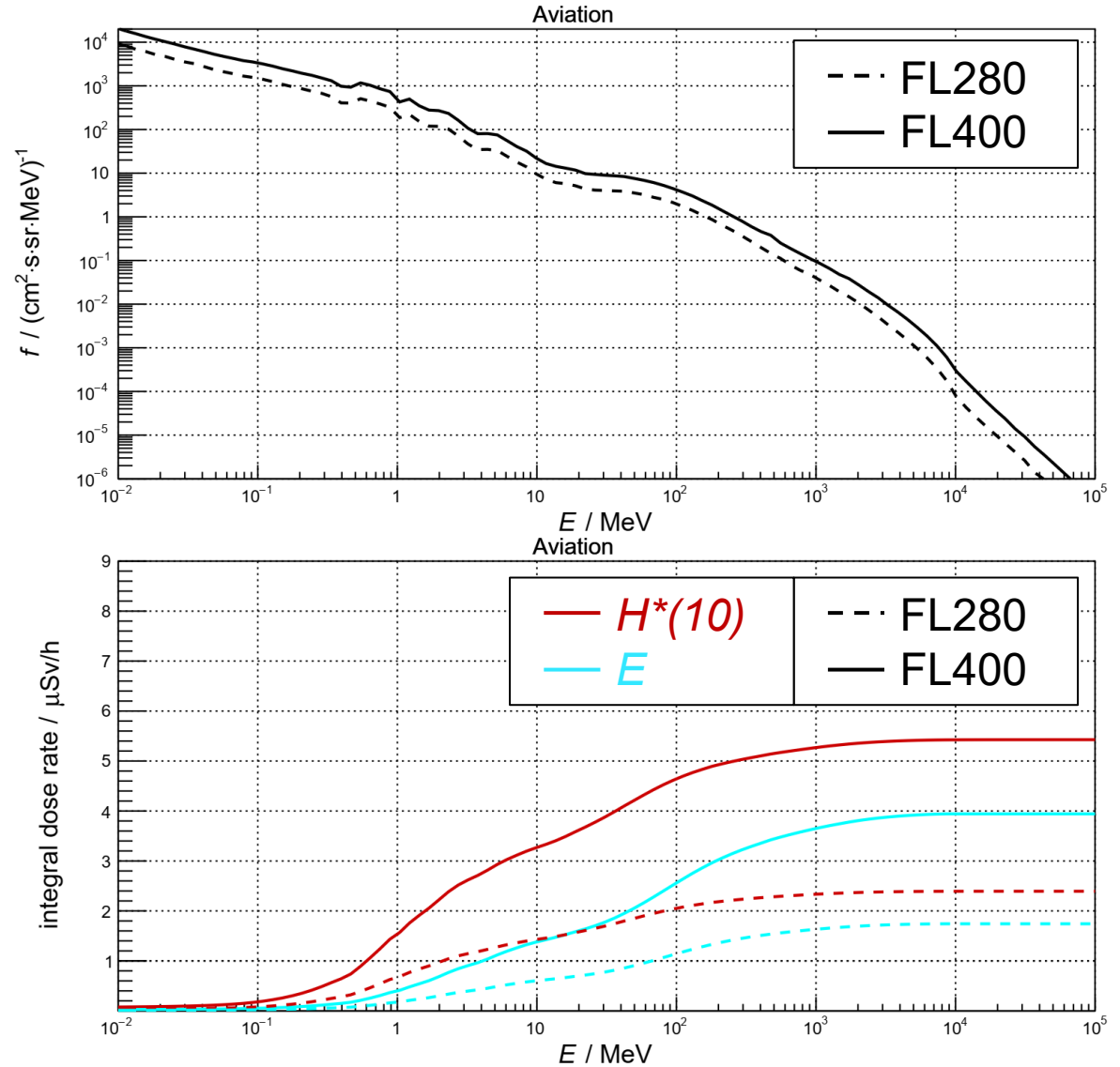
- Measurement $H^*(10)$:
 - total: $51 \pm 3 \mu\text{Sv}$
 - neutron: $23 \pm 2 \mu\text{Sv}$ (45%)
- Model/PANDOCA $H^*(10)$:
 - total: $54 \mu\text{Sv}$
 - neutron: $22.5 \mu\text{Sv}$ (42%)



Ambient dose equivalent rate vs. latitude measured with neutron dose rate monitor LB6411-Pb

Neutrons in Aviation – Neutron spectra at different altitudes (model)

- Energies $\sim 0.1\text{MeV}$ – few GeV contribute
- Ambient dose eq. $H^*(10)$: $2.5 - 5 \mu\text{Sv/h}$
- Effective dose E : $2 - 4 \mu\text{Sv/h}$



Neutrons in Low Earth Orbit International Space Station (ISS)

- GCR, trapped particles (radiation belts)
- Neutrons from interactions with ISS, cargo etc.



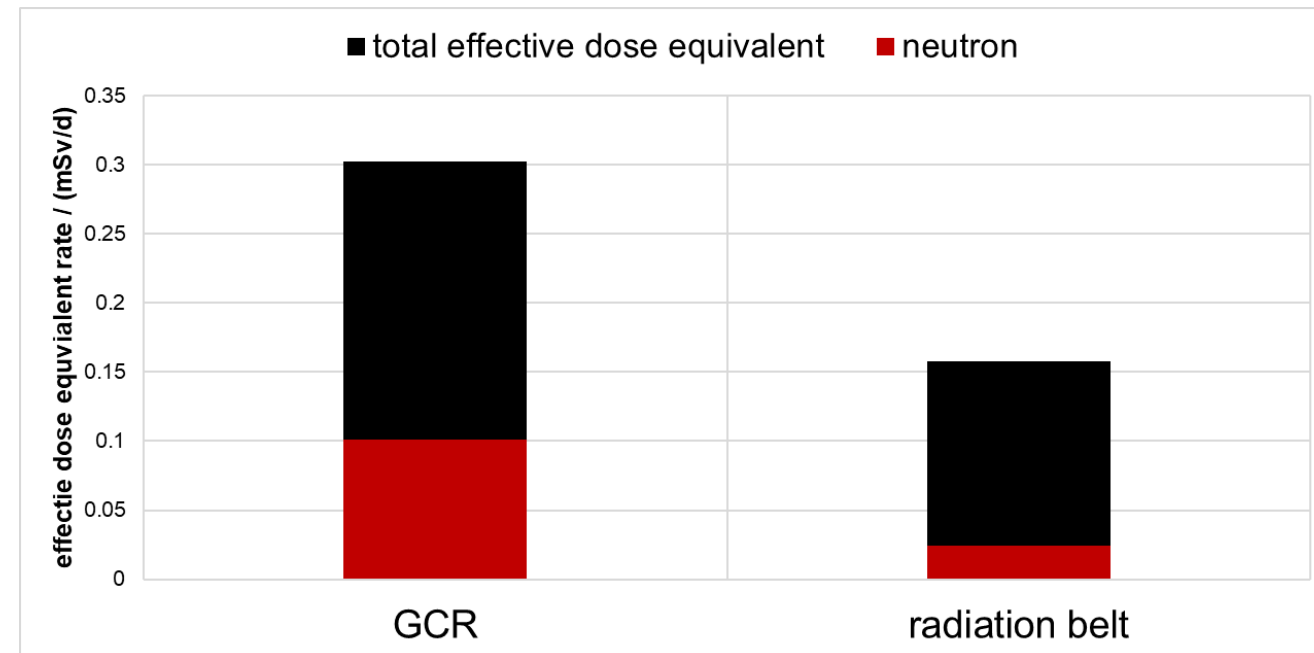
NASA/ESA T. Pesquet

Neutrons in Low Earth Orbit International Space Station (ISS)

- GCR, trapped particles (radiation belts)
- Neutrons from interactions with ISS, cargo etc.
- GCR (COLUMBUS lab.)*
 - total ~ 0.3 mSv/d
 - neutrons ~ 0.1 mSv/d (33%)
- Radiation belt (COLUMBUS lab.)*
 - total ~ 0.16 mSv/d
 - neutrons ~ 0.02 mSv/d (12%)

*model calculations of the total effective dose equivalent for solar minimum

Berger et al., Long term variations of galactic cosmic radiation on board the International Space Station, on the Moon and on the surface of Mars, JSWSC, 2020, [doi: 10.1051/swsc/2020028](https://doi.org/10.1051/swsc/2020028)



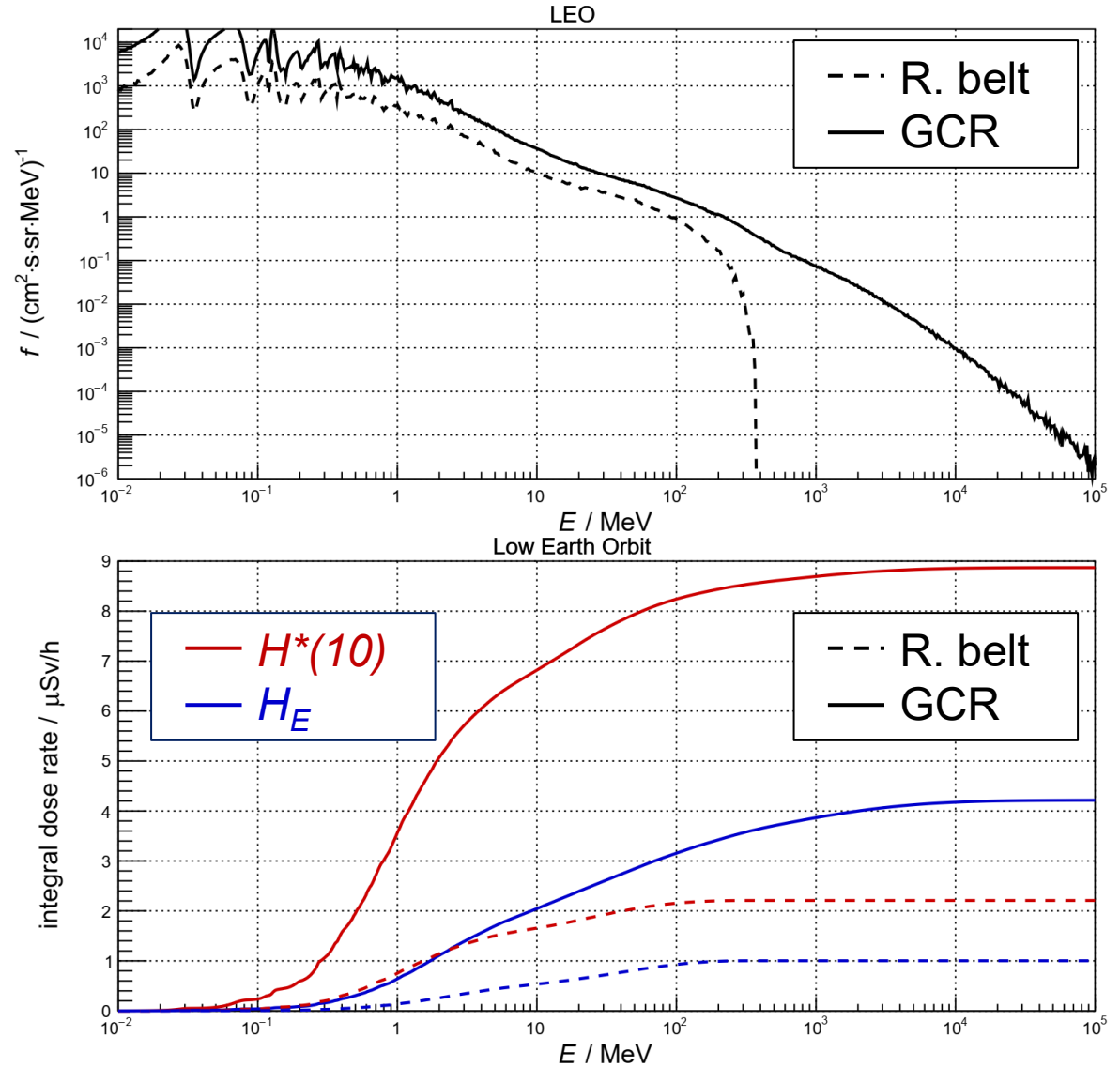
Neutron spectra in Low Earth Orbit International Space Station (model)

■ Radiation Belt:

- energies $\sim 0.1\text{ MeV} - 100\text{ MeV}$
- ambient dose eq. $H^*(10)$: $\sim 2\text{ }\mu\text{Sv/h}$
- effective dose equivalent H_E : $\sim 1\text{ }\mu\text{Sv/h}$

■ GCR:

- energies $\sim 0.1\text{ MeV} - \text{few GeV}$
- ambient dose eq. $H^*(10)$: $\sim 9\text{ }\mu\text{Sv/h}$
- effective dose equivalent H_E : $\sim 4\text{ }\mu\text{Sv/h}$



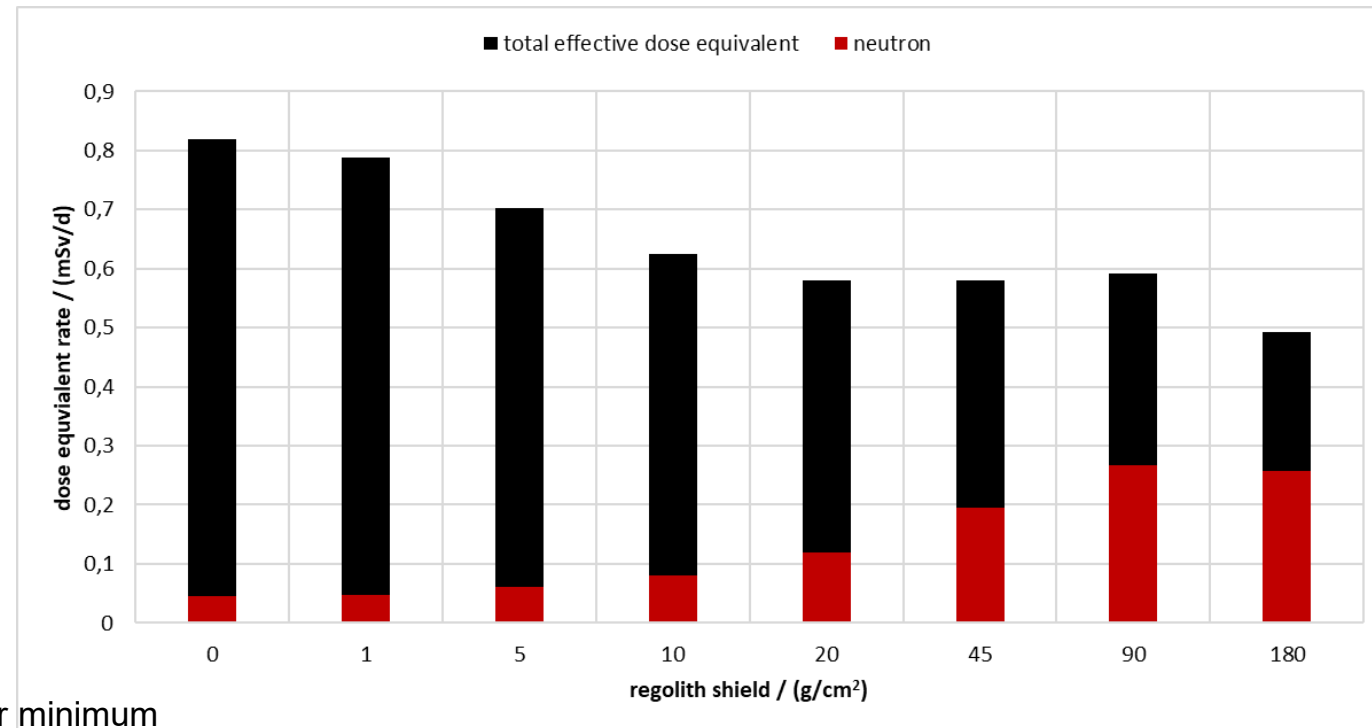
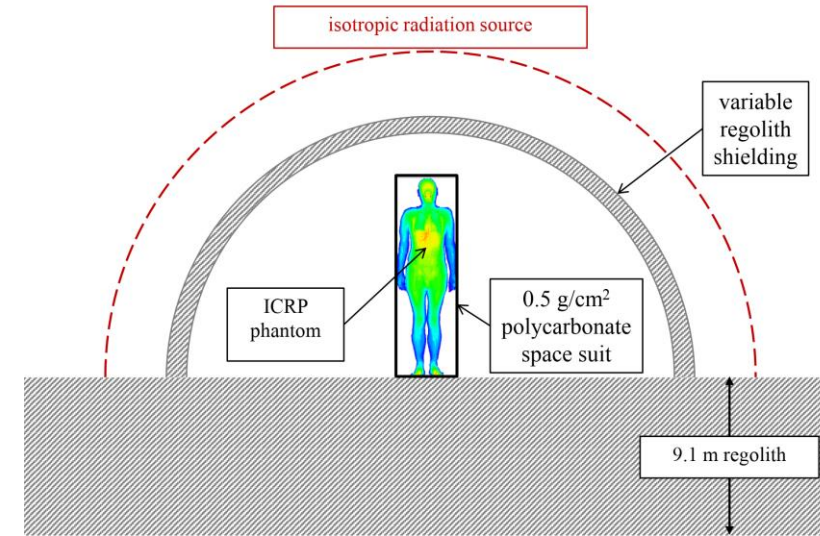
Neutrons on the Lunar Surface

- GCR, solar energetic particles
- Albedo neutrons from interactions with regolith, potentially habitats



Neutrons on the Lunar Surface

- GCR, solar energetic particles
- Albedo neutrons from interactions with regolith, potentially habitats
- GCR (space suit)*
 - total ~ 0.8 mSv/d
 - neutrons ~ 0.05 mSv/d (6%)
- Neutron contribution up to ~50%



*model calculations of the total effective dose equivalent for solar minimum

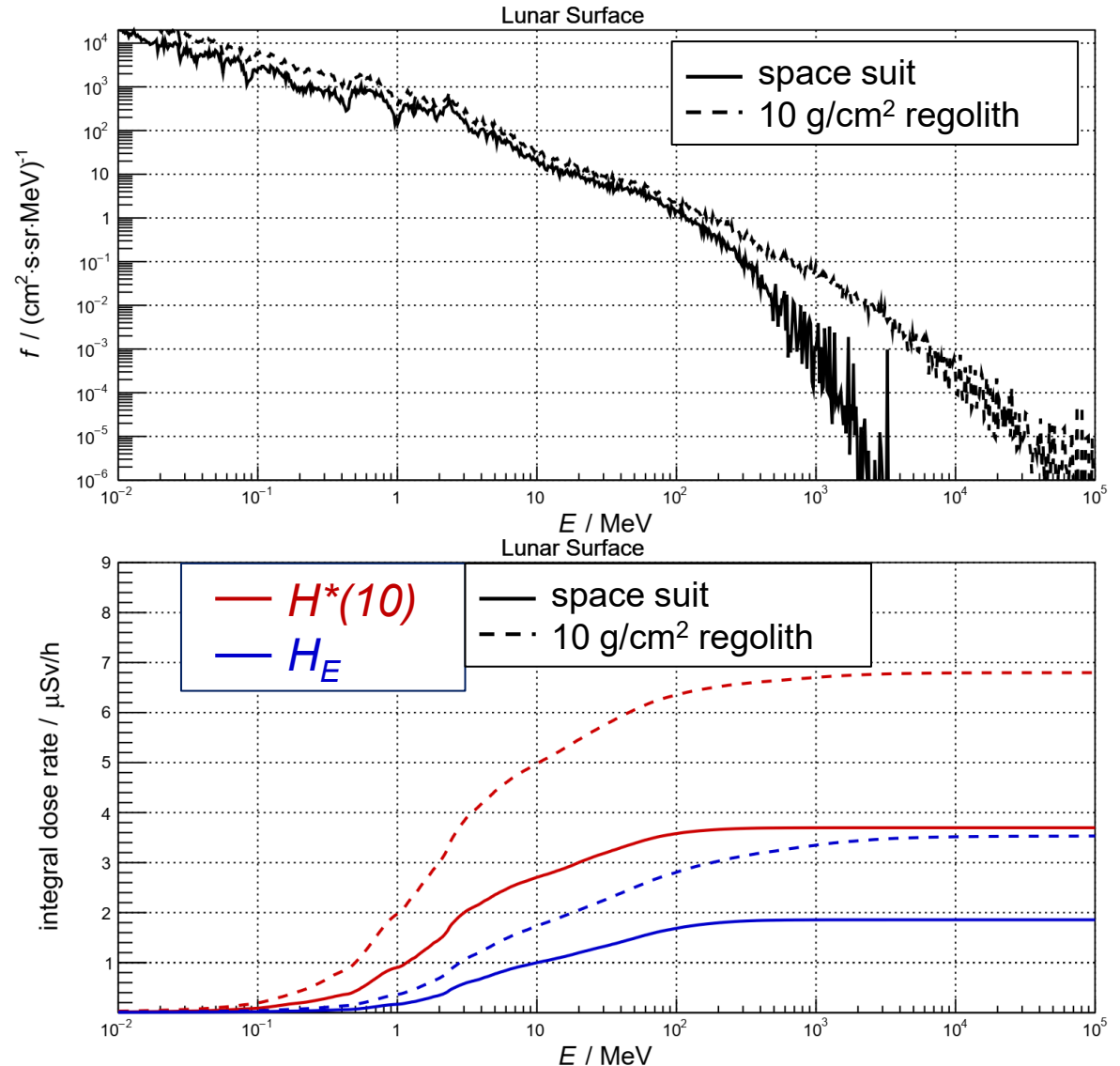
Neutron spectra on the Lunar Surface (model)

■ Space Suit:

- ~ only albedo neutrons
- energies ~0.1 MeV – 100 MeV
- ambient dose eq. $H^*(10)$: ~4 $\mu\text{Sv/h}$
- effective dose equivalent H_E : ~2 $\mu\text{Sv/h}$

■ Habitat (10g/cm² regolith):

- energies ~0.1 MeV – few GeV
- ambient dose eq. $H^*(10)$: ~7 $\mu\text{Sv/h}$
- effective dose equivalent H_E : ~3.5 $\mu\text{Sv/h}$



- Neutrons are one of the most important components of the mixed secondary field from cosmic radiation

- Neutrons contribute 5% to 60% to the total dose ($H^*(10)$, E , H_E) from GCR
 - Aviation: up to $\sim 5 \mu\text{Sv/h}$
 - Space: up to $\sim 10 \mu\text{Sv/h}$

- Relevant energy range $\sim 0.1 \text{ MeV}$ to hundreds/thousands GeV

- Importance of neutrons increases with shielding (atmospheric/spacecraft/habitat)

Thema: Radiation Exposure from Neutrons in Spaceflight and Aviation

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