

Equilibrium Reconstruction in SUNIST-2

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ABSTRACT

- Equilibrium evolution and real-time reconstruction are critical challenges in magnetic confinement fusion, especially for the fast-changing plasmas in SUNIST-2.
- We present a Deep Operator Neural Network (**DeepONet**)^[1] framework that unifies Grad-Shafranov constraints, and diagnostic data.
- Trained on only ~100 discharges, our model achieves **real-time reconstruction without labels** and shows strong generalization.
- This framework can also be extended to build plasma evolution models, enabling data-driven plasma control.

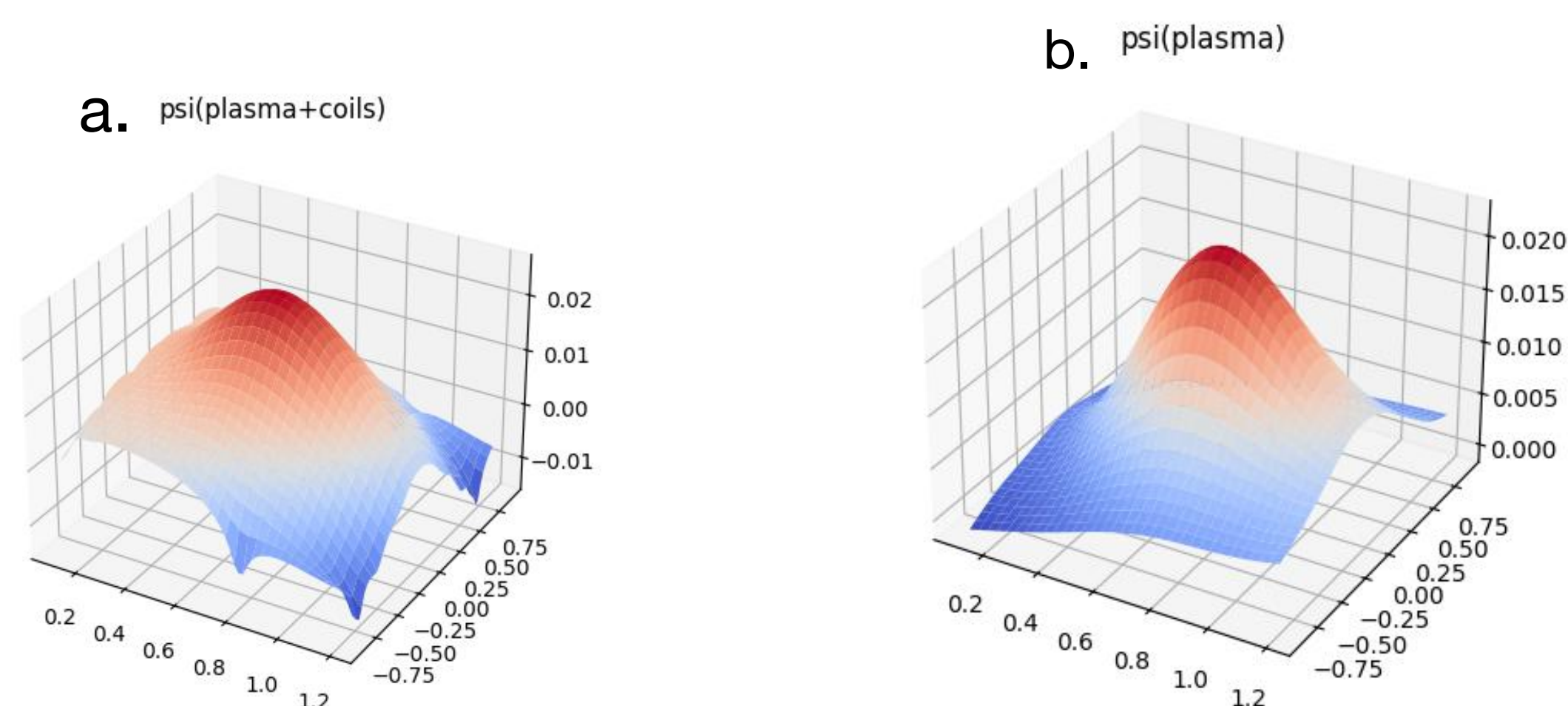
BACKGROUND

- Real-time control needs **real-time** equilibrium.
- SUNIST-2 plasmas change fast and violently.
- rt-EFIT limited by **speed-accuracy trade-off**.
- Surrogate models** with neural networks → balance speed & accuracy, enable robust real-time reconstruction.

METHODS

Dataset

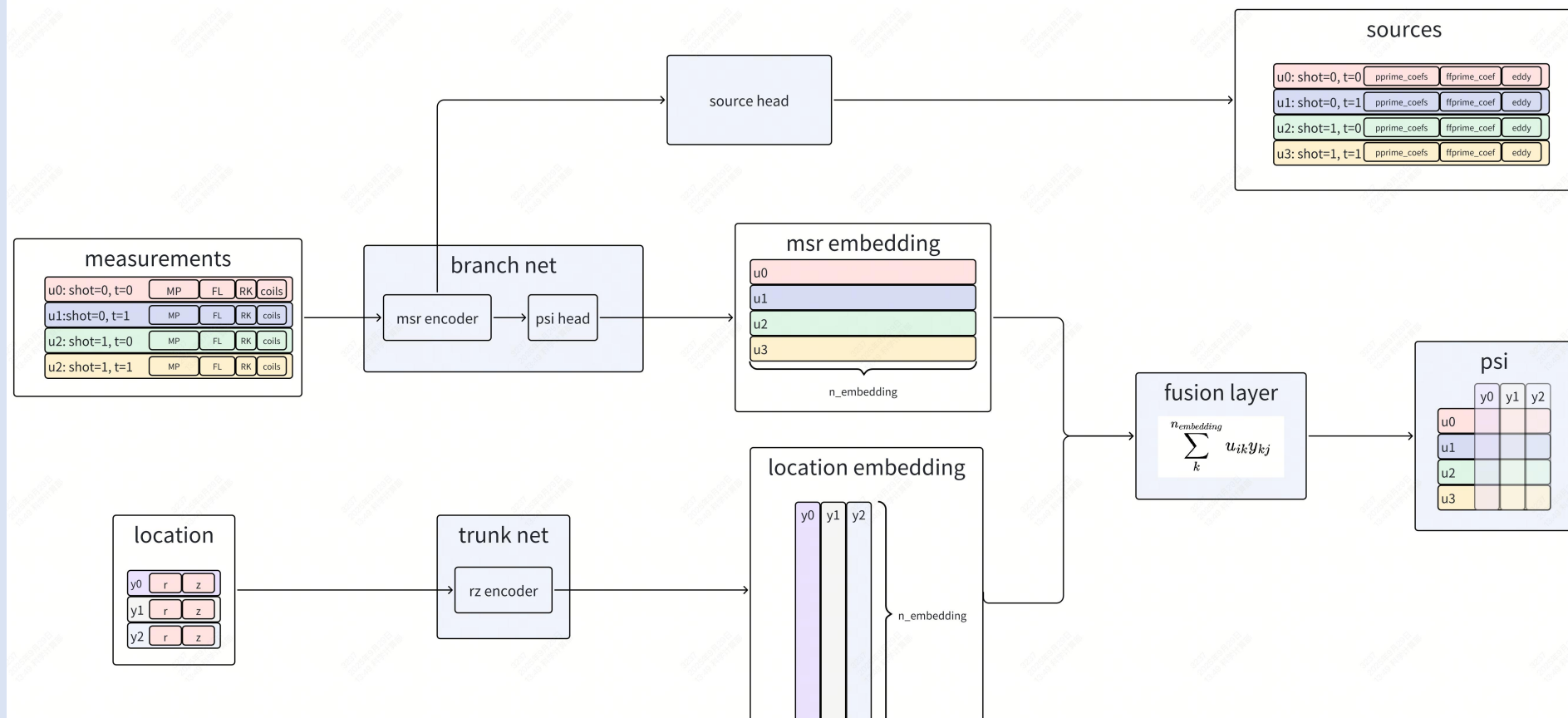
- The dataset contains measured magnetic signals (**magnetic probes, flux loops, and Rogowski coils**) and **PF coil** data.
- A key trick is to predict **ψ contributed only by the plasma**, rather than the combined effect of plasma and coils. The reason is that the ψ landscape becomes much smoother when only the plasma contribution is considered (**a. plasma + coils vs. b. plasma only**), making it easier for the neural network to learn and fit.



Network Architecture & Losses

- Our framework (c) builds upon GS-DeepNet^[2] with some modifications. It adopts the **DeepONet paradigm**, with a branch network encoding measurement signals and a trunk network encoding spatial locations, which are fused to predict ψ . The fusion layer is implemented as a **simple Einstein summation**. A **shared encoder layer** (msr encoder in c) is introduced to enhance generalization.
- We use a polynomial basis for p' and ff' , which is the same as the EFIT paradigm. The **source output** predicts the coefficients of p' and ff' , as well as the eddy currents of the vacuum vessel.

C. network architecture



d. losses

losses	
data loss	magnetic probe, flux loop
physical loss	GS equation
least square loss	$\ Ax-b\ $
constraint loss	non-negative pressure

METHODS

- The loss function combines multiple components: data loss, physical loss, **least-squares loss**, and constraint loss. The least-squares loss, computed from the source output, replaces the conventional least-squares solving procedure used in EFIT.

Training Strategy

We adopt a **stepwise training strategy**:

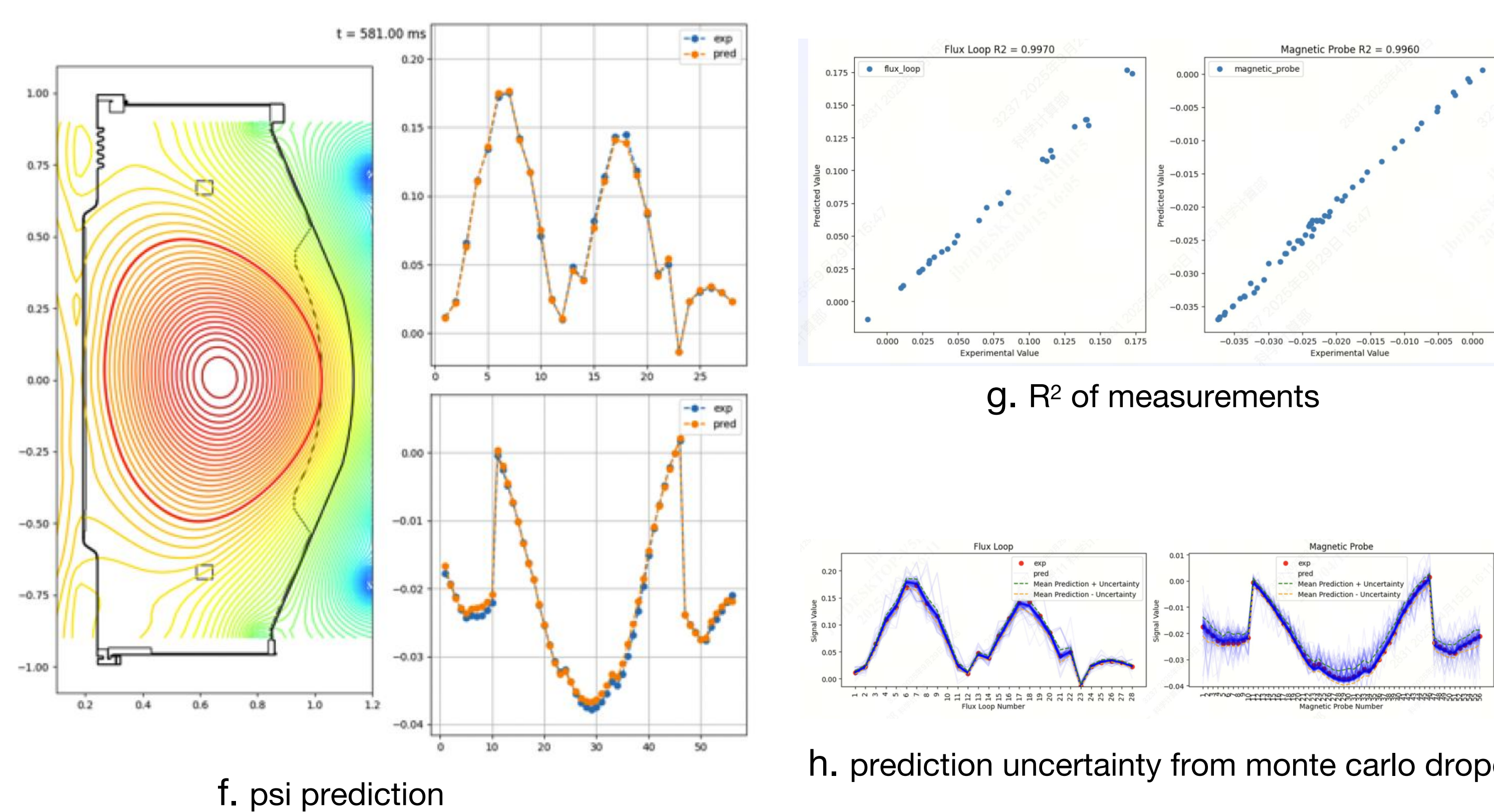
- train ψ -related modules to capture equilibrium representation.
- optimize source-related modules, and freeze the ψ -related modules.
- jointly fine-tune all modules for end-to-end learning.

Thes strategy is really important to get a stable training result.

OUTCOME

Training Result

- Our PINN-based framework successfully reconstructs the plasma equilibrium in SUNIST-2 as shown in **f** (left panels). The measurement signals (right panels) exhibit excellent agreement between prediction and measurement.
- Quantitatively, the model achieves **$R^2 = 0.9970$** for flux loops and **$R^2 = 0.9960$** for magnetic probes as show in **g**, demonstrating high accuracy and robustness.
- Monte Carlo Dropout is employed to estimate prediction uncertainty, providing reliable error bounds as demonstrated in **h**.



CONCLUSION & OUTLOOK

Conclusion

- Source network** improves robustness over least-squares.
- Plasma ψ separation** reduces training difficulty.
- Stepwise training** accelerates convergence and improves stability.

Outlook

- Temporal modeling** for better eddy current prediction and reconstruction continuity.
- Unified framework** for equilibrium evolution and reconstruction.

ACKNOWLEDGEMENTS / REFERENCES

- [1] Lu, L., Jin, P., Pang, G. et al. Nat Mach Intell 3, 218–229 (2021).
- [2] Joung, S., Ghim, Y.C., Kim, J. et al. Sci Rep 13, 15799 (2023).