Characteristics of Tungsten Impurity Sources and Transport in KSTAR Plasmas



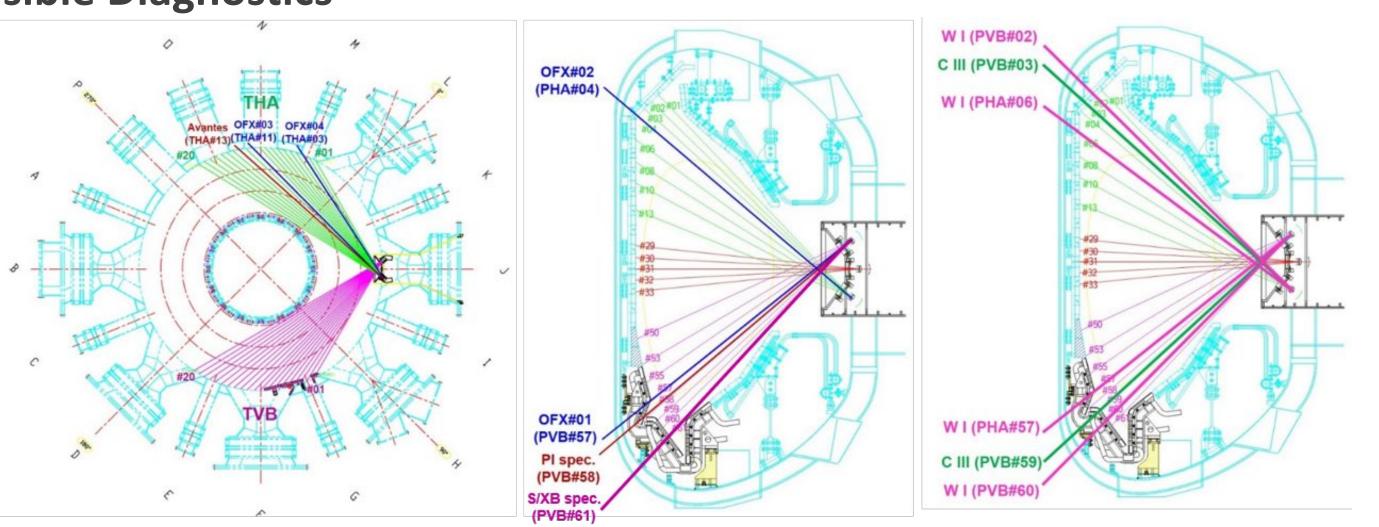
<u>Juhyeok Jang^{1*}</u>, Hyung Ho Lee¹, Alex Grosjean², Dongcheol Seo¹, Junghoo Hwang³,Junhyeok Yoon³, Changmin Shin³, Changrae Seon¹, Seungtae Oh¹ and Eun-nam Bang¹ ¹KFE (Daejeon, Korea), ²University of Tennessee Knoxville (Tennesse, USA), ³KAIST (Daejeon, Korea) Corresponding author: jjh4368@kfe.re.kr

Abstract

- Tungsten (W), the divertor material for ITER, can strongly degrade plasma performance through radiative loss in plasmas. To study physics issues in tungsten sources and transport, KSTAR was upgraded with tungsten monoblock divertor cassettes in 2023.
- KSTAR diagnostics were also upgraded for W measurement, and dedicated W transport experiment in KSTAR were conducted.
- The core plasma radiation and effective charge level is increased by the tungsten impurity transported from the divertor. The core plasma radiation shows asymmetric pattern with tungsten divertor.
- Plasma configuration and plasma-target distance affect W source generation significantly. RMPinduced ELM mitigation further suppressed tungsten generation at the divertor and reduced their core accumulation. Detailed transport analysis is also planned to clarify the underlying mechanisms.
- These findings demonstrate that plasma shaping and RMP application provide control strategies for tungsten impurities in KSTAR plasmas.

Experimental Setup

Visible Diagnostics



Lines of sight of toroidal & Poloidal visible diagnostics (2024)

- H_α monitor : hydrogen/deuterium line intensity at plasma edge [3]
- Visible Bremsstrahlung (VB): bremsstrahlung radiation profile $\rightarrow Z_{\text{eff}}$ profile [4]
- Impurity Filterscope: some channels were modified for impurity monitoring > W I line filter (400.9 nm) was used to estimate W source from divertor.
- Visible Survey Spectroscopy (VSS): impurity line wavelength, width, intensities ➤ Avantes 1ch, OceanFX 4ch: operational in 2024
- S/XB spectrometer 1ch : erosion rate of W divertor target [5]
- All visible diagnostic channels have been calibrated for 2023 2024 campaign.

VIIV & FIIV spectrometers

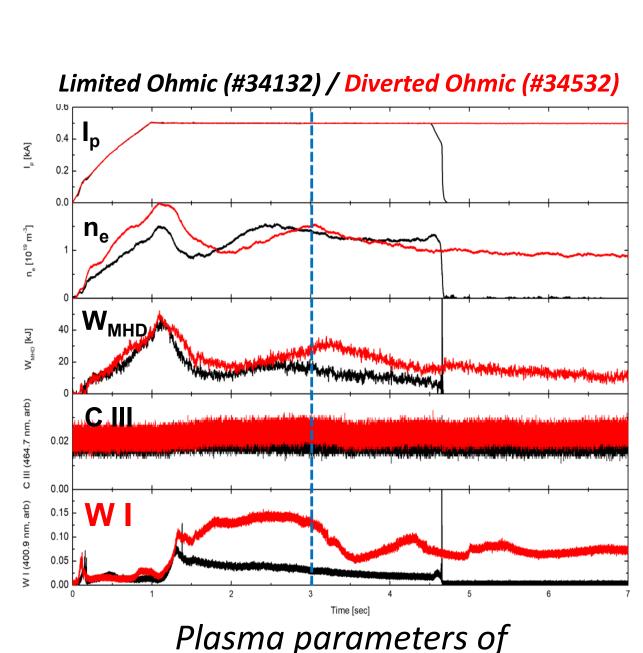
Diagnostics	Observed impurity species	Time res.	Target wavelength	EUV (CAES,
Core VUV survey #1 (PI BI-CCD)	He, C, Ne, Fe, and W	40.8 ms	29.0 – 51.0 nm	VUV (core sur
Core VUV survey #2 (Andor CCD with MCP) [2]	Kr and W	30.6 ms	12 – 32 nm	-0.5
Divertor VUV imaging spectrometer	W (?)			VIVIE 5 28 Th
Compact Advanced EUV Spectrometer (CAES) [4]	B, C, N, O, Ar, and W	102.37 ms	1.8 – 8.3 nm	

 Routine updates of VUV and EUV spectral data to the MDSplus database began with the 2024 KSTAR campaign [6].

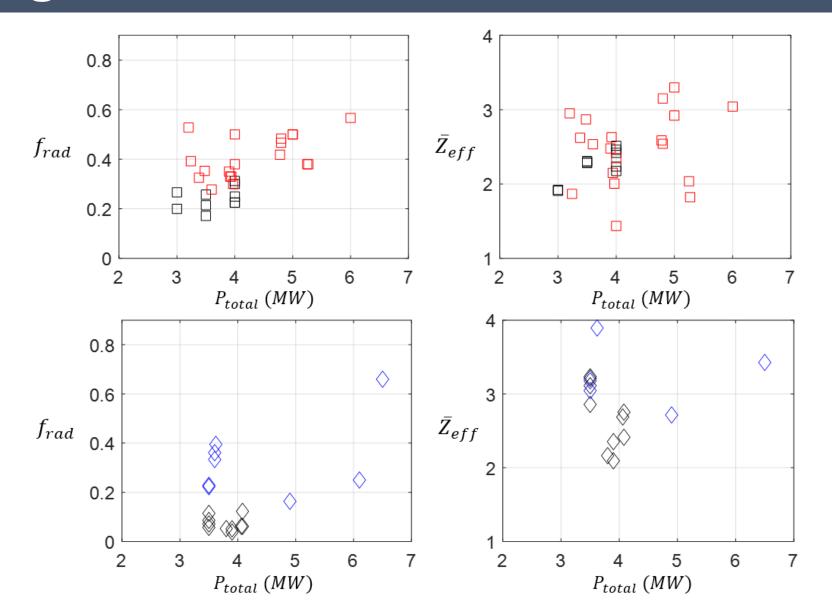
 Routine operation of CAES, for W spectroscopy, also began with the 2024 KSTAR campaign [7].

Divertor VUV imaging spectrometer is under maintenance (mirror box issue).

Radiation and Zeff with Tungsten Source



Limited / Diverted Ohmic discharges

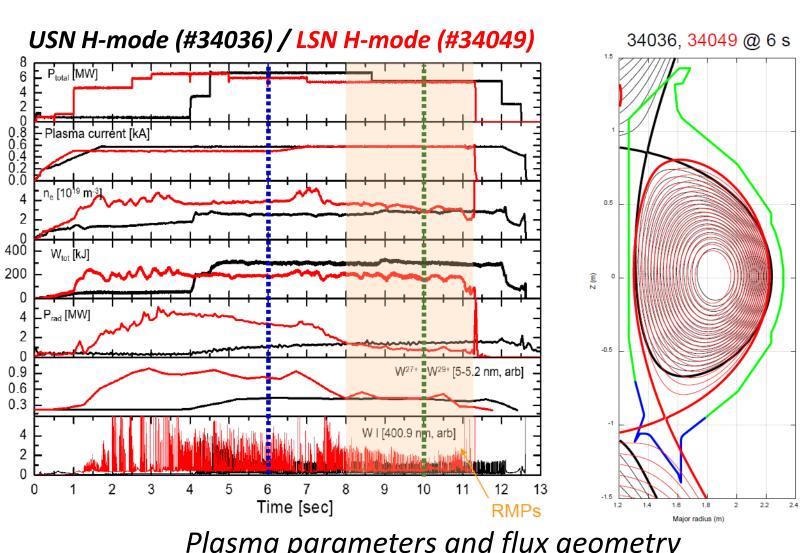


 f_{rad} and \bar{Z}_{eff} before / after W divertor ☐ LN H-mode with Tungsten divertor (2023) ☐ LN H-mode with Graphite divertor (2022) ♦ UN L-mode with Tungsten divertor (2023) ♦ UN L-mode with Graphite divertor (2022)

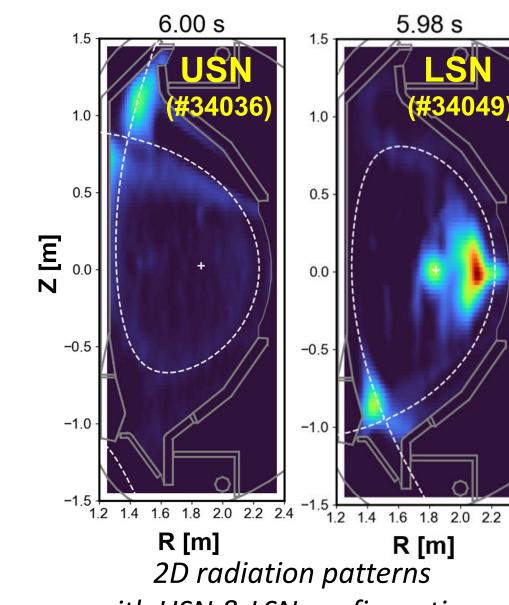
- Visible filterscope channel show that tungsten source is generated after the tungsten divertor installation.
- Radiation fraction (f_{rad}) and line-averaged effective charge (\bar{Z}_{eff}) were typically elevated with tungsten source from lower divertor in KSTAR.

Plasma Shape Effect on W Source & Transport

W source measurements



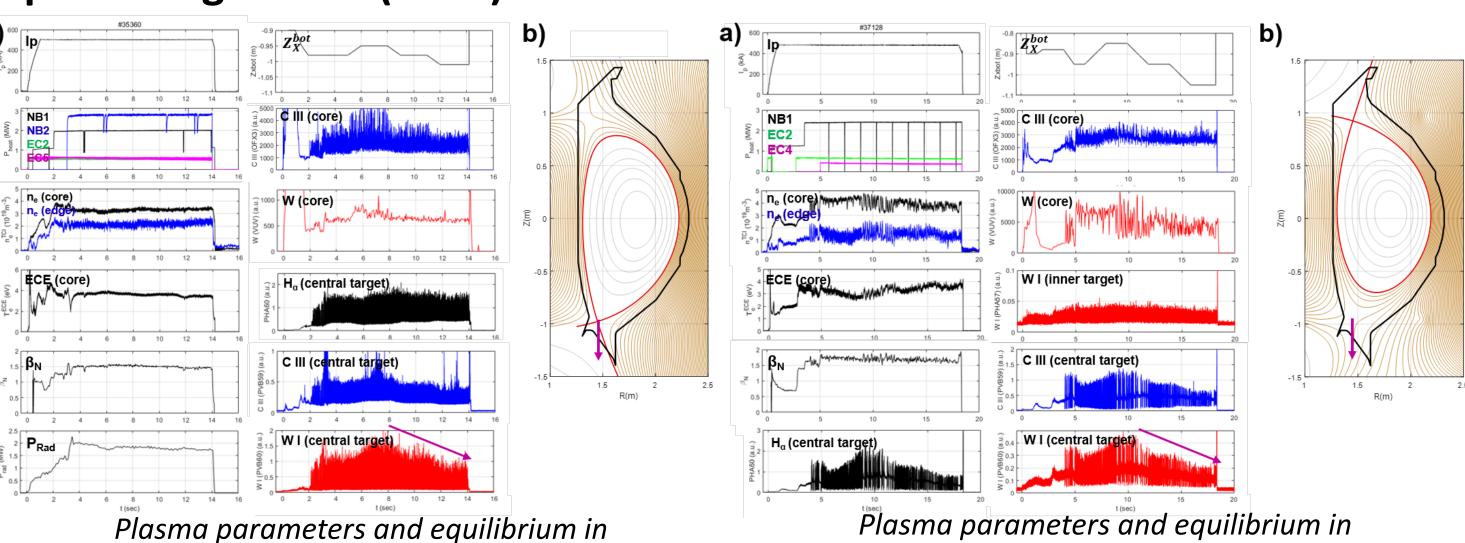




with USN & LSN configurations

- Visible filterscope (W I) showed tungsten source generated from lower divertor
- CAES showed that core W concentration (W 27+ or W29+) in LSN is much higher than that in USN \rightarrow resulting in increased plasma radiation (bolometer image)
- Core W radiation pattern: off-axis due to centrifugal effects

X-point height scan (2024)



lower X-point height scan in **LSN** (#35360) 7.0 sec (+3.0 cm) 10.0 sec (0.0 cm) 13.0 sec (-3.0 cm)

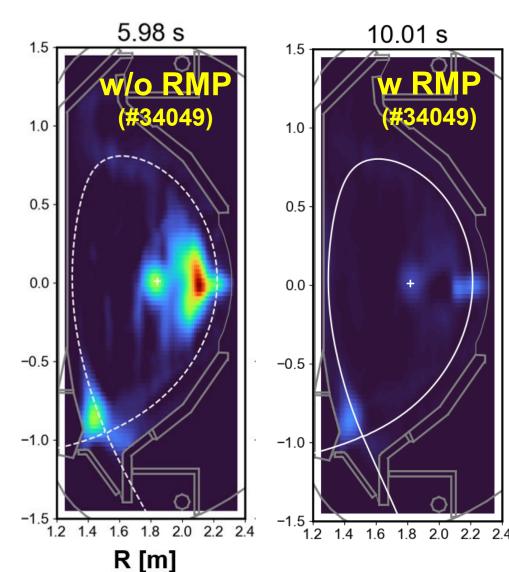
Flux geometry in LSN with X-point displacement

 Lower X-point height scanned by ±3 cm in LSN / USN configurations

lower X-point height scan in **USN** (#37128)

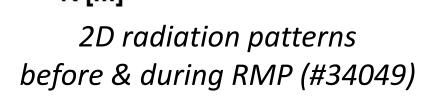
- Significant W source reduction with downward X-point shift observed in both LSN / USN plasmas.
- Slight reduction in core W emission also observed.
- Mechanism of W source reduction not fully understood → requires transport modeling

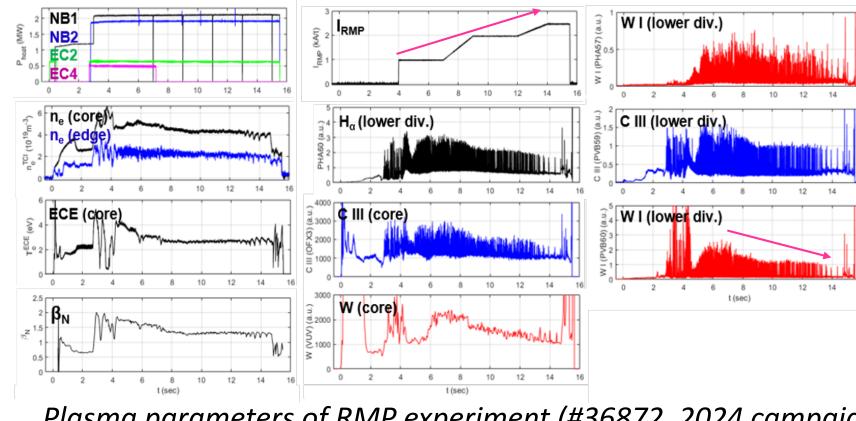
RMP Effect on W Source & Transport



X: W I filterscope channel

X: outer strike point





Plasma parameters of RMP experiment (#36872, 2024 campaign)

- ELM mitigation by n=1 RMP
 - → Reduction of W source by sputtering
- Bolometer images showed that plasma radiation by W impurity reduced
- VSS and filterscope showed that W I and C III intensities reduced in both lower divertor and plasma core.

Summary & Future Works

- Visible and VUV diagnostics in KSTAR were upgraded for W transport study.
- After tungsten divertor installation in KSTAR, radiation fraction (f_{rad}) and effective charge (Z_{eff}) elevated significantly.
- Tungsten source strongly depends on plasma configuration, plasma—target distance and ELM activity.
- RMPs effectively mitigate ELMs and suppress W impurity generation.
- Future work
 - Quantitative W influx via S/XB method
 - Follow-up experiments with L-mode plasma and more W filterscope channels
 - Transport modeling with SOLEDGE3X-ERO [10]

Acknowledgements / References

This research was supported by the R&D Program "KSTAR Experimental Collaboration and Fusion Research (Grant No. EN2501)" through the Korea Institute of Fusion Energy (KFE), funded by the Government funds, Republic of Korea.

[1] Lee, H.H., et al., 3rd International Fusion and Plasma Conference, Seoul (2024) [2] Jang, J., et al., 3rd International Fusion and Plasma Conference,

Seoul (2024) [3] Na, H.K., et al., Fus. Eng. Des., 86 (2011) 66 [4] Jang, J., et al., Rev. Sci. Instrum. 95 (2024) 013503 [5] Van Rooij G.J., et al., J. Nucl. Mater. 438 (2013) S42

[6] Seon, C.R. et al., Rev. Sci. Instrum., 88 (2017) 083511 [7] Song, I., et al., Rev. Sci. Instrum., 88 (2017) 093509 [8] Zhang, W., et al., Nucl. Fusion 64 (2024) 086004 [9] Dai, S.Y., et al., Nucl. Fusion 63 (2020) 025003 [10] Bufferand, H., et al., Nucl. Mater. Energy, 18 (2019) 82