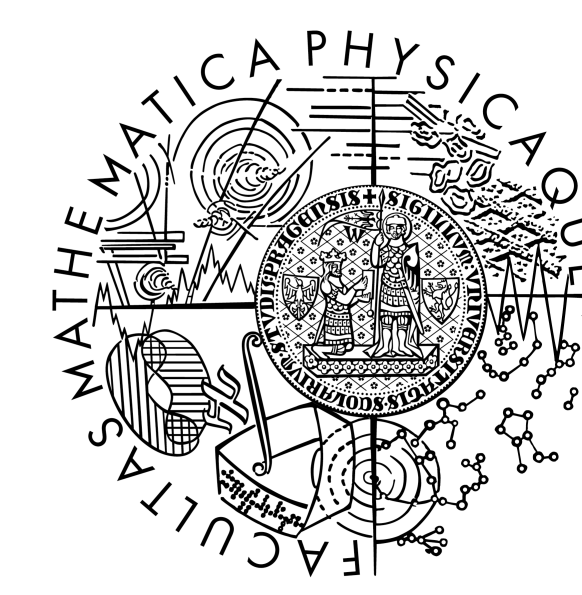


# What can we say about the dipole photon strength in $^{57}\text{Fe}$ from the $(n_{\text{th}}, \gamma)$ data?



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## Introduction

In the statistical model of  $\gamma$  decay, the photon strength function (PSF),  $f^{(XL)}$ , determines the partial radiative width  $\Gamma_{i\gamma f}$  from the initial level  $i$  to the final level  $f$  as

$$\Gamma_{i\gamma f} = \sum_{XL} f^{(XL)} \xi_{XL} E_{\gamma}^{2L+1} D_{J_i} = \sum_{XL} f_i^{(XL)} E_{\gamma}^{2L+1} D_{J_i}, \quad (1)$$

where  $D_{J_i}$  is the average spacing for given  $J_i^{\pi}$  in the vicinity of the initial level and the factors  $\xi_{XL}$  reflect the fluctuations of  $\Gamma_{i\gamma f}$  or the individual PSF values,  $f_i^{(XL)}$ . These fluctuations are assumed to follow the Porter-Thomas distribution. The summation goes over all allowed types  $X$  and multiplicities  $L$  of transition connecting the initial and final level. In practice the decay of highly excited levels is usually dominated by dipole transitions,  $E1$  or  $M1$ , with the only possible relevant mixed transition being  $M1 + E2$ . For the sake of clarity we omit the  $XL$  labeling hereafter.

- the  $E1$  PSF above neutron separation energy,  $S_n$ , dominated by the giant electric dipole resonance (GEDR)
- the  $E1$  PSF shape below  $S_n$  still uncertain, often found inconsistent with standard Lorentzian description of GEDR
- other resonances play a role - pygmy  $E1$ , spin-flip and scissors  $M1$
- the first observation of low-energy enhancement (LEE) in  $^{56,57}\text{Fe}$  by the Oslo method [1]
- the exact magnitude depends on normalization of the data [2, 3], see Fig. 1
- the presence and magnitude of LEE has impact on  $(n, \gamma)$  cross section predictions

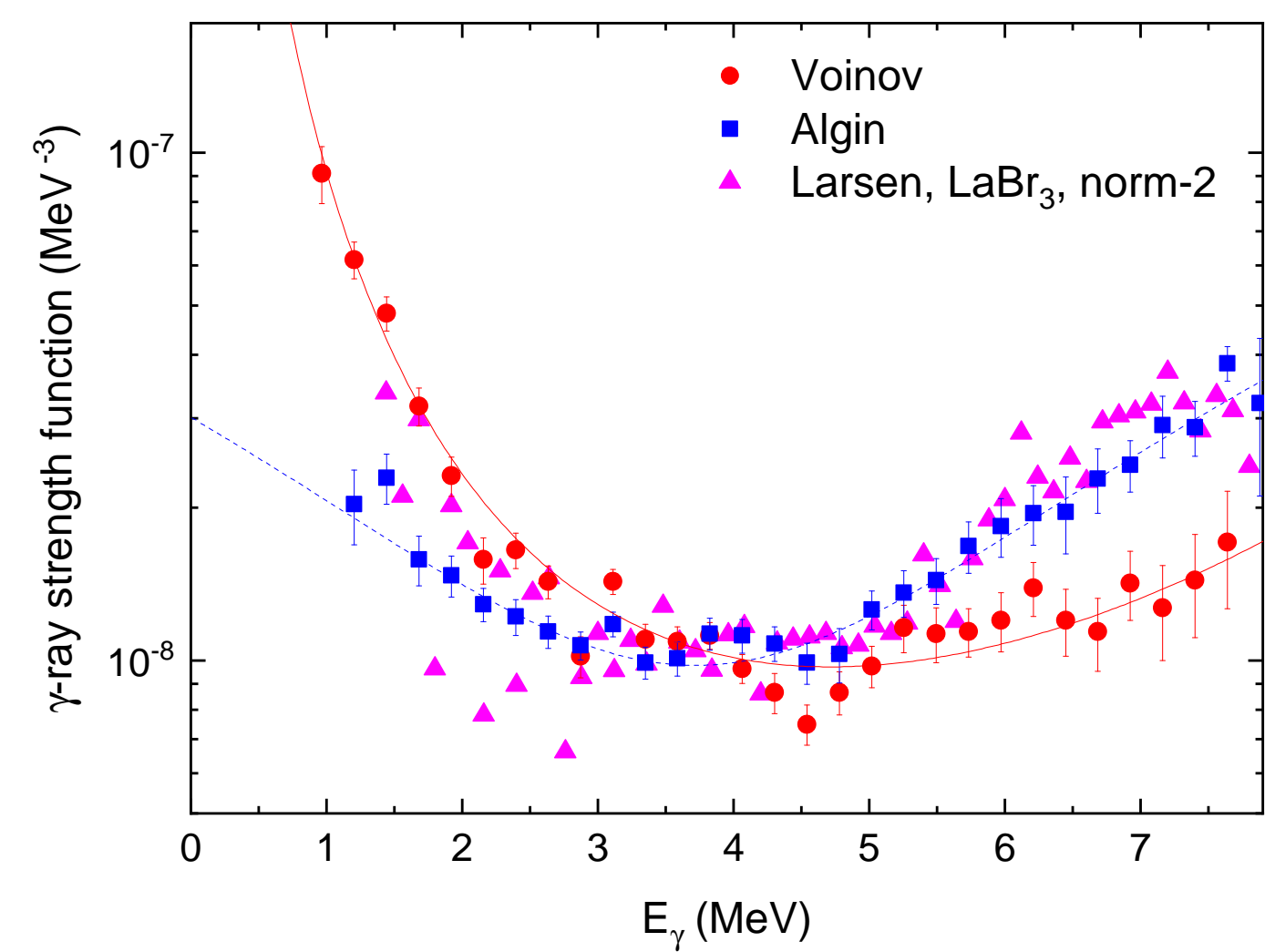


Figure 1: The dipole PSF as determined by the Oslo method. The datasets correspond to Refs. [1, 2, 3], respectively. The uncertainties of Larsen dataset are omitted for the sake of clarity.

## Thermal neutron capture measurements

The  $^{56}\text{Fe}(n_{\text{th}}, \gamma)$  was measured at Cold neutron PGAA facility at Budapest [4] and thermal neutron two-step  $\gamma$  cascade (TSC) facility at Řež [5]. The singles spectrum from Budapest was used to get intensity values, while the TSC spectra were used to check the placement of some transitions.

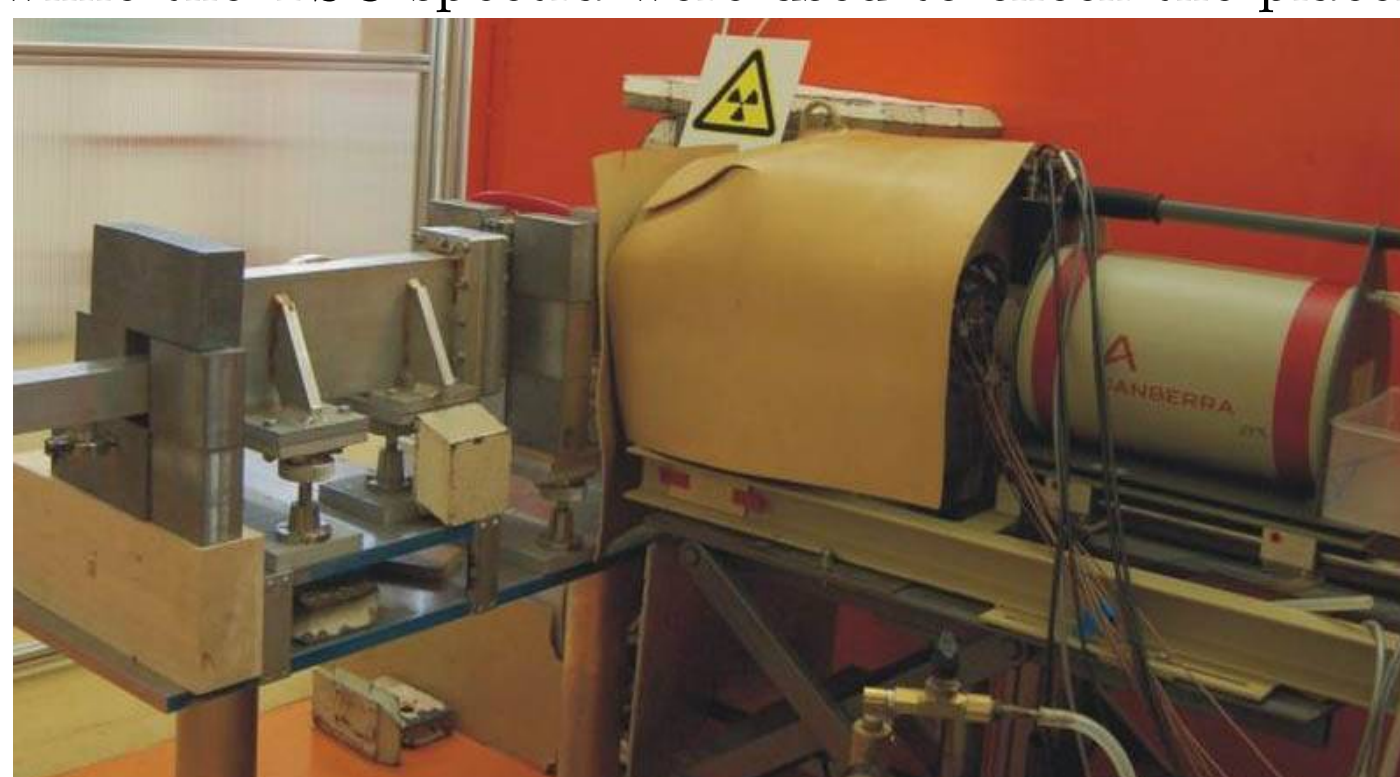


Figure 2: The PGAA apparatus at Budapest Neutron Center. The photo was taken from <https://www.iperi-onhs.eu/tool/70/>.



Figure 3: The TSC gamma coincidence facility at LWR-15 research reactor, NPI CAS Řež at the time of measurement in 2005.

## Intensities of transitions

A detailed level scheme constructed by Firestone *et al.* [6], authors deduced **transition probabilities**  $P_{\gamma}$  per neutron capture.

- 98 levels in  $^{56}\text{Fe}$  with 448  $\gamma$ -rays between them, substantial revision w.r.t. ENSDF [7] and references therein
- total observed **intensity**  $P_{\gamma} E_{\gamma} / S_n = 0.989(14)$
- 24 unplaced transitions account only for 0.14% of total observed intensity
- **primary transitions** - 85 observed, 32 of known  $XL$ , see Fig. 4
- 3 high-energy  $\gamma$ -rays observed, which can not be primaries in the present level scheme  $\Rightarrow$  we can assume their  $P_{\gamma}$  from the corresponding secondaries, shown as “unobserved” in Fig. 4
- improved thermal neutron capture cross section  $\sigma_{\gamma} = 2.394(19)$  b.

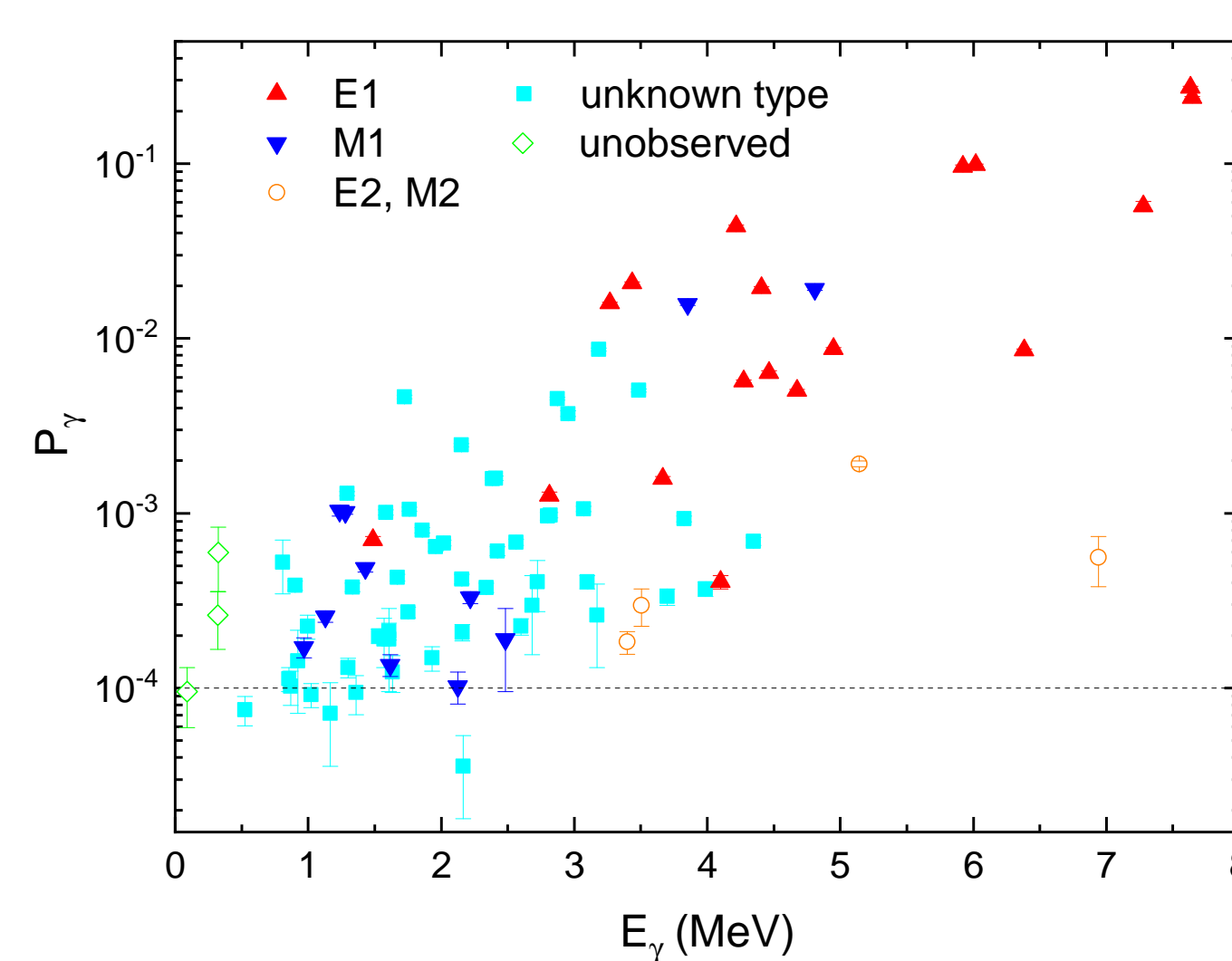


Figure 4: Probability of a primary transition as a function of  $\gamma$ -ray energy. The black dashed line corresponds to the observation threshold assumed in further analysis.

According to DICEBOX simulations, missing transition probability is  $\sum P_{\gamma} \approx 0.3\%$  for primary transitions, and  $\sum P_{\gamma} \approx 0.47\%$  for feeding of GS + 14 keV levels *assuming* the constant threshold of  $10^{-4}$  shown in Fig. 4. **Unique dataset** going down to **very low  $\gamma$ -ray energy**.

Despite the significant progress, many levels do not have a definite  $J^{\pi}$  assignment, hence many transitions remain of unknown type. In further analysis we assume these are dominated by dipole transitions.

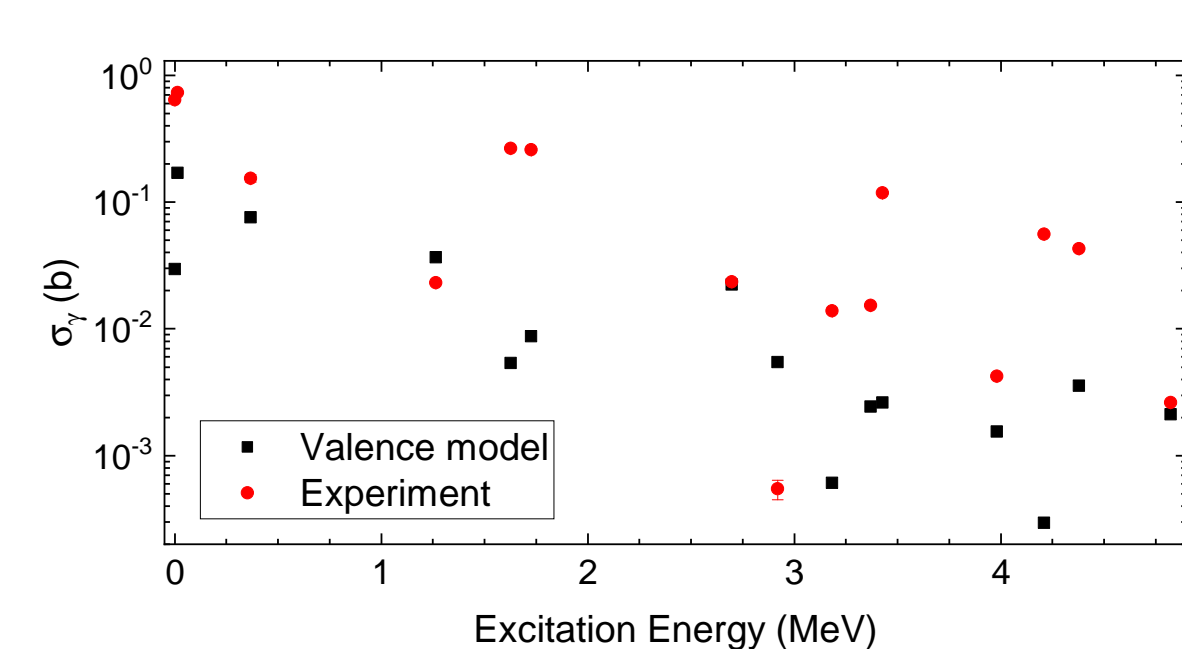


Figure 5: The possible influence of non-statistical effects. Plotted is the partial thermal neutron capture cross section as a function of excitation energy. The experimental points are deduced from probability of corresponding primary transition, the predictions of valence model are calculated according to Ref. [8].

## Photon strength function

The individual PSF values  $f_i$  are obtained from probabilities of transitions as

$$f_i = \frac{\Gamma_{i\gamma f}}{E_{\gamma}^{2L+1} D_{J_i}} = \frac{\Gamma_{i\gamma} (P_{i\gamma f} / \sum_f P_{i\gamma f})}{E_{\gamma}^{2L+1} D_{J_i}}, \quad (2)$$

we omit the  $i$  and  $f$  indices for the sake of clarity.

- individual values are expected to fluctuate according to Porter-Thomas distribution, see Eq. (1)
- $\Rightarrow$  to obtain the mean PSF value we need to average
- $\times$  only handful of transitions in relatively wide, high-energy  $\gamma$ -ray interval
- need to assume a PSF shape within each interval, e.g. constant
- ! influence of threshold in low-energy intervals – the result slightly depends on nuclear level density model
- ! the resulting mean PSF value can be below threshold
- many primaries are of unknown type  $\Rightarrow$  assumptions about  $E1$ -to- $M1$  PSF ratio

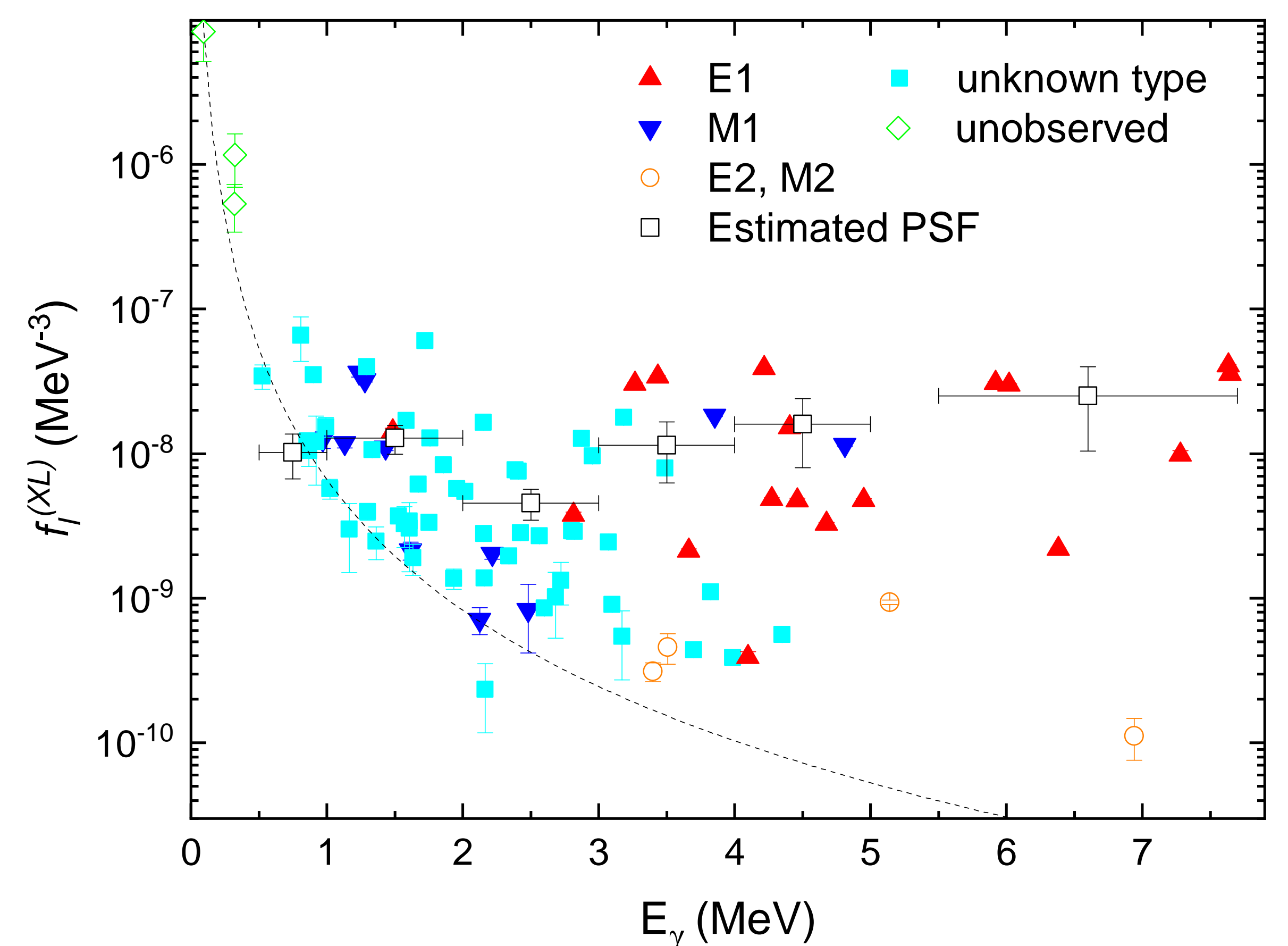


Figure 6: PSF deduced from intensities of primary transitions as a function of  $\gamma$ -ray energy. The black dashed line corresponds to the threshold in Fig. 4.

To convert probabilities of primary transitions to PSF values, we have used  $D_{J_i} = D_0 = 22.0(17)$  keV and  $\Gamma_{i\gamma} = 1474$  meV from Mughabghab's atlas [8]. The resonance that dominantly contributes to the thermal capture is the bound state according to Mughabghab, hence the choice of  $\Gamma_{i\gamma}$ .

The average total radiative width is quoted as 900(470) meV therein. The uncertainty probably reflects the expected significant fluctuations of  $\Gamma_{i\gamma}$  in such a light nucleus, see Fig. 7.

In any case, the value of  $\Gamma_{i\gamma}$  does not influence the *shape* of deduced PSF.

The integrated TSC data can be used for comparison with DICEBOX simulations under different assumptions. The sensitivity is limited because of sizable P-T fluctuations when intensities of primaries are not taken from experiment, see Fig. 8.

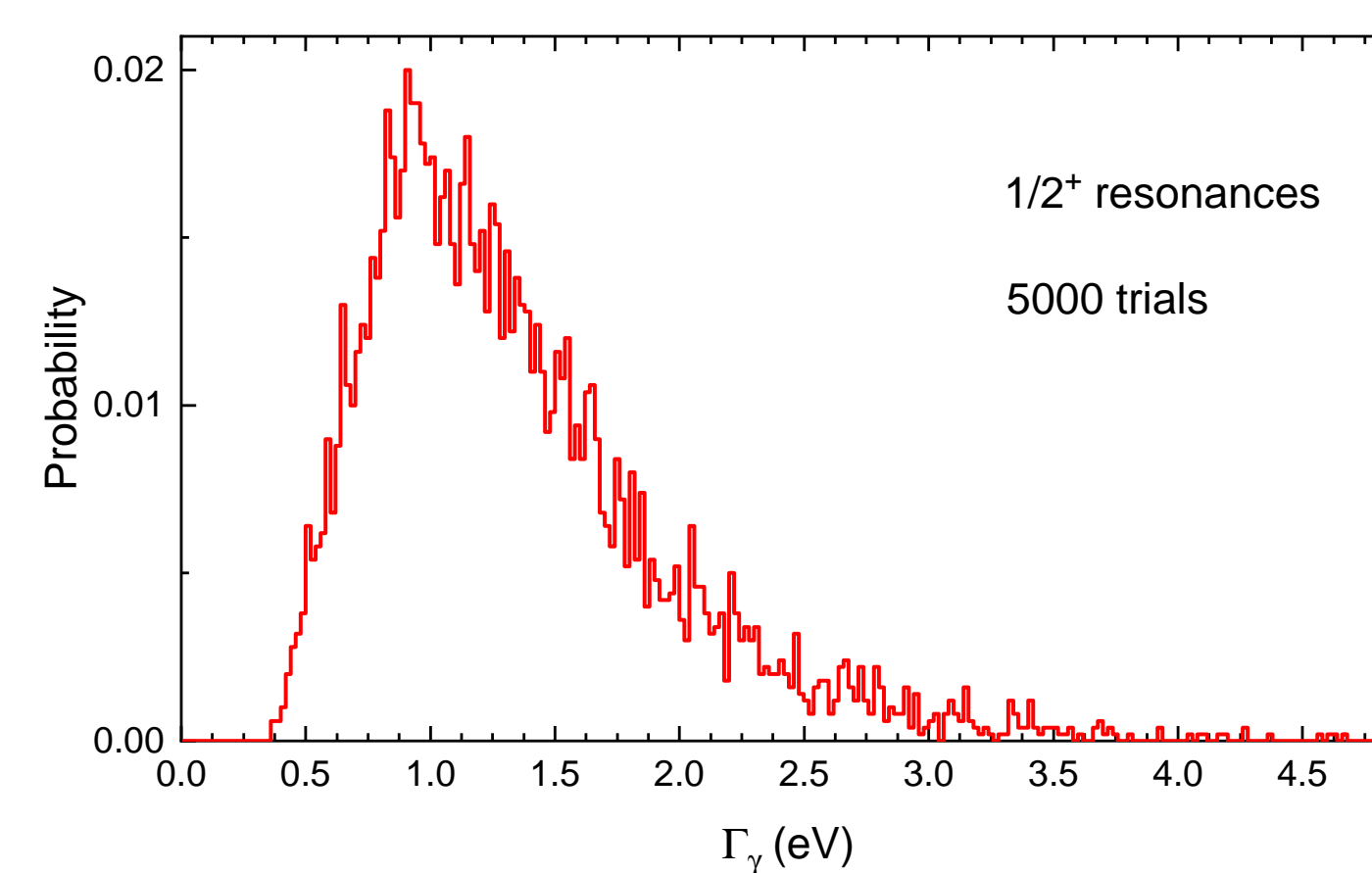


Figure 7: Distribution of total radiative widths as simulated by DICEBOX for  $1/2^+$  level at  $S_n$  in  $^{57}\text{Fe}$ . Only  $P_{\gamma}$  of primary transitions were varied in the simulations.

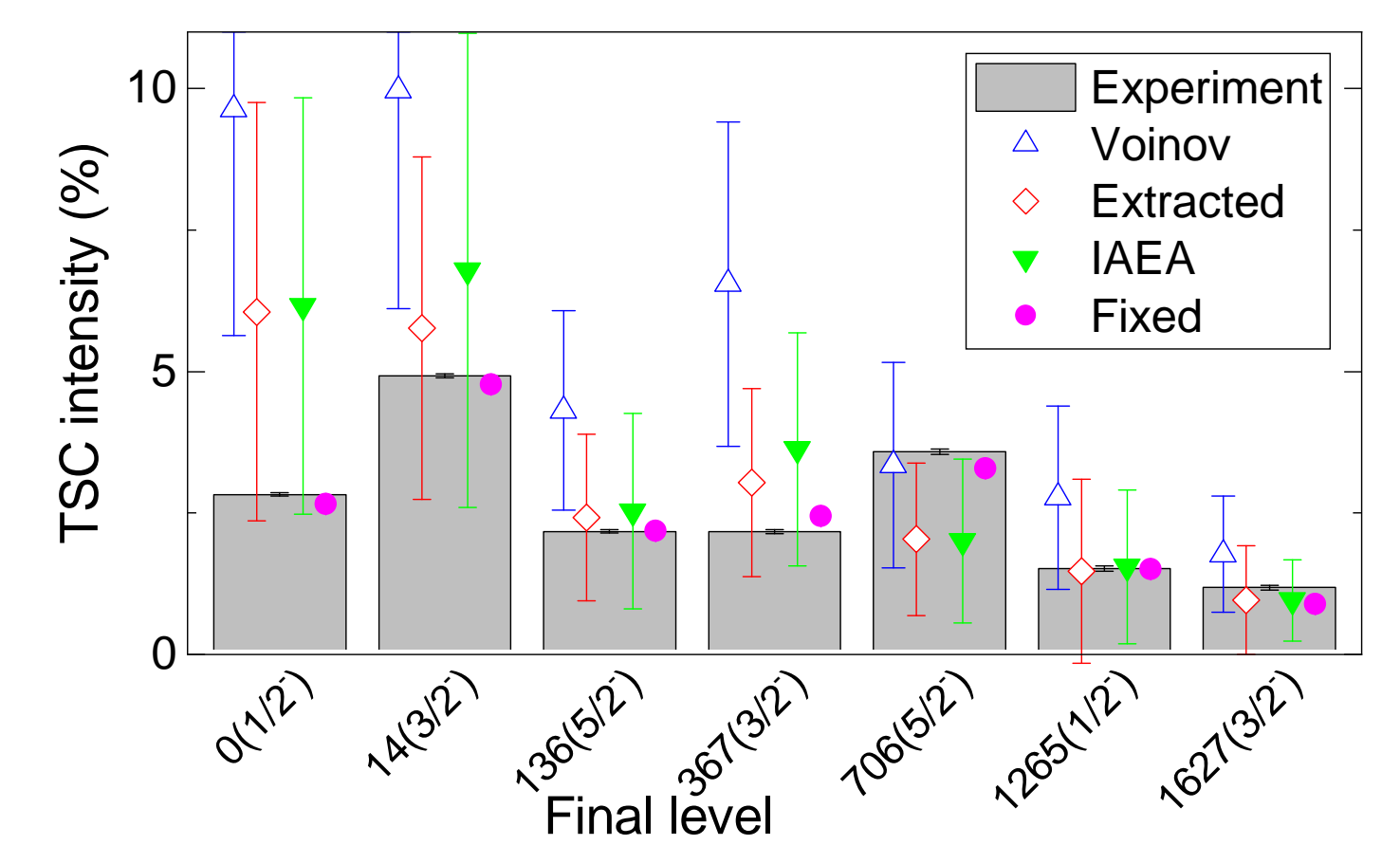


Figure 8: Comparison of experimental and simulated TSC intensity integrated over 4 MeV central interval. The “IAEA” PSFs were taken from S. Goriely *et al.*, Eur. Phys. J. A 55 (2019) 172.

## Results

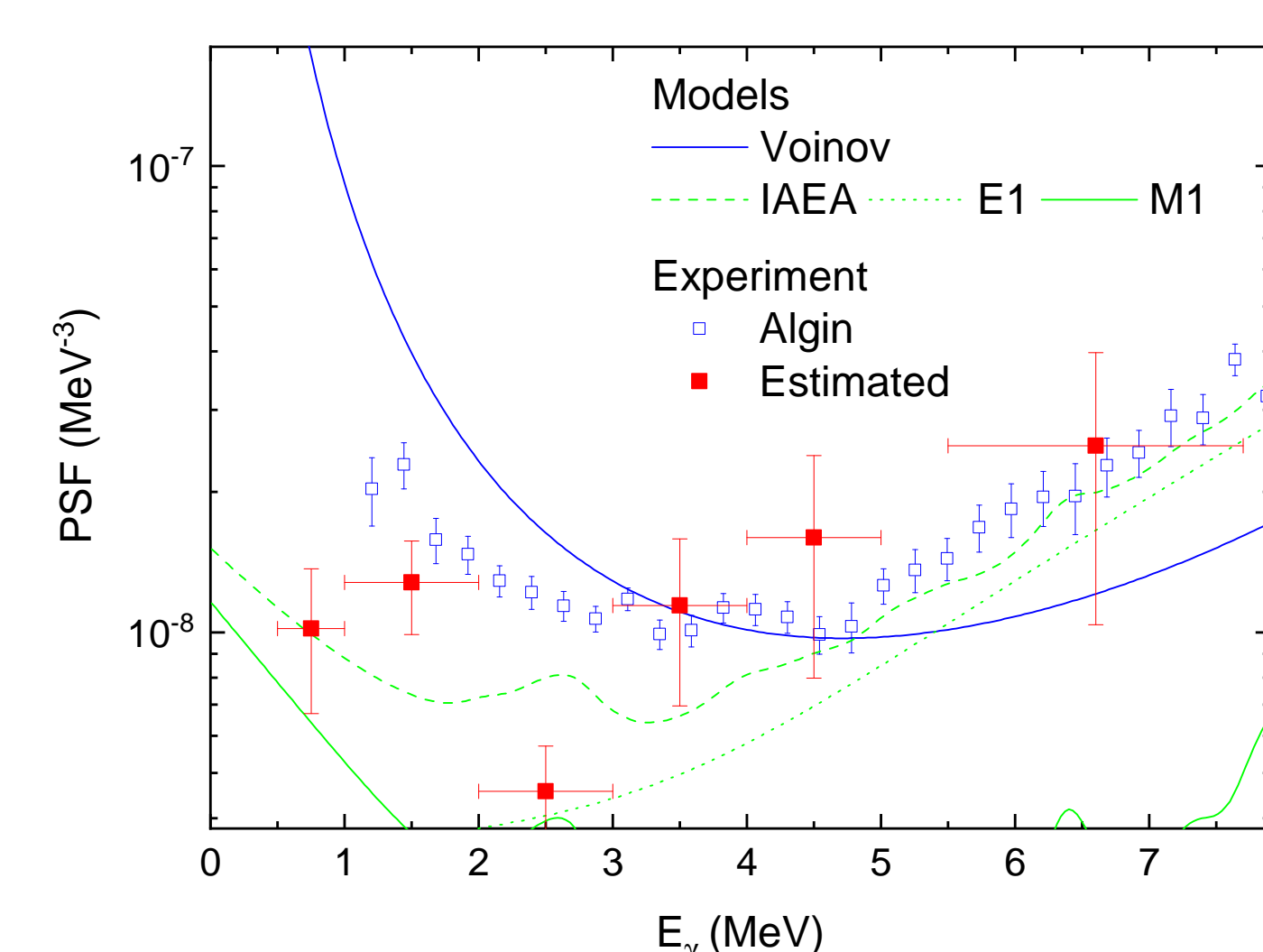


Figure 9: Selected PSF data and models.

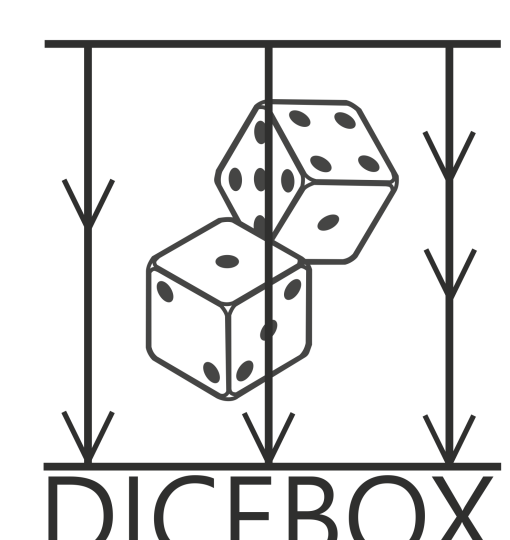
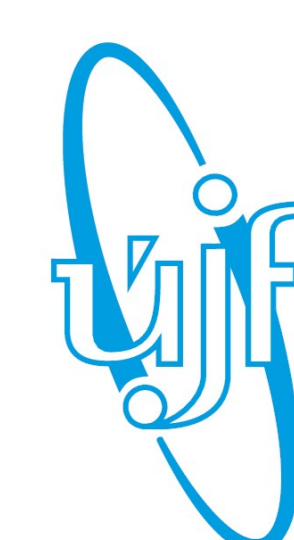
- PSF was extracted from probabilities of primary transitions
- sizable uncertainties of at best 20%, dominated by P-T fluctuations
- a dip at 2-3 MeV observed
- no clear sign of strong LEE down to 0.5 MeV
- $\Rightarrow$  data in disagreement with PSF by Voinov *et al.* [1]
- other models can not be ruled out due to uncertainties

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