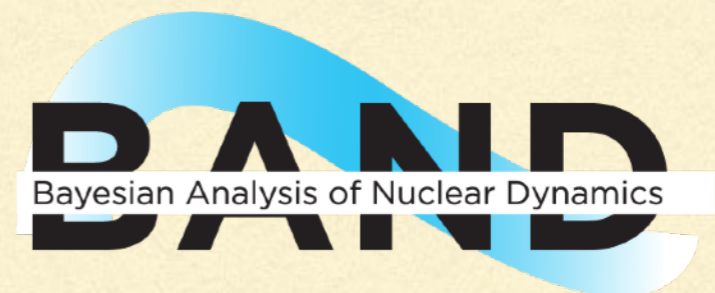

Uncertainty quantification in R-matrix analyses using Bayesian methods

Daniel Phillips
Ohio University



OHIO
UNIVERSITY

with Daniel Odell, Carl Brune,
James deBoer, and Som Paneru



**RESEARCH SUPPORTED BY THE DOE OFFICE OF SCIENCE, THE
SSAP, AND THE NSF OAC**

The Bayes-ics

Thomas Bayes (1701?-1761)



<http://www.bayesian-inference.com>

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Posterior



Model evidence

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Posterior

Model evidence

Typically evaluated by MCMC sampling

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Likelihood Prior
↓ ↓
↑ ↑
Posterior Model evidence

Typically evaluated by MCMC sampling

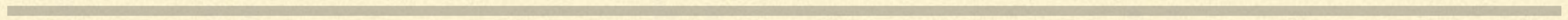
Marginalization: $\text{pr}(x|\text{data}, I) = \int dy \text{pr}(x, y|\text{data}, I)$

Allows us to integrate out “nuisance” parameters, e.g., those associated with systematic uncertainties

The benefits of Bayesian parameter estimation

The benefits of Bayesian parameter estimation

- Straightforward to introduce additional nuisance parameters to model experimental imperfections. Marginalizing over them includes impact of those imperfections on parameters and evaluated quantities



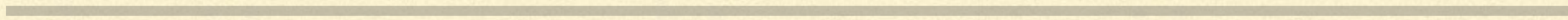
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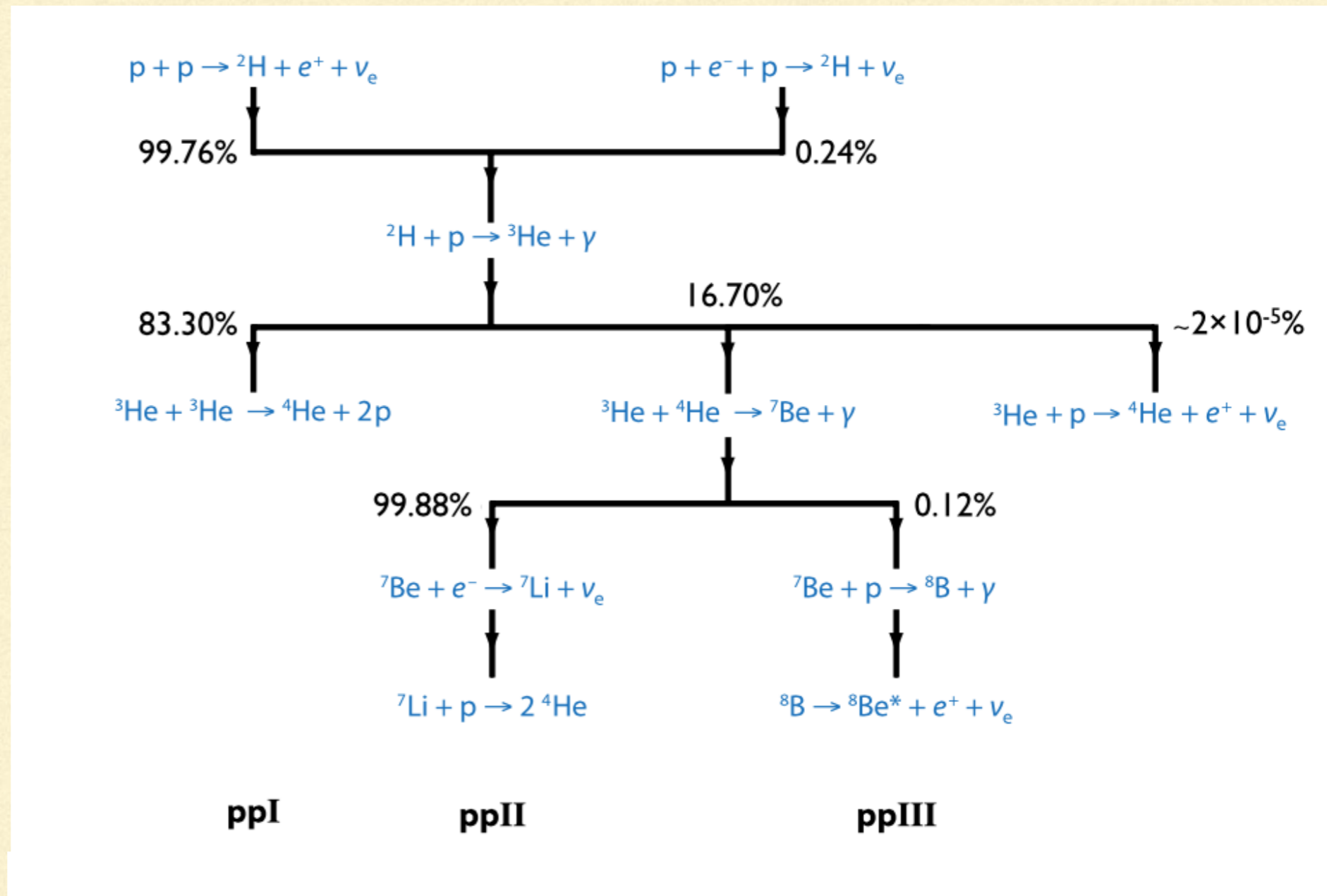
Outline

- What is all this fuss about Bayesian methods? Why should I care?
 - Bayesian R-matrix analysis of ${}^3\text{He} + {}^4\text{He} \rightarrow {}^7\text{Be} + \gamma$ and ${}^3\text{He} + {}^4\text{He}$ elastic scattering
 - Set up Odell, Brune, DP, deBoer, Paneru, *Frontiers in Physics* (2022)
Paneru, Brune Connolly, Odell, Poudel, DP, et al. *Phys. Rev. C* (to appear)
 - Experimental imperfections
 - Why the full posterior?
 - Error propagation
 - Bayesian R-matrix analysis of dt fusion Odell, Brune, DP. *Phys. Rev. C* (2022)
 - Summary and Future Work
-

Why is ${}^3\text{He}({}^4\text{He},\gamma)$ important?

Adelberger et al., Rev. Mod. Phys. 83, 195 (2011)

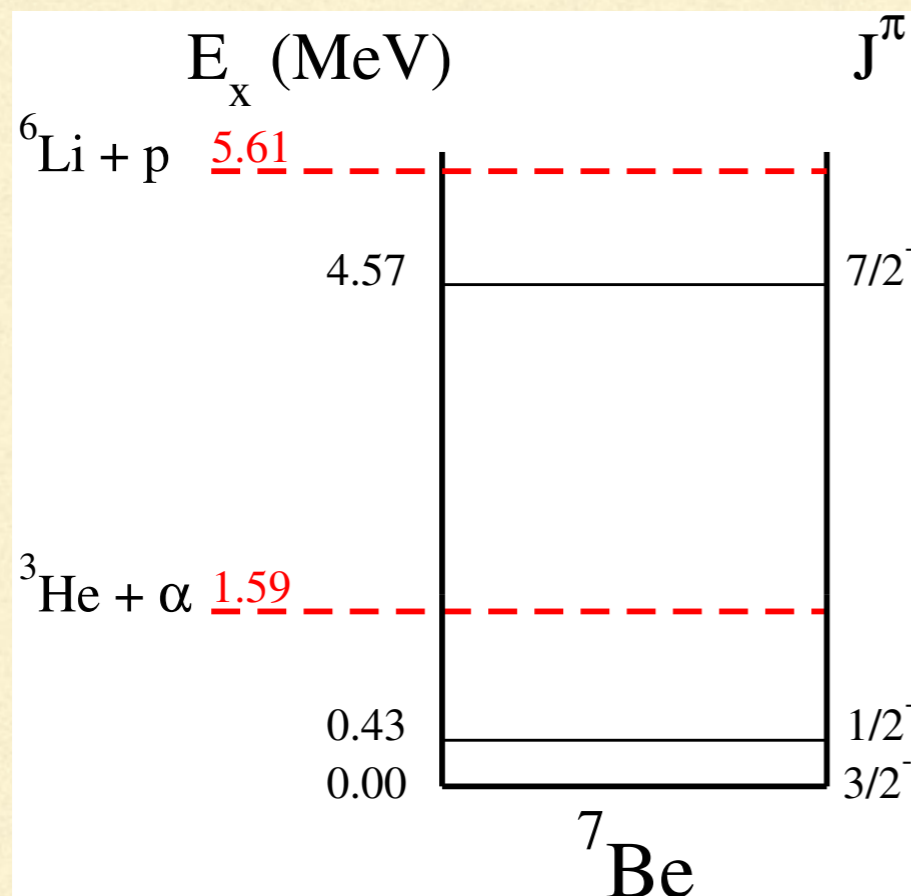
- Accurate knowledge of ${}^3\text{He}({}^4\text{He},\gamma)$ needed to reliably predict amount of ${}^7\text{Be}$ in the Sun
- Therefore key for prediction of ${}^8\text{B}$ solar neutrino flux
- BBN implications, but I will not discuss those here



Build an R-matrix model

Odell, Brune, DP, deBoer, Paneru, *Frontiers in Physics* (2022)

- Goal: describe scattering and capture data up to the $p^6\text{Li}$ threshold
- $3/2^-$ and $1/2^-$ bound states with prior ranges for ANCs from 1 to 5 MeV



Background & resonance levels

| | E (MeV) | Γ_a (MeV) |
|---------|---------|------------------|
| $1/2^-$ | 21.6 | [-200,200] |
| $3/2^-$ | 21.6 | [-100,100] |
| $5/2^-$ | 7 | [0,100] |
| $7/2^-$ | [2,10] | [0,10] |
| $1/2^+$ | 14 | [0,100] |
| $3/2^+$ | 12 | [0,100] |
| $5/2^+$ | 12 | [0,100] |

Pick data sets

- 88 S-factor data
 - Seattle (S)
 - Weizmann
 - Luna (L)
 - Erna
 - Notre Dame
 - Atomki
 - Scattering data
 - SONIK*: 45 I from 0.385 to 3.127 MeV
 - Barnard: 646 from 1.49 to 3.27 MeV
- *Paneru et al., arXiv:2211.14641, Phys. Rev. C (in press)

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 - Specify CMEs
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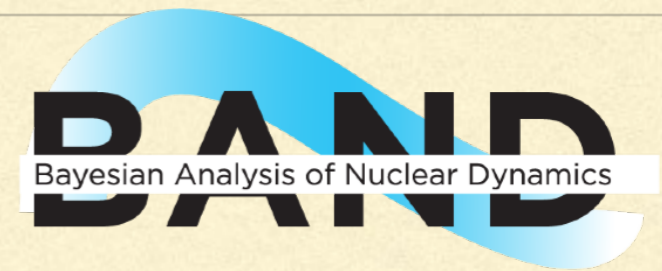
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Two analyses:

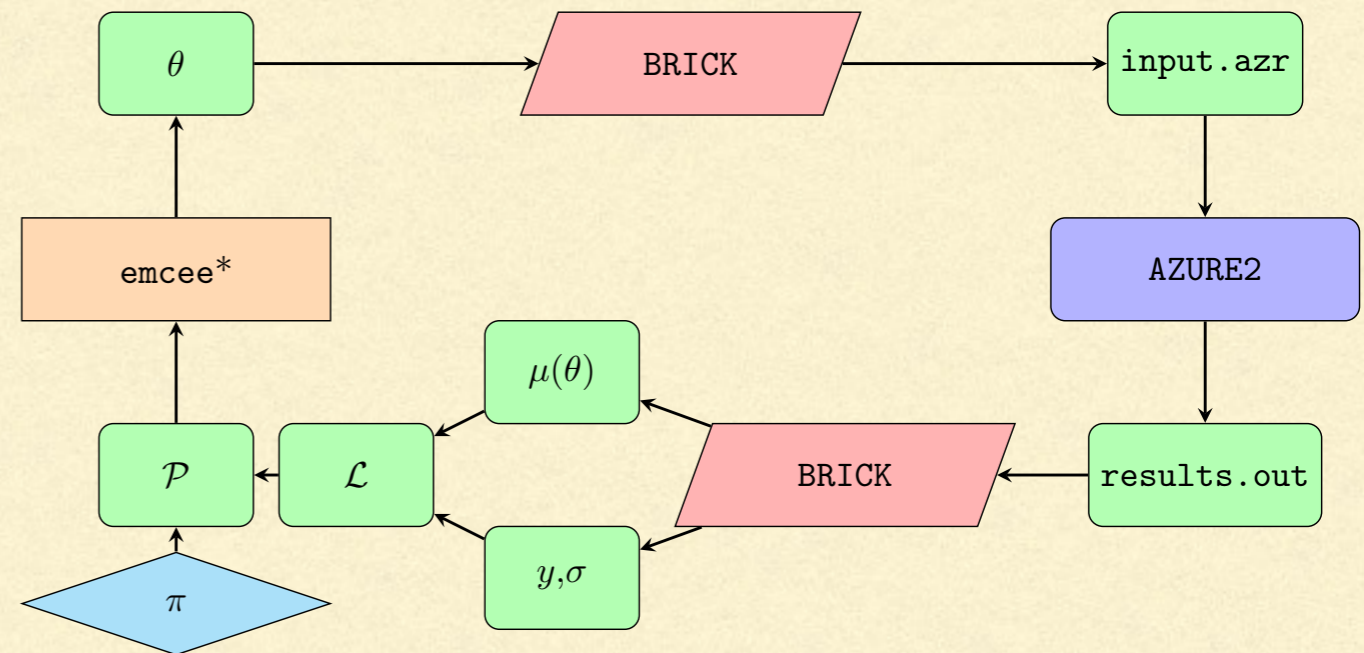
Capture + SONIK

Capture + SONIK + Barnard

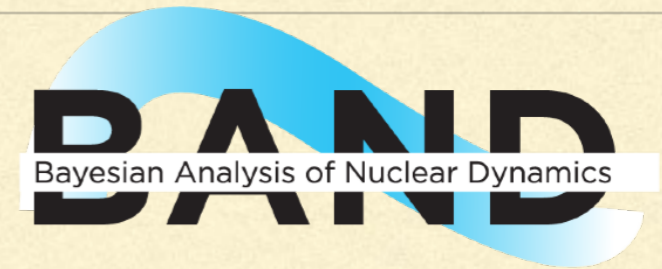
Throw a BRICK at it



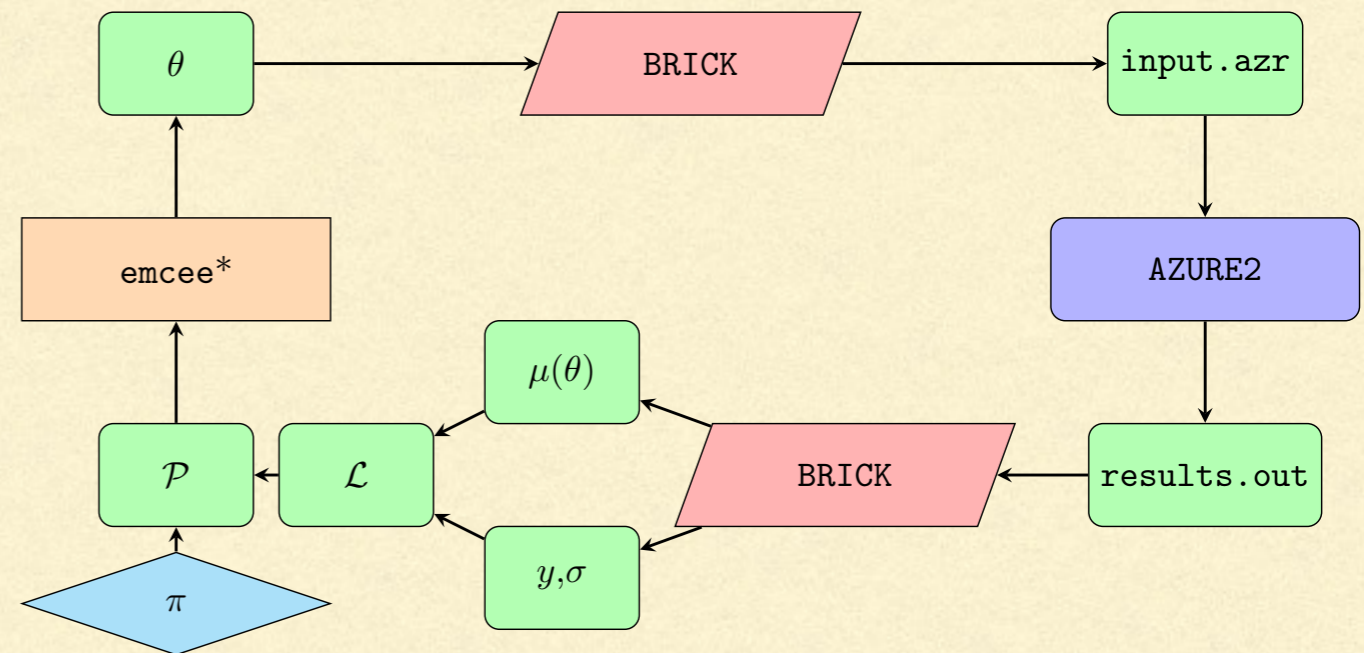
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- BAND Framework v0.2
bandframework.github.io
- AZURE2 must be installed
- User specifies R-matrix model & data set in AZURE2
- Specify priors & likelihood



Throw a BRICK at it

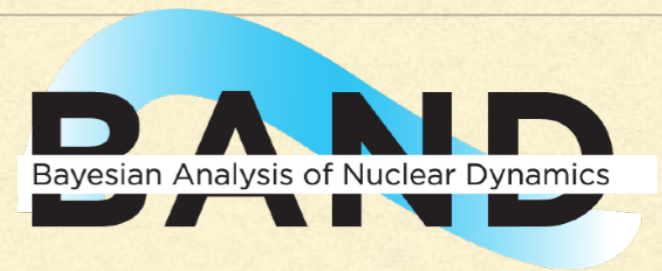


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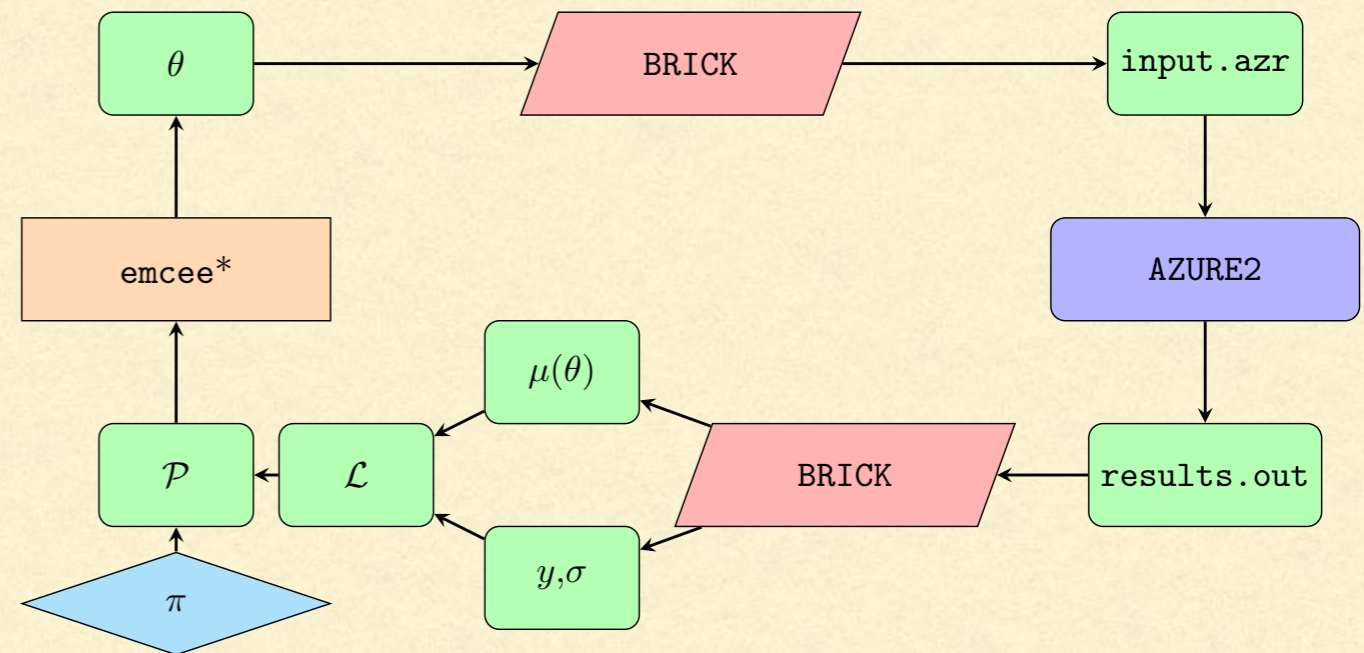
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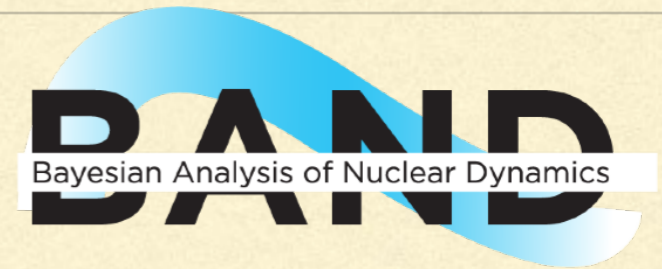
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Data



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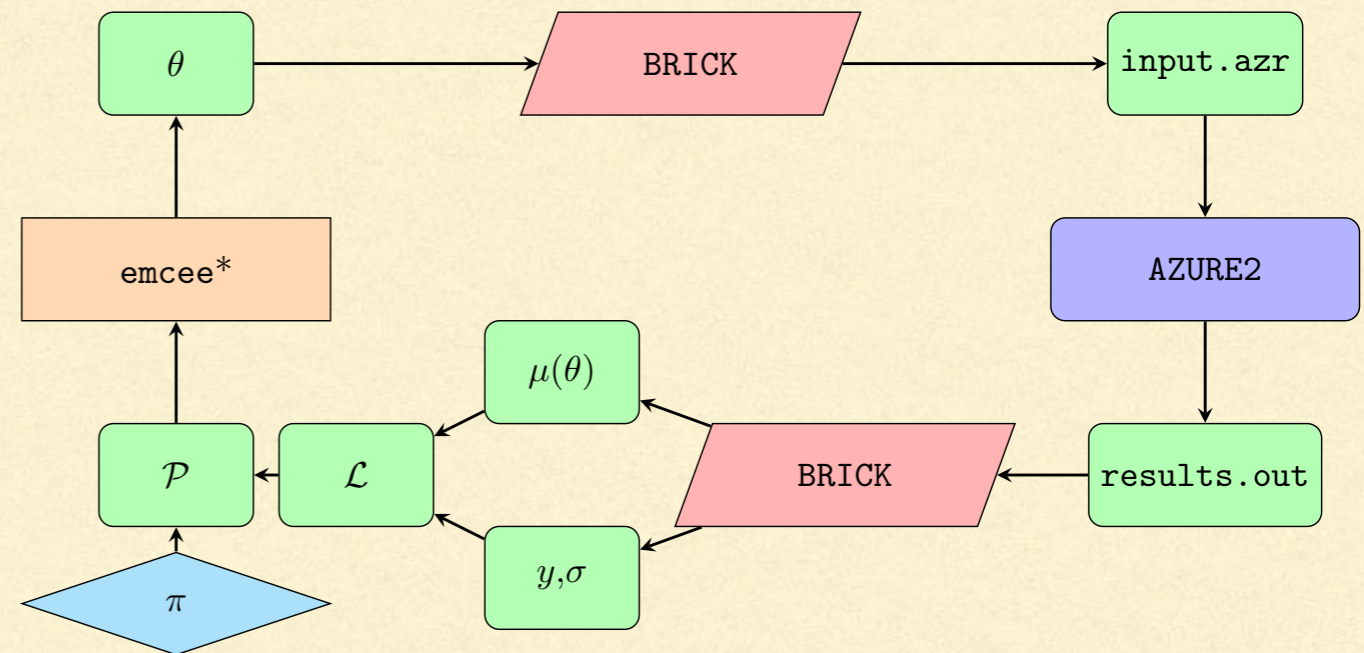
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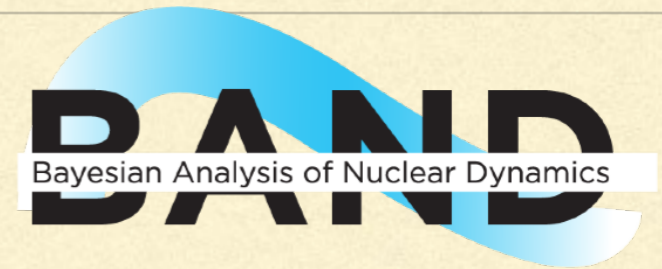
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Data

R-matrix number



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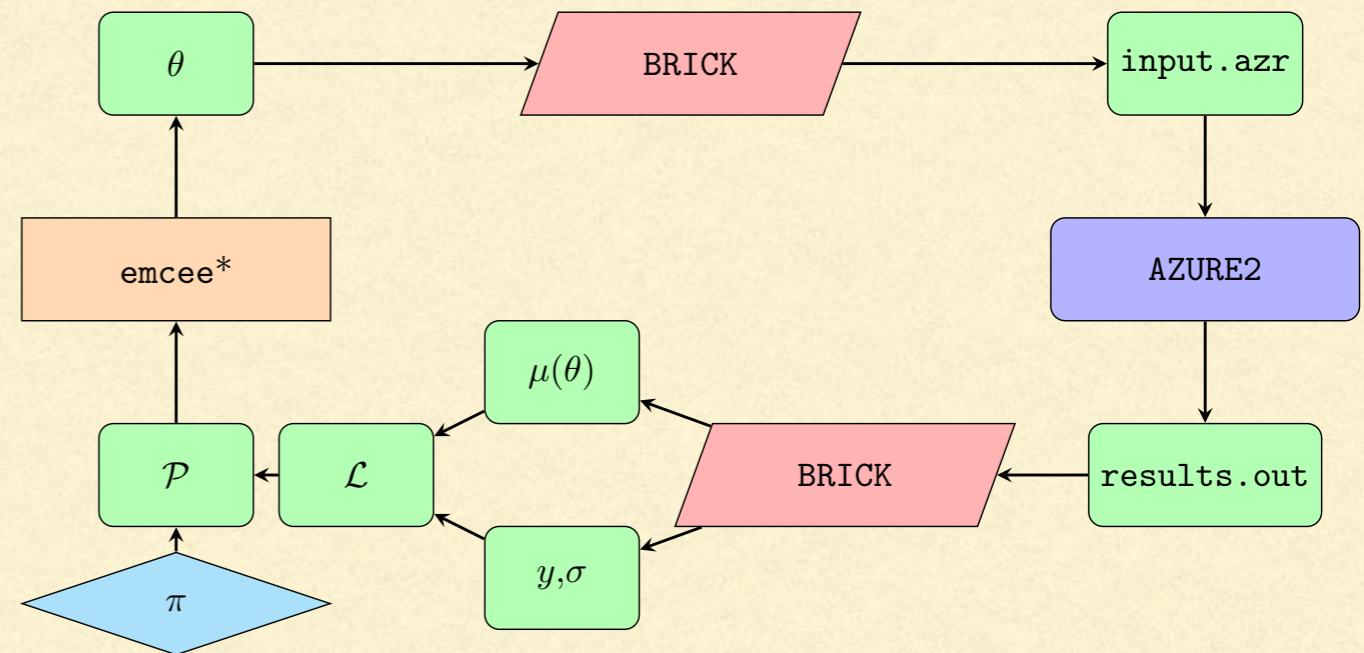
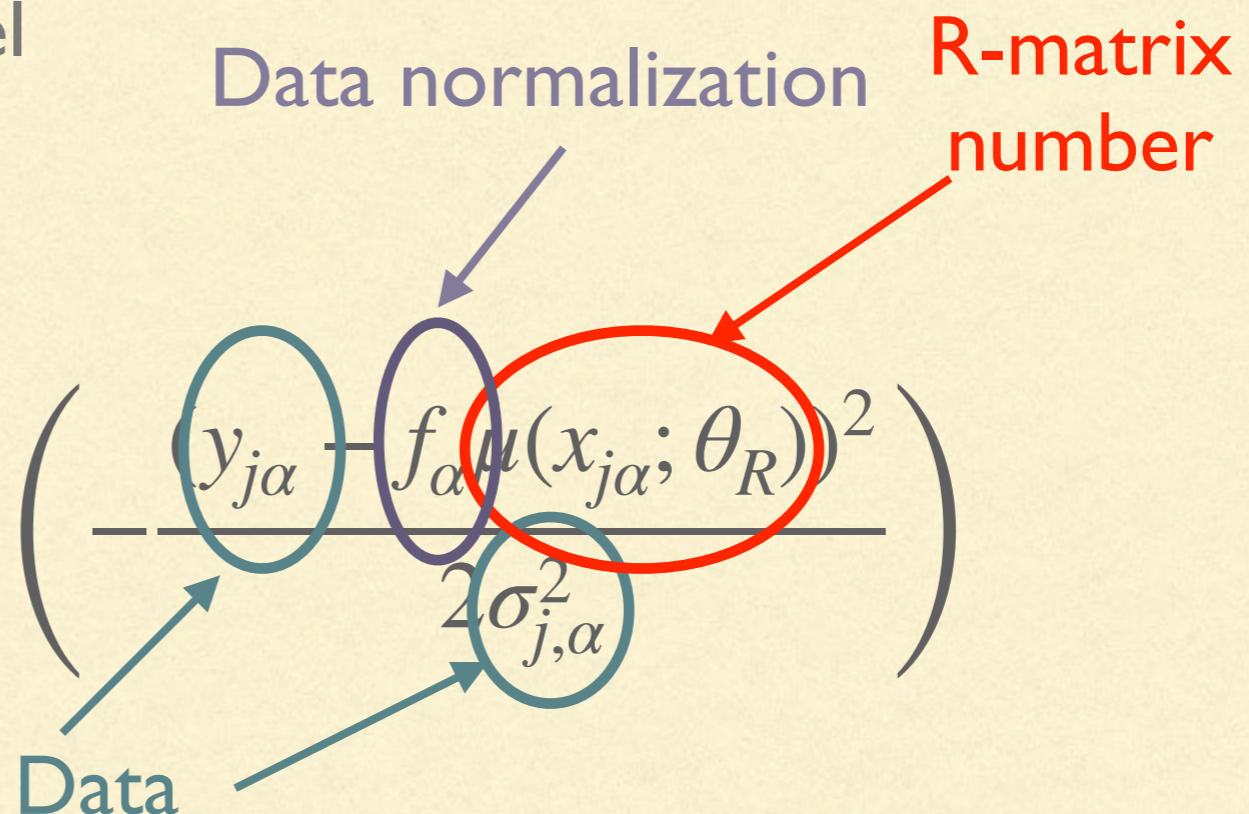
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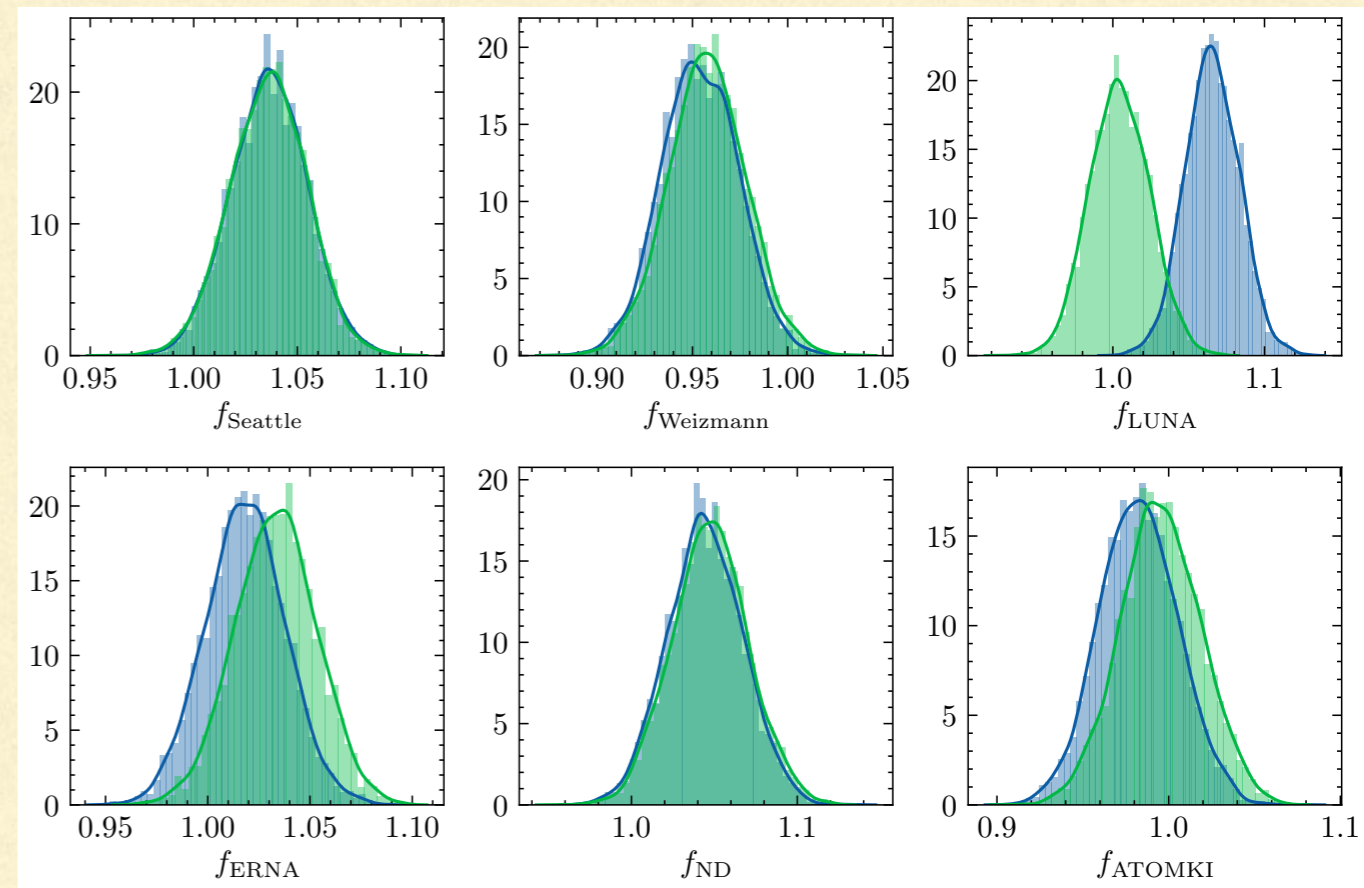
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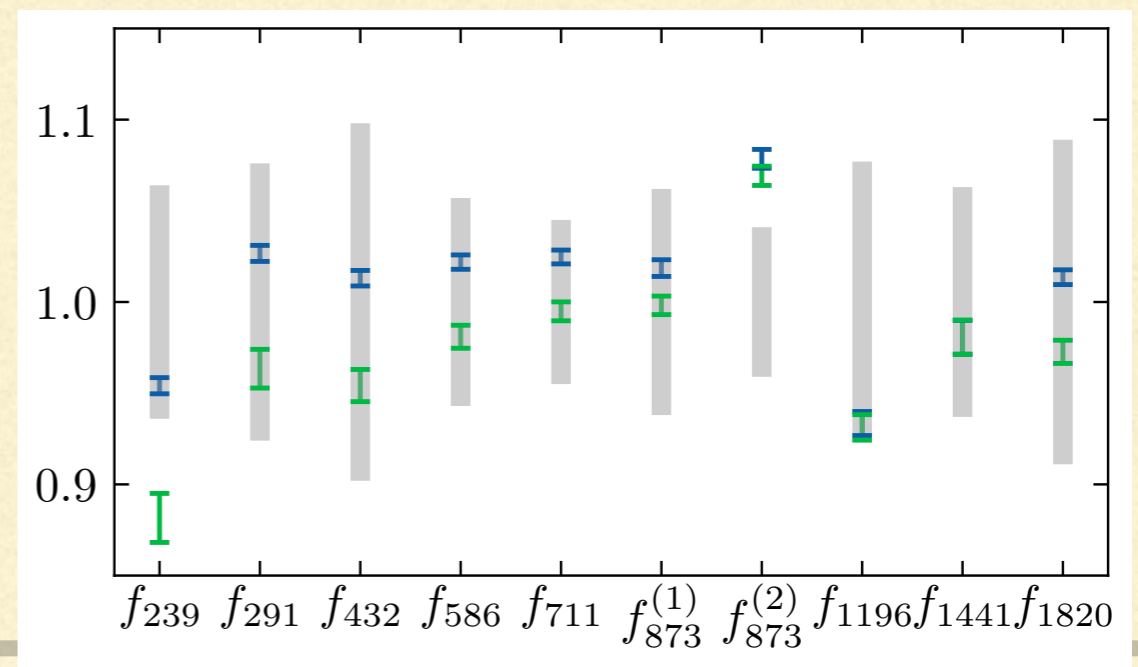
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-

Modeling of normalization uncertainties

- Analysis includes common-mode errors for all data sets, implemented by factor f_α to avoid d'Agostini bias
- For SONIK data set this normalization factor is assigned for each beam energy
- Almost all normalizations come out inside quoted CMEs, all are within $2 \times \text{CME}$, apart from LUNA in CSB analysis
- “Dialogue with the data”



CS
CSB



More sophisticated normalization modeling

Paneru, Brune, Connelly, Odell, ..., DP, et al., PRC (to appear)

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More sophisticated normalization modeling

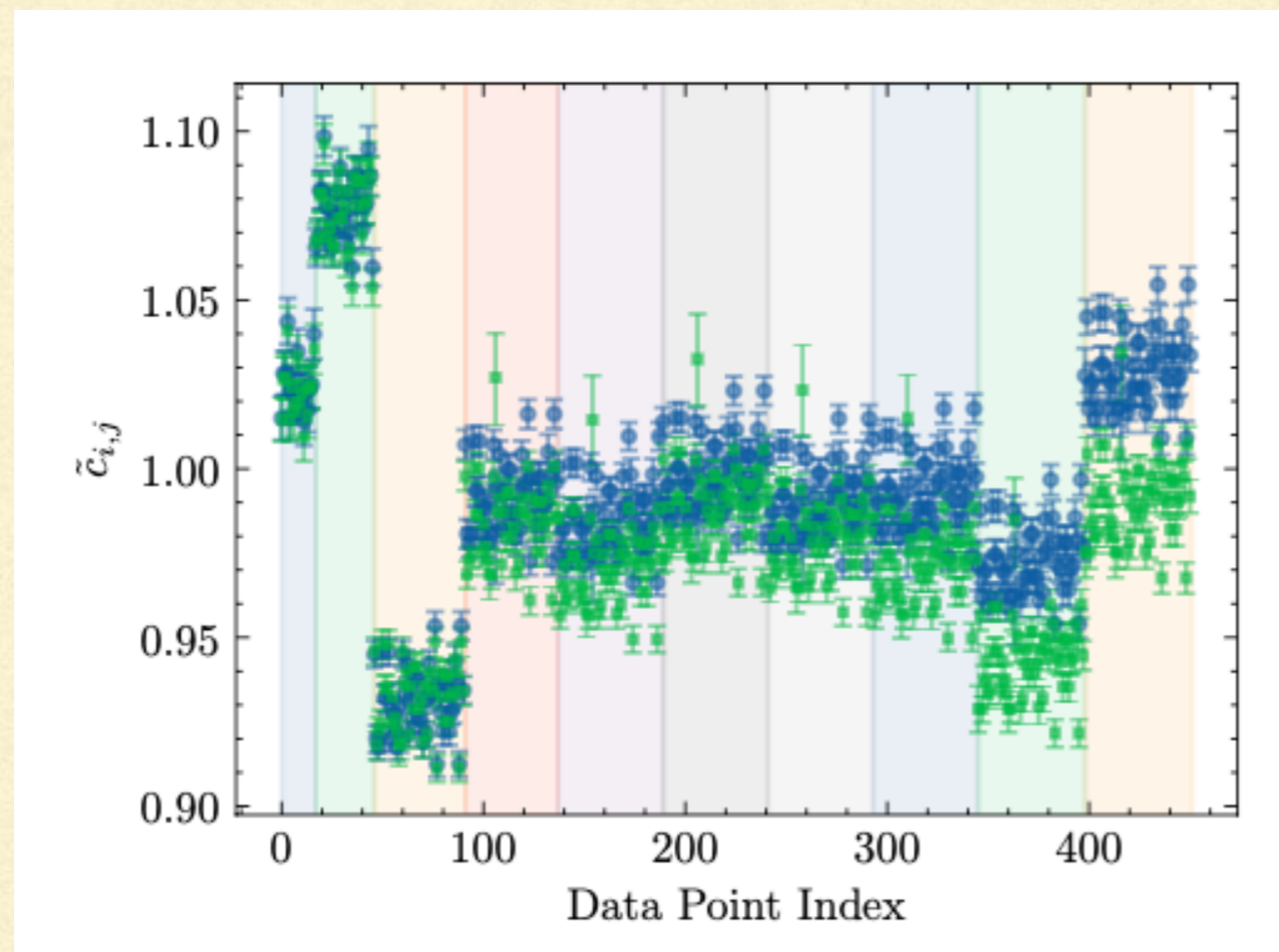
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$$y_{\text{exp}} = f_{\text{SONIK}} f_E f_{\text{det}} y_R + \delta y_{\text{exp}}$$

$$\tilde{c}_{i,j} = f_E f_{\text{det}}$$

Green: R-matrix

Blue: EFT

Beam energy shift

Beam energy shift

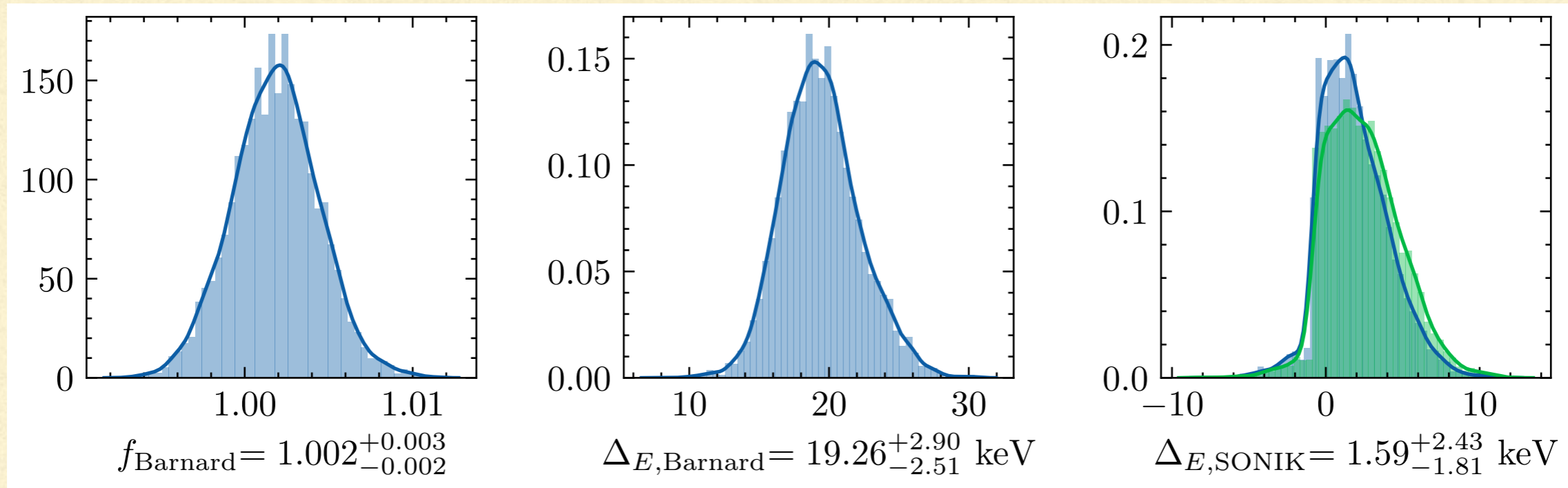
- Shift energy of Barnard data set by a constant to account for possible miscalibration of beam energy: $E \rightarrow E + \Delta$. Prior a Gaussian with standard deviation 40 keV ← information in paper

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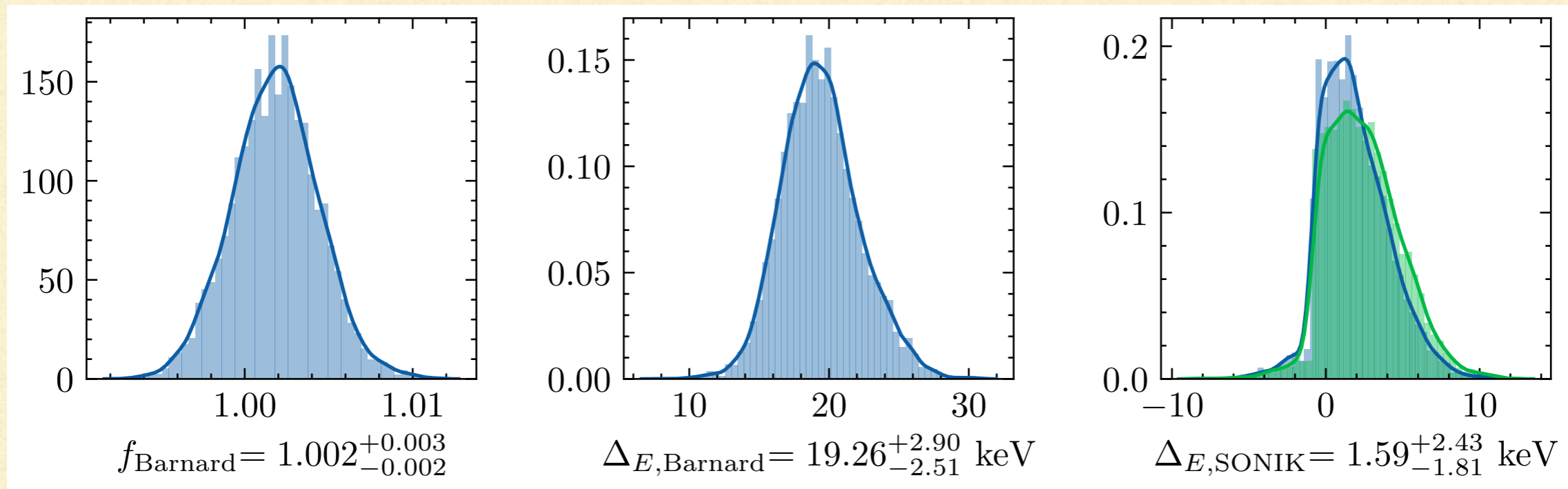
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Beam energy shift

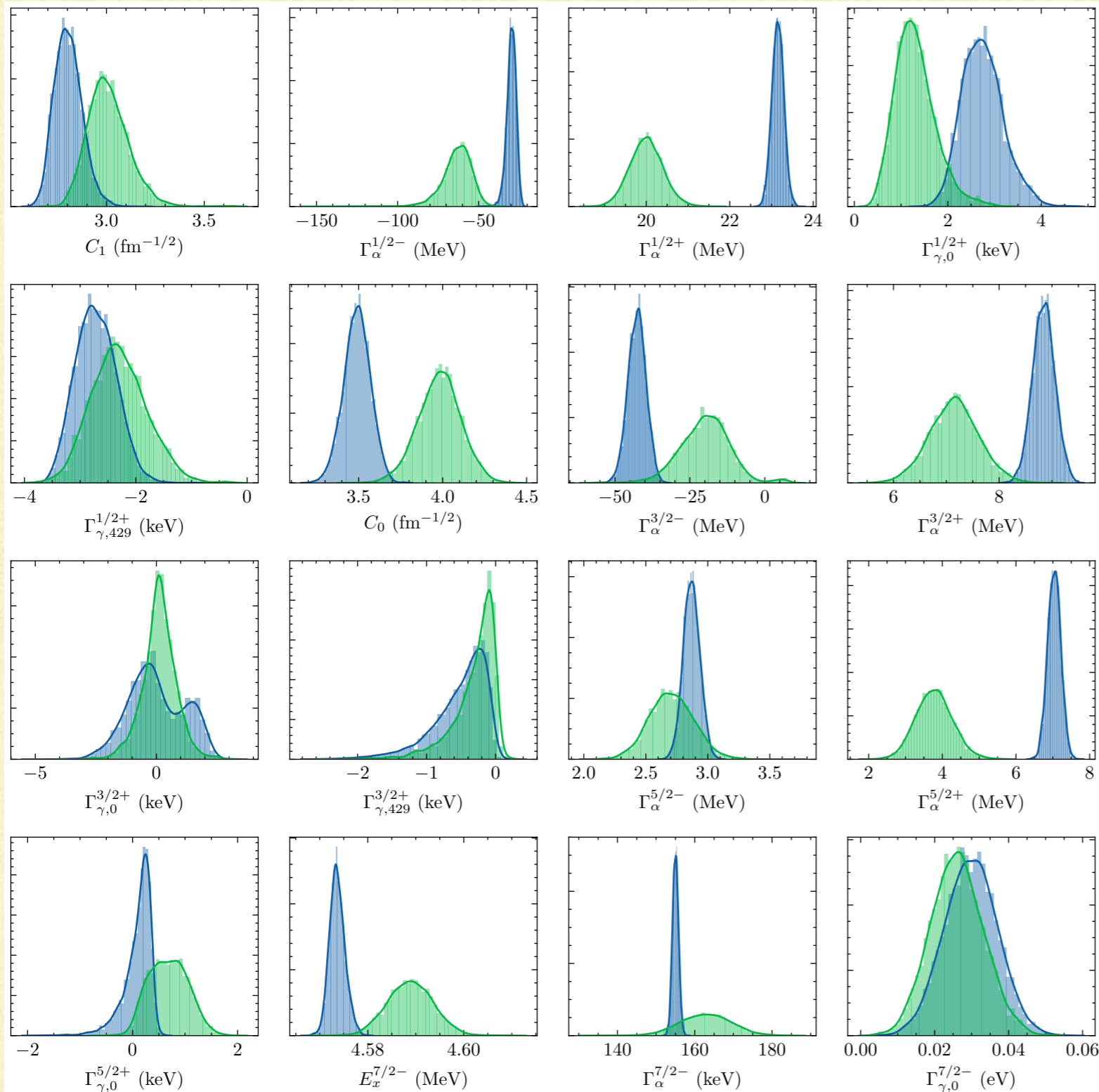
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- No significant change in $\vec{\theta}_{\text{Barnard}}$ due to this though

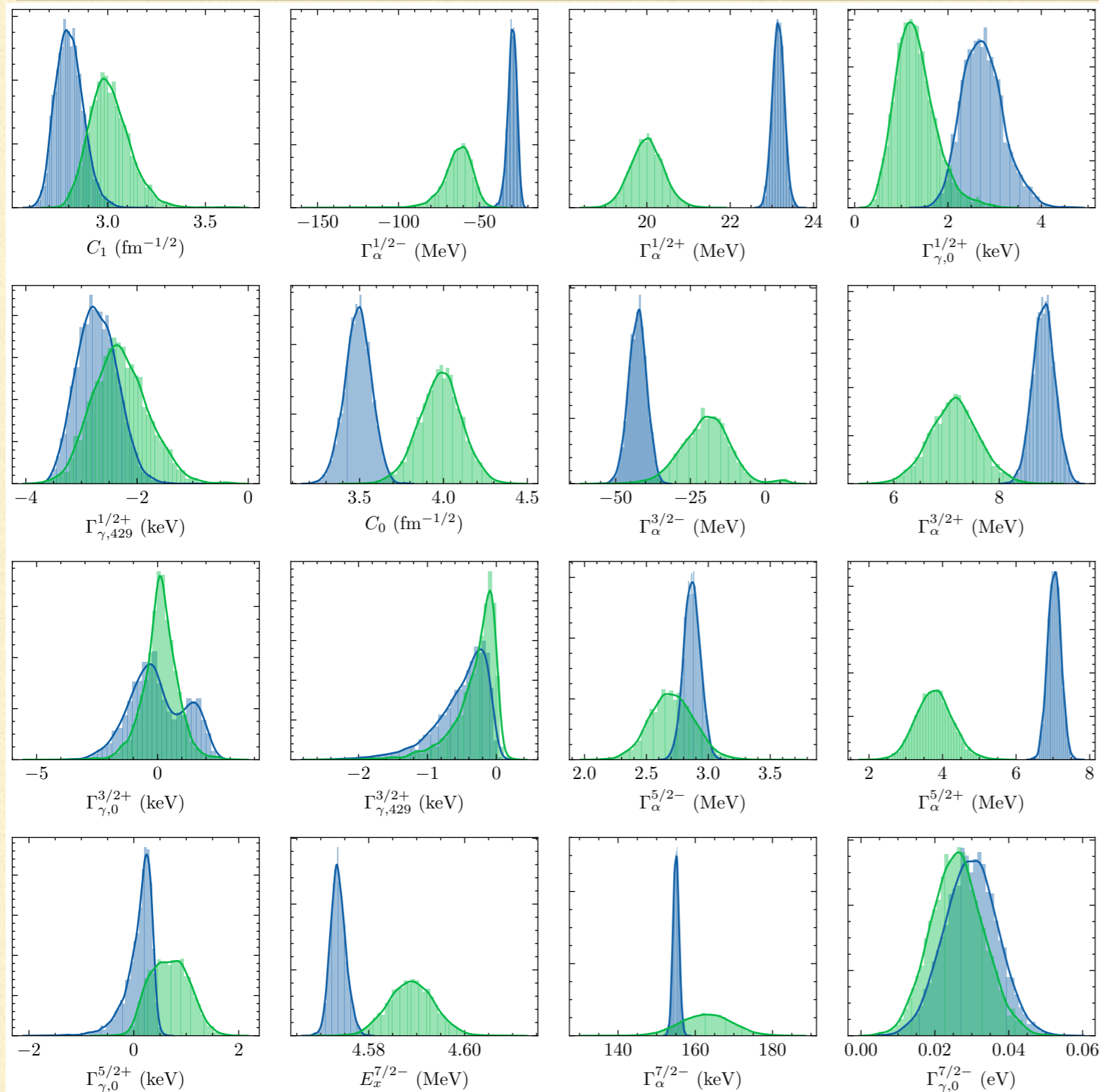
Posteriors for R-matrix parameters



Capture + SONIK

Capture + SONIK + Barnard

Posteriors for R-matrix parameters



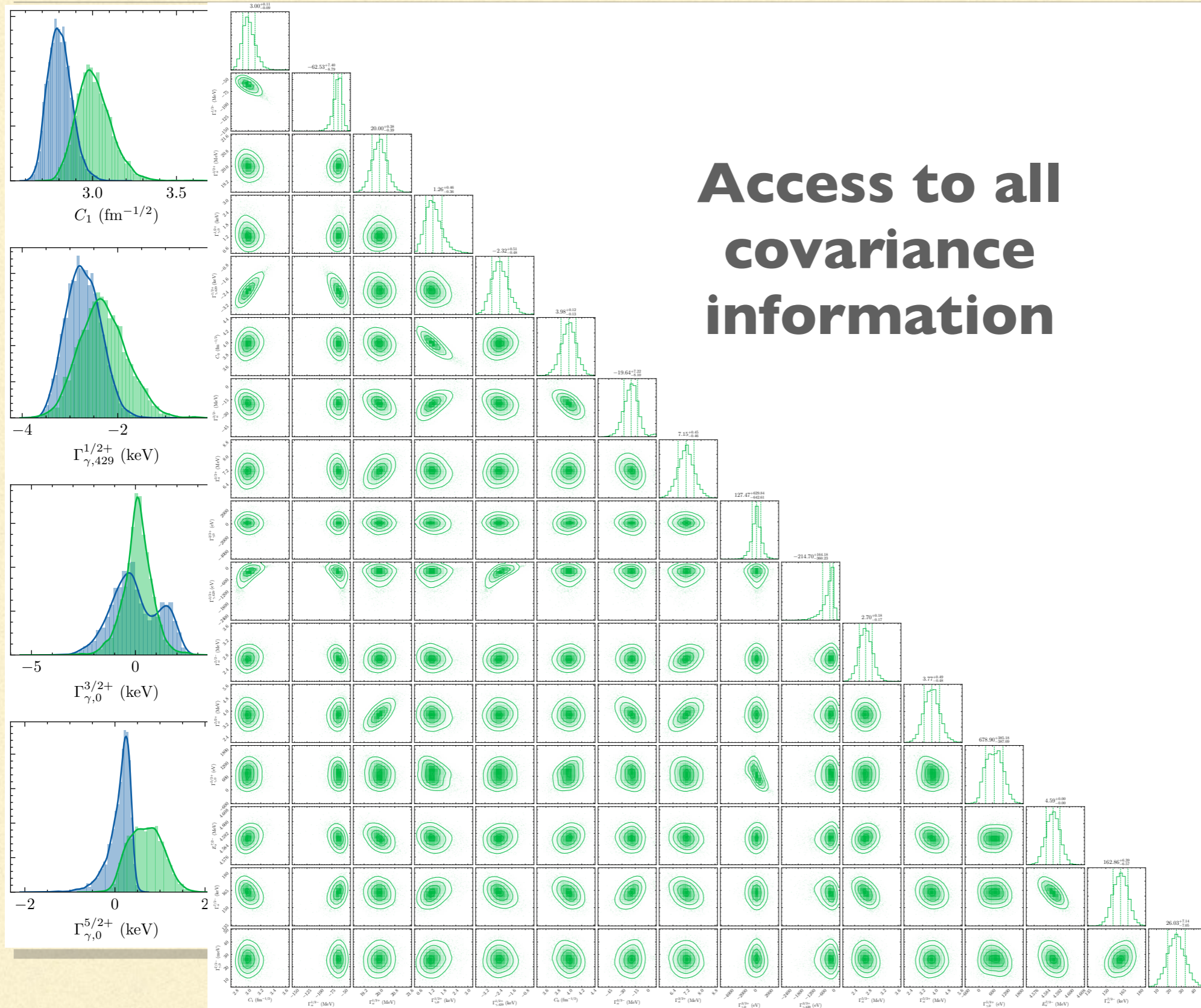
Capture + SONIK

Capture + SONIK + Barnard

Notable points:

- ANCs
- $\Gamma_{\alpha}^{7/2-}$
- Non-Gaussianity

Posteriors for R-matrix parameters



Access to all
covariance
information

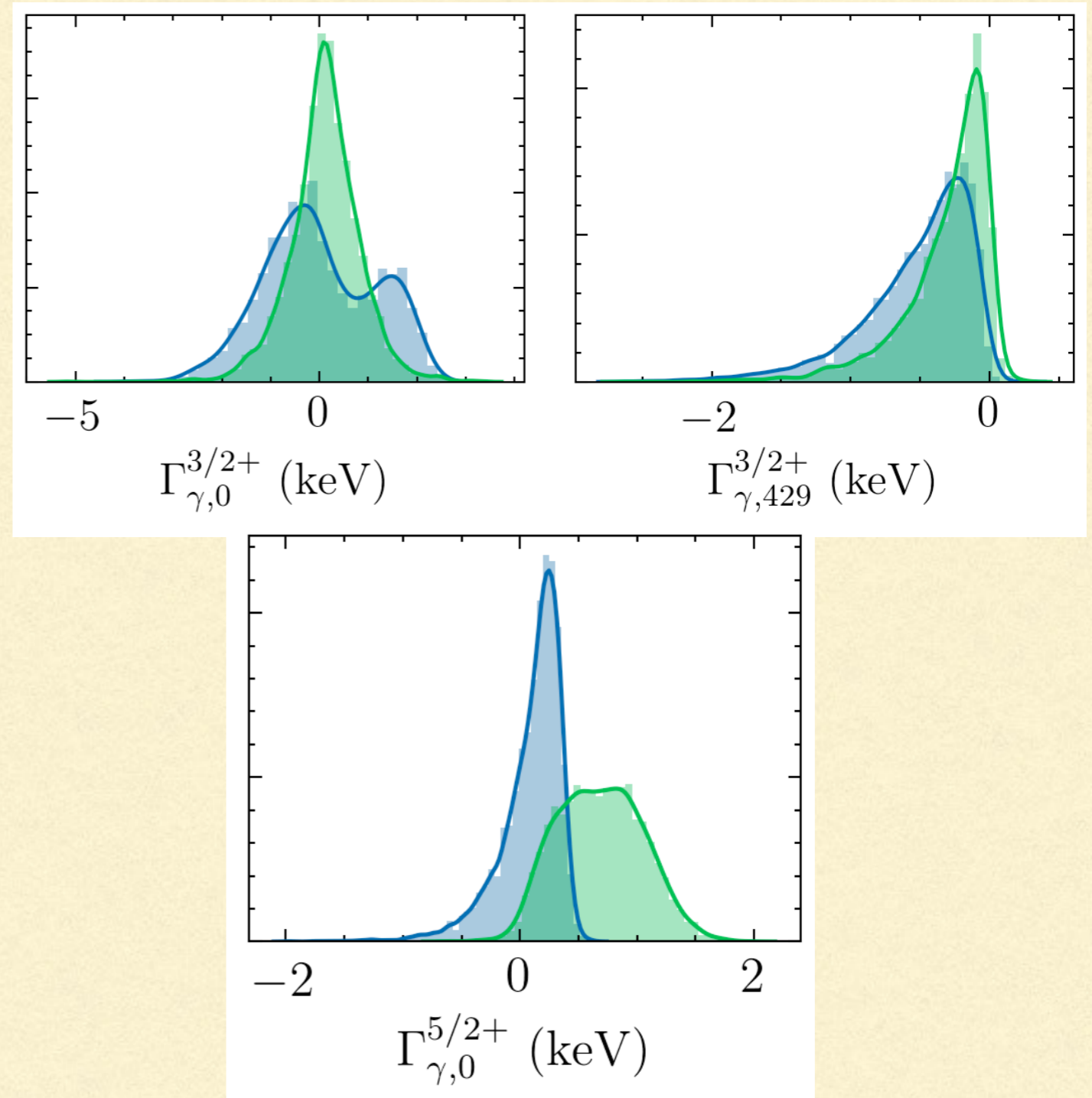
re + SONIK
SONIK + Barnard

ble points:
Cs
-
n-Gaussianity

But why didn't you just use MINUIT?

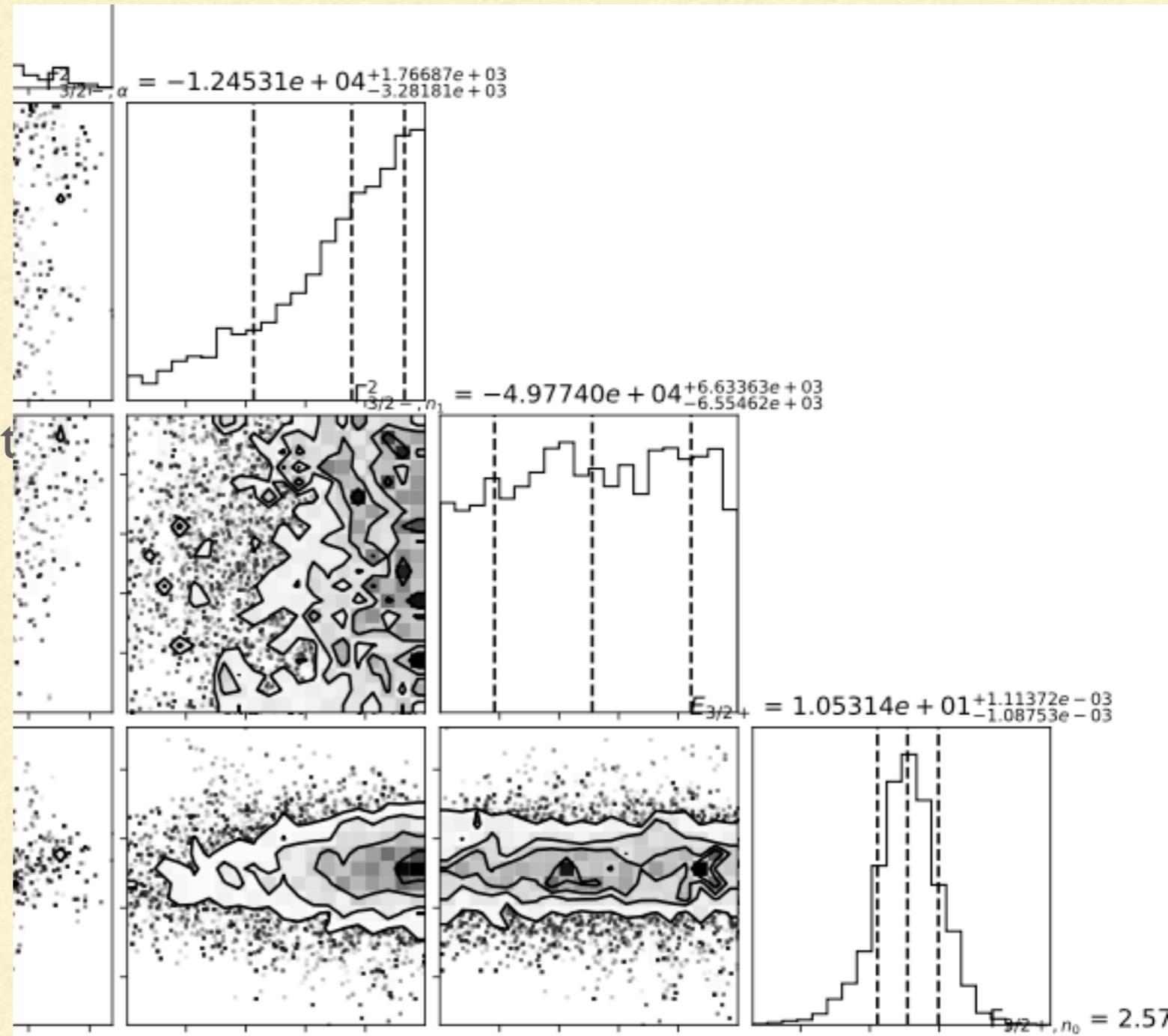
But why didn't you just use MINUIT?

- Diagnose non-Gaussianity



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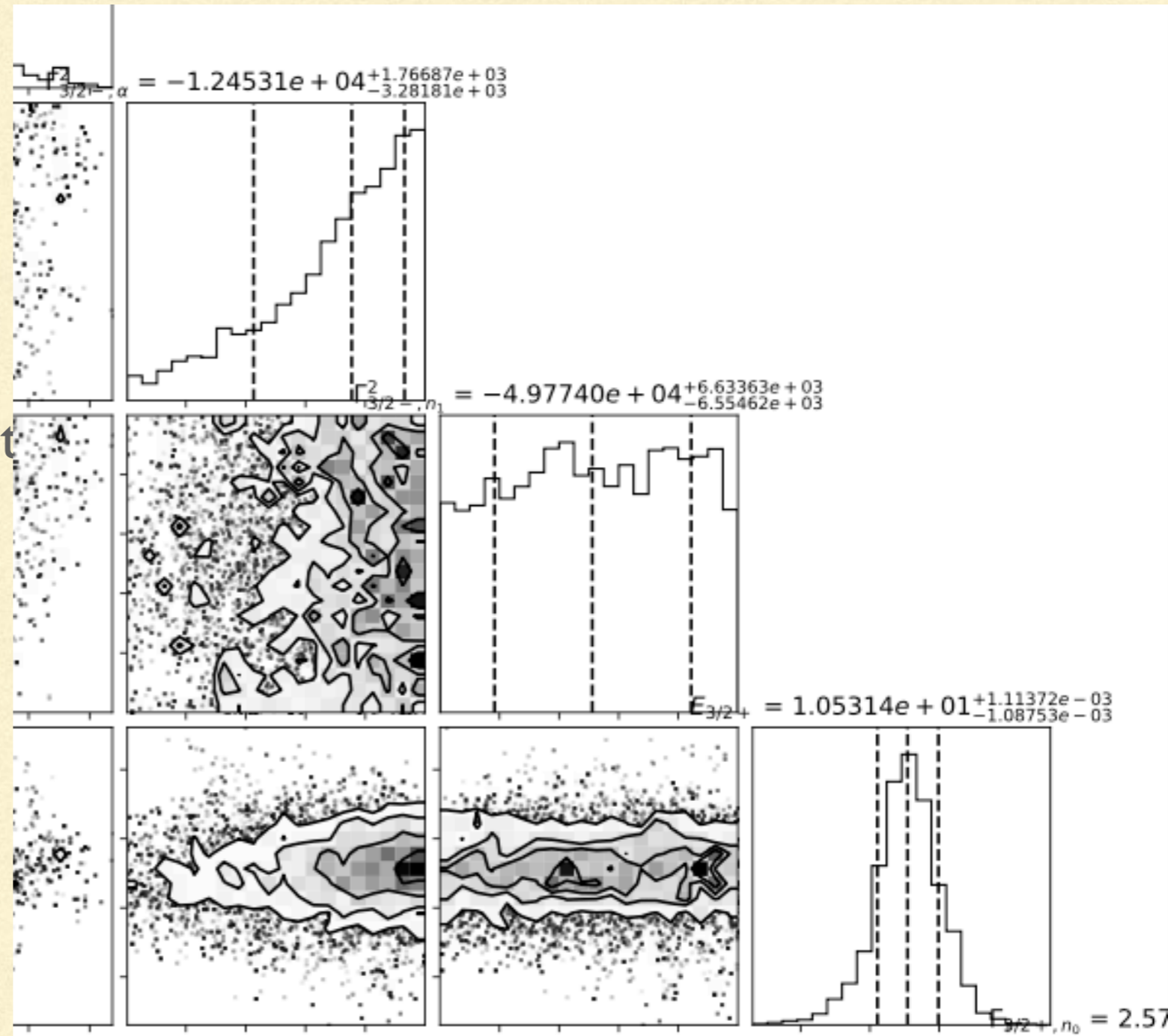
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$^{13}\text{C}(\alpha, n)$, courtesy James deBoer

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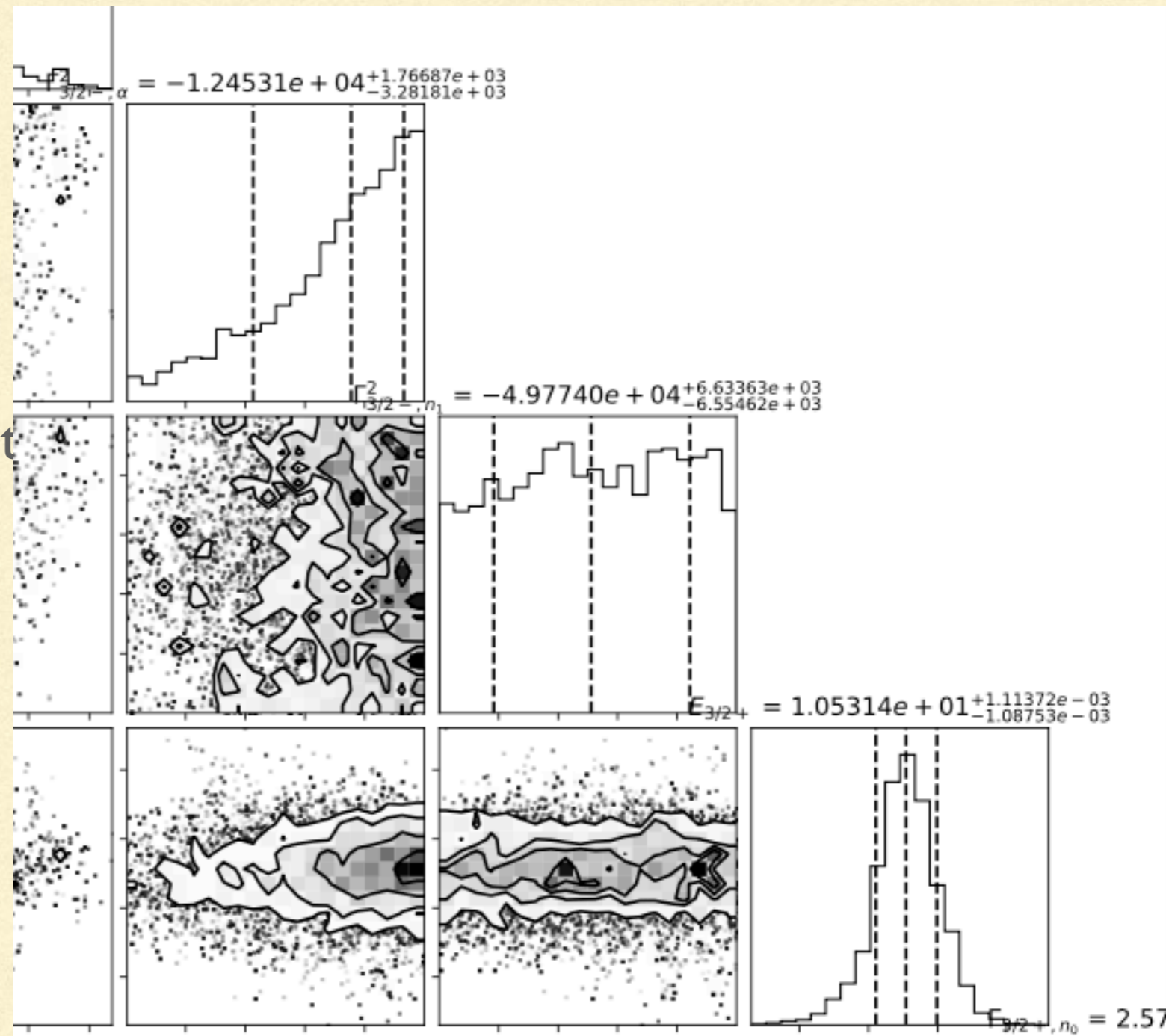
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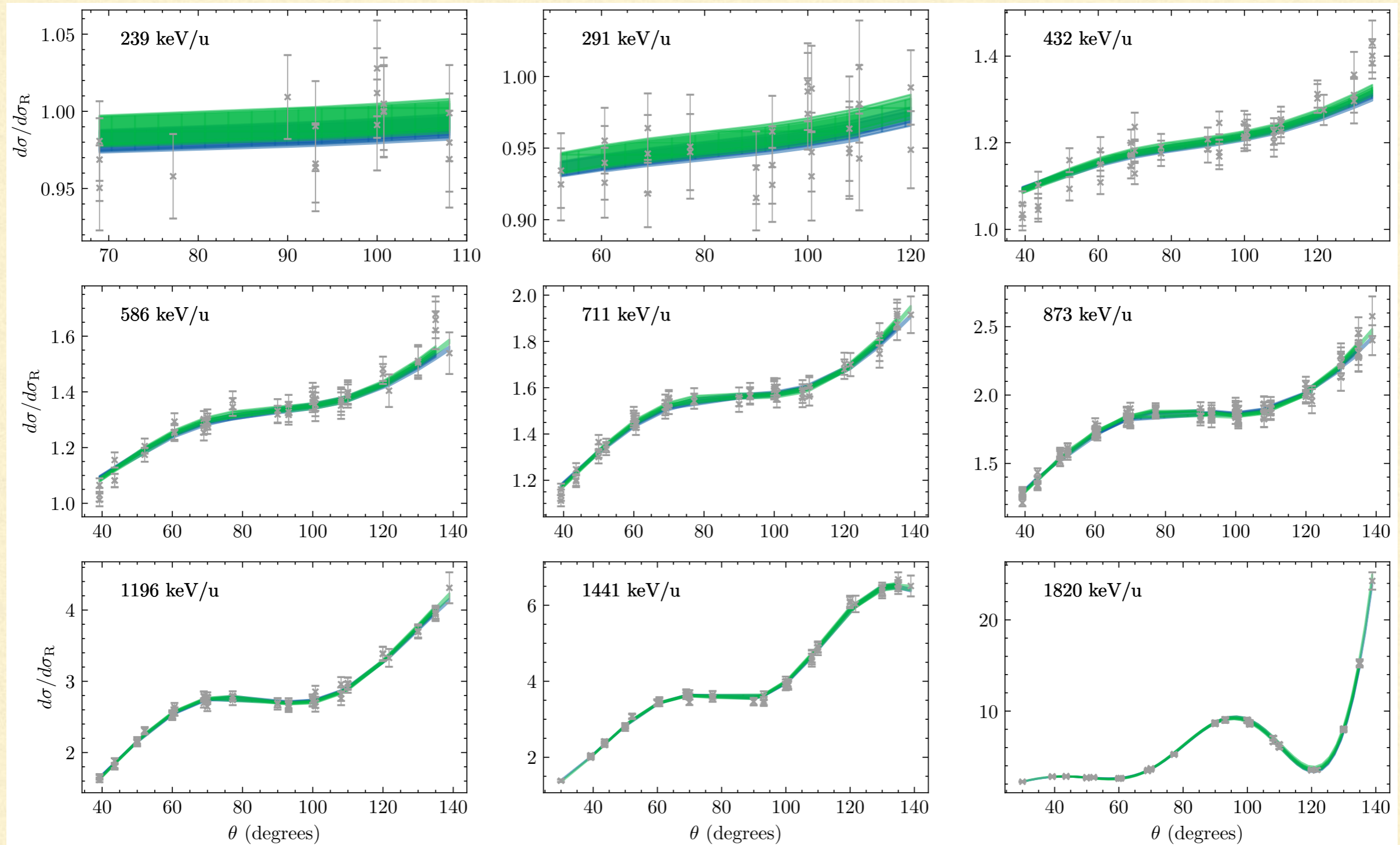
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- (Note that it's also clear when prior is affecting shape of posterior.)
- Also, error propagation....

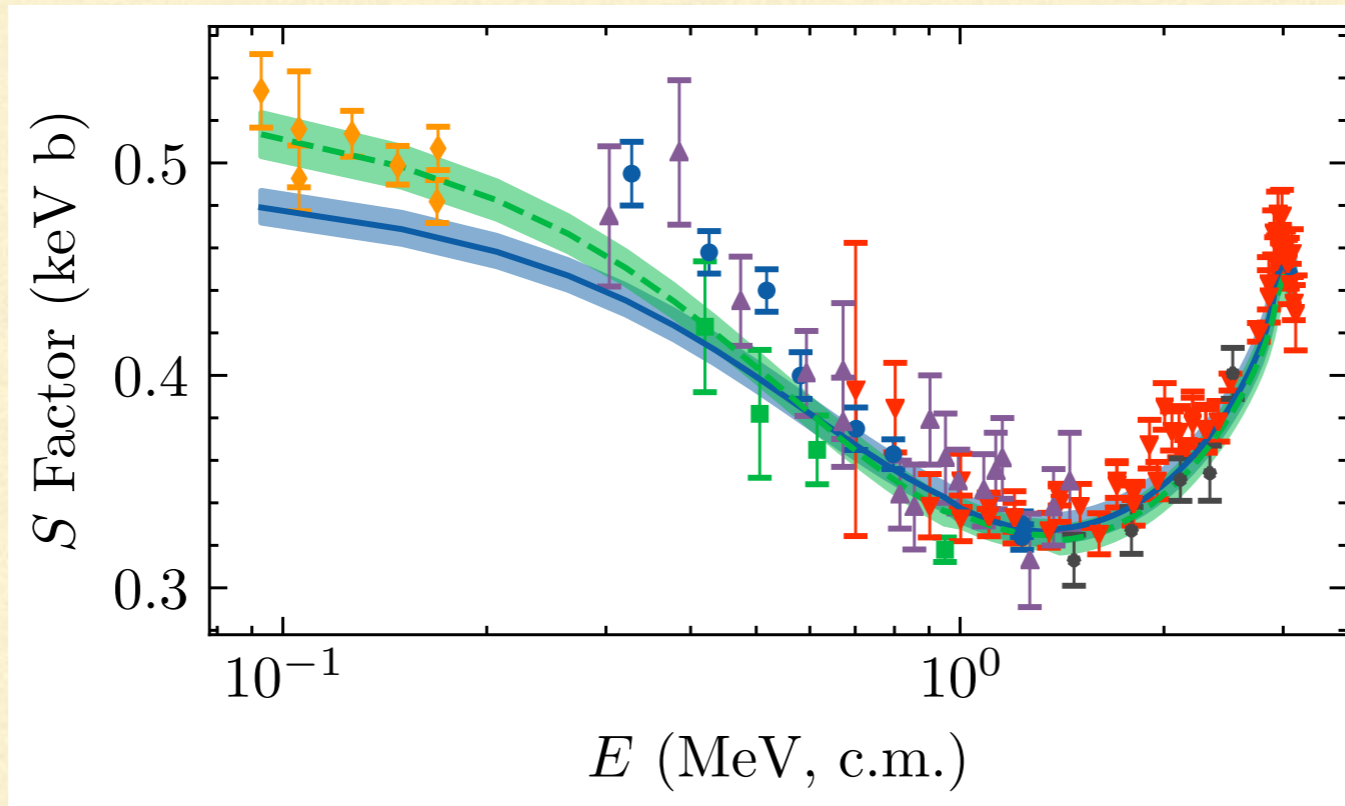


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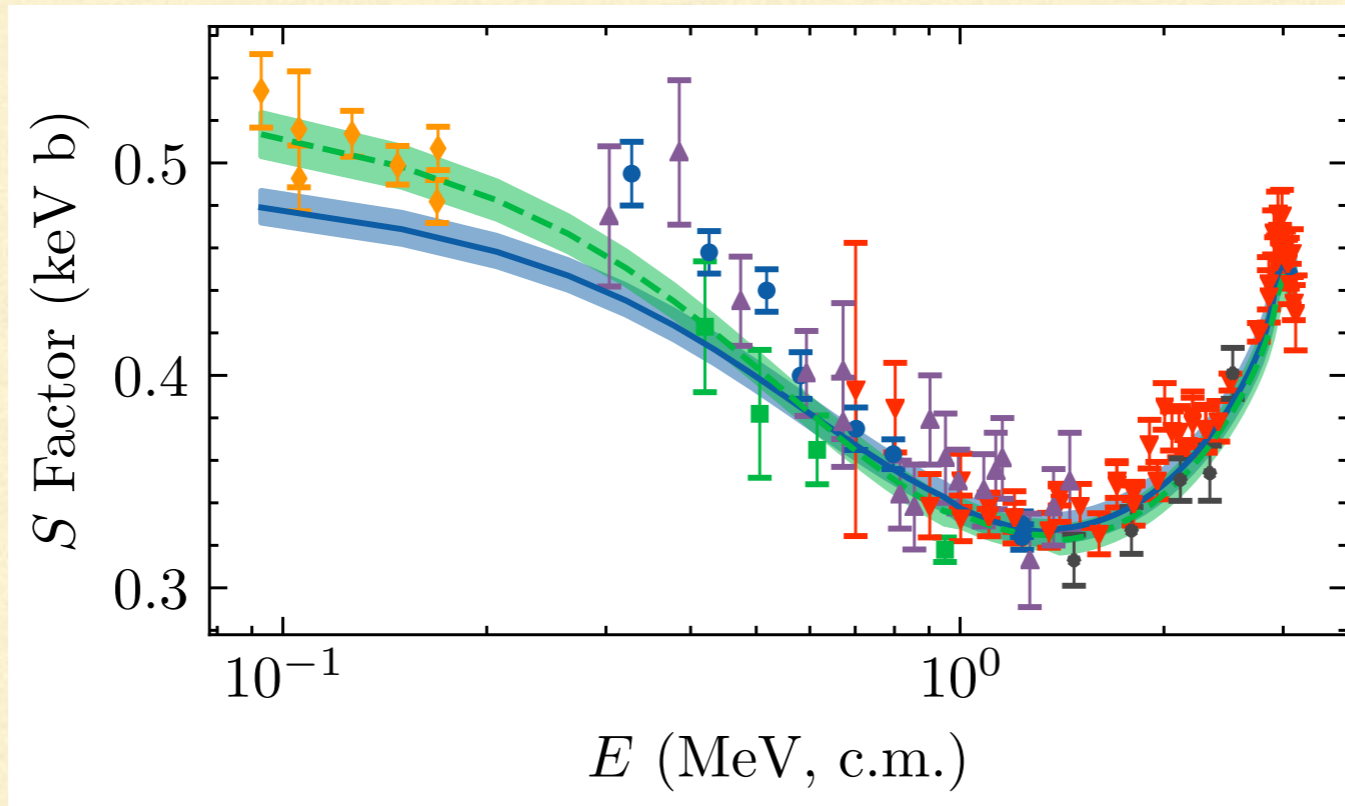
Speaking of which: SONIK data looks good



What about S-factor at solar energies?

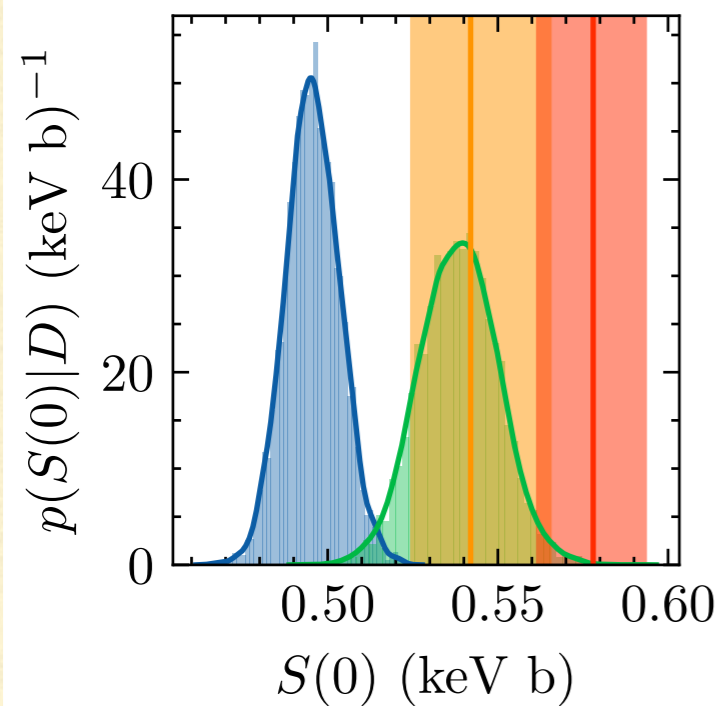
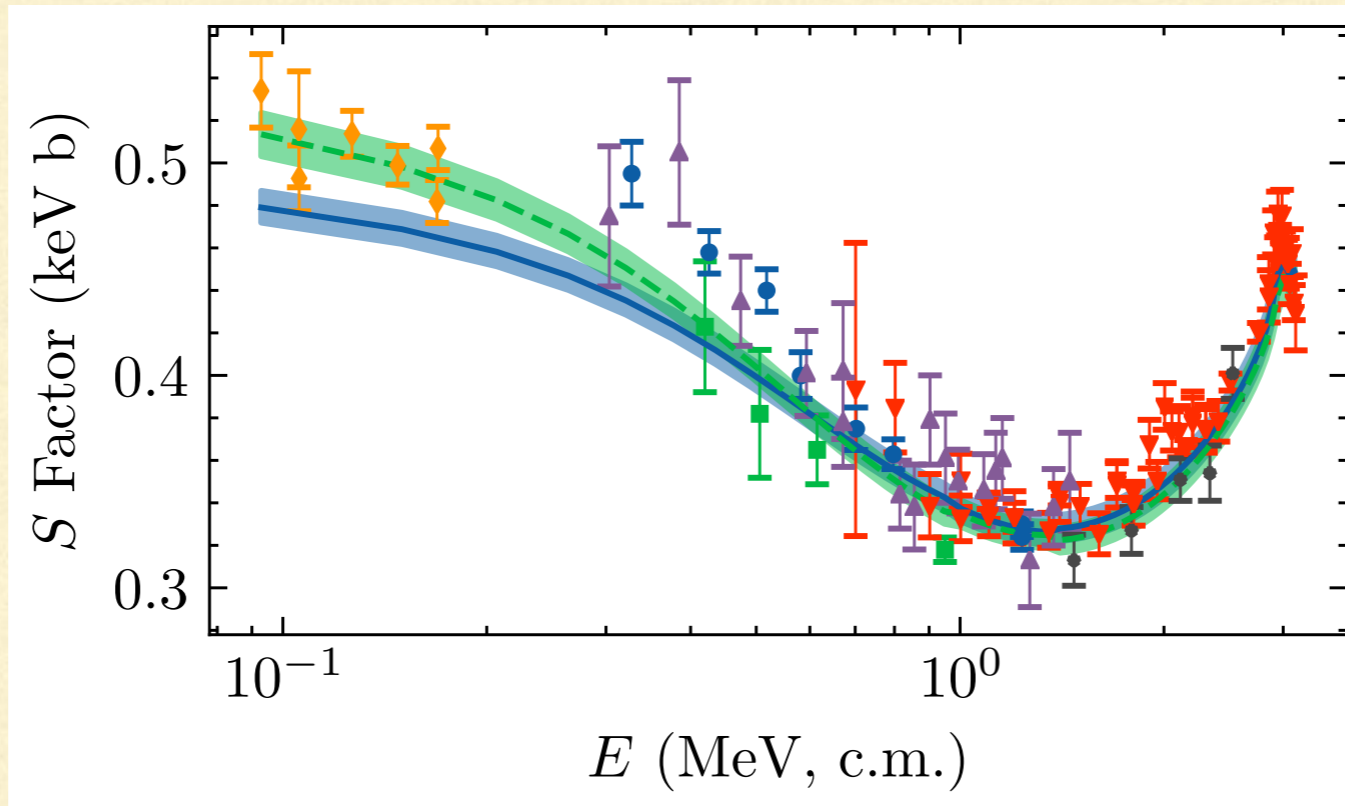


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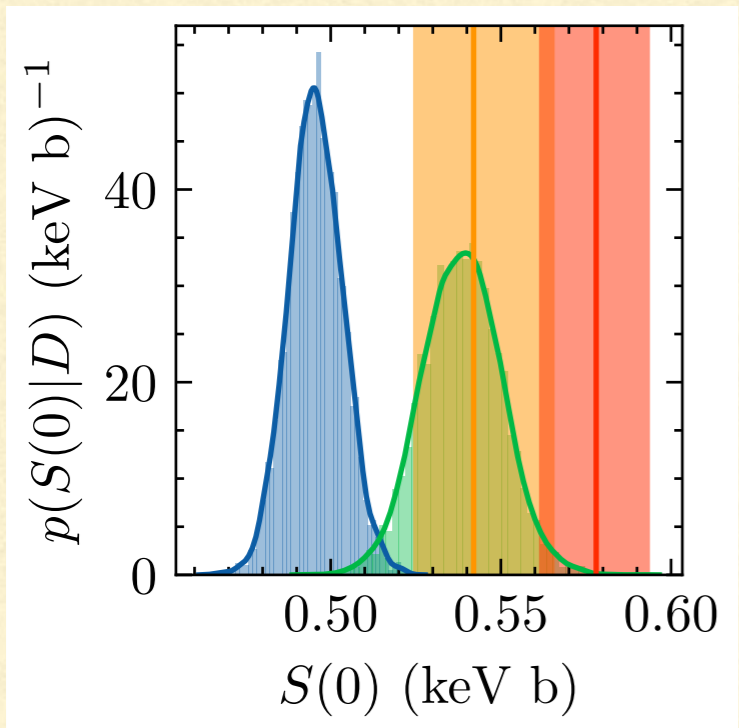
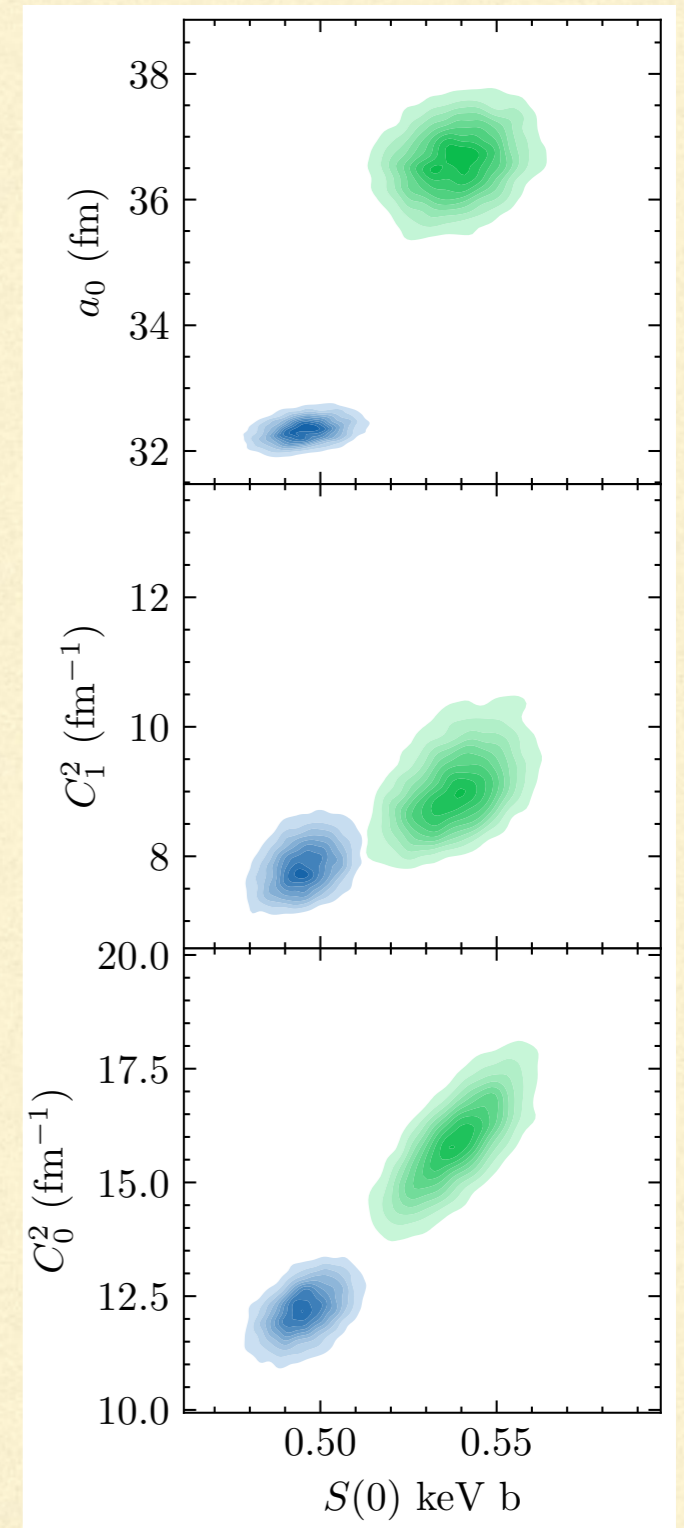
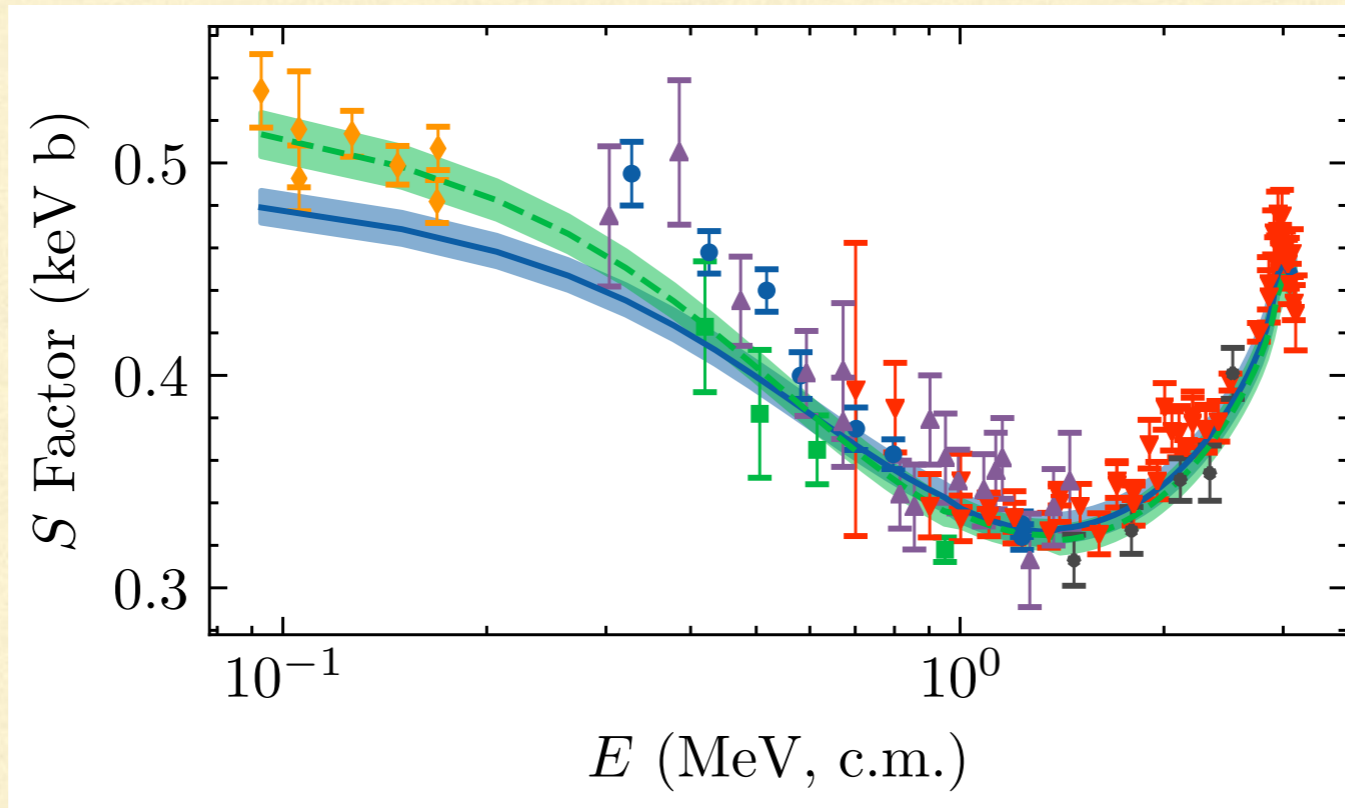
- Blue: CSB
- Green: CS
- Orange: de Boer et al.
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Bayesian analysis of the dt reaction

Odell, Brune, DP, Phys. Rev. C (2022) cf. de Souza et al., Phys. Rev. C (2019)

Bayesian analysis of the dt reaction

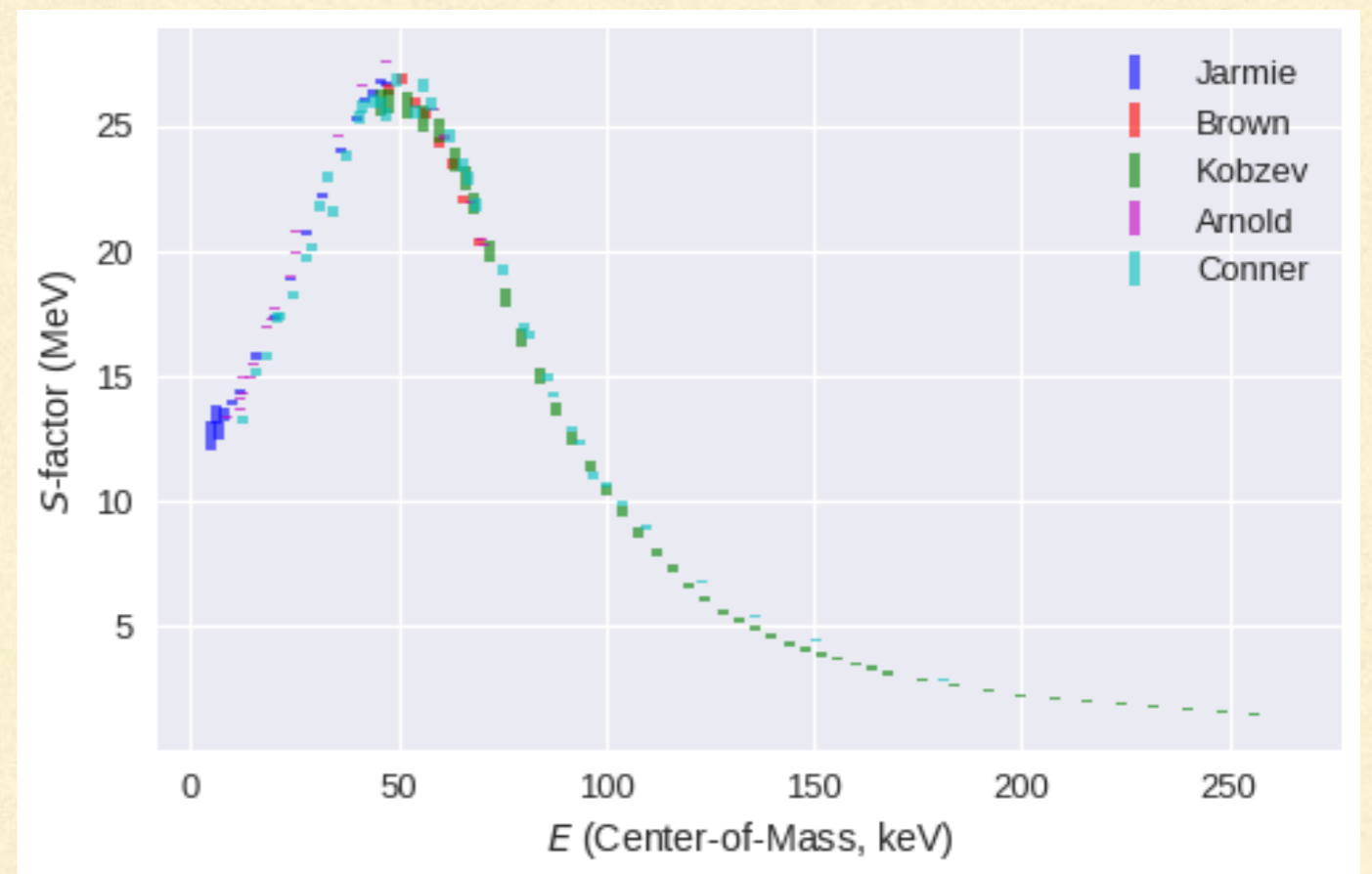
Odell, Brune, DP, Phys. Rev. C (2022) cf. de Souza et al., Phys. Rev. C (2019)

- ${}^2\text{H} + {}^3\text{H} \rightarrow {}^4\text{He} + \text{n}$ reaction important for applications; also part of BBN chain
 - Here we will use five data sets: Jarmie, Brown, Kobzev, Arnold, Conner
 - Precise data for $E_{\text{c.m.}}=5\text{-}260$ keV with stated normalization errors of 1.26-2.5%
-

Bayesian analysis of the dt reaction

Odell, Brune, DP, Phys. Rev. C (2022) cf. de Souza et al., Phys. Rev. C (2019)

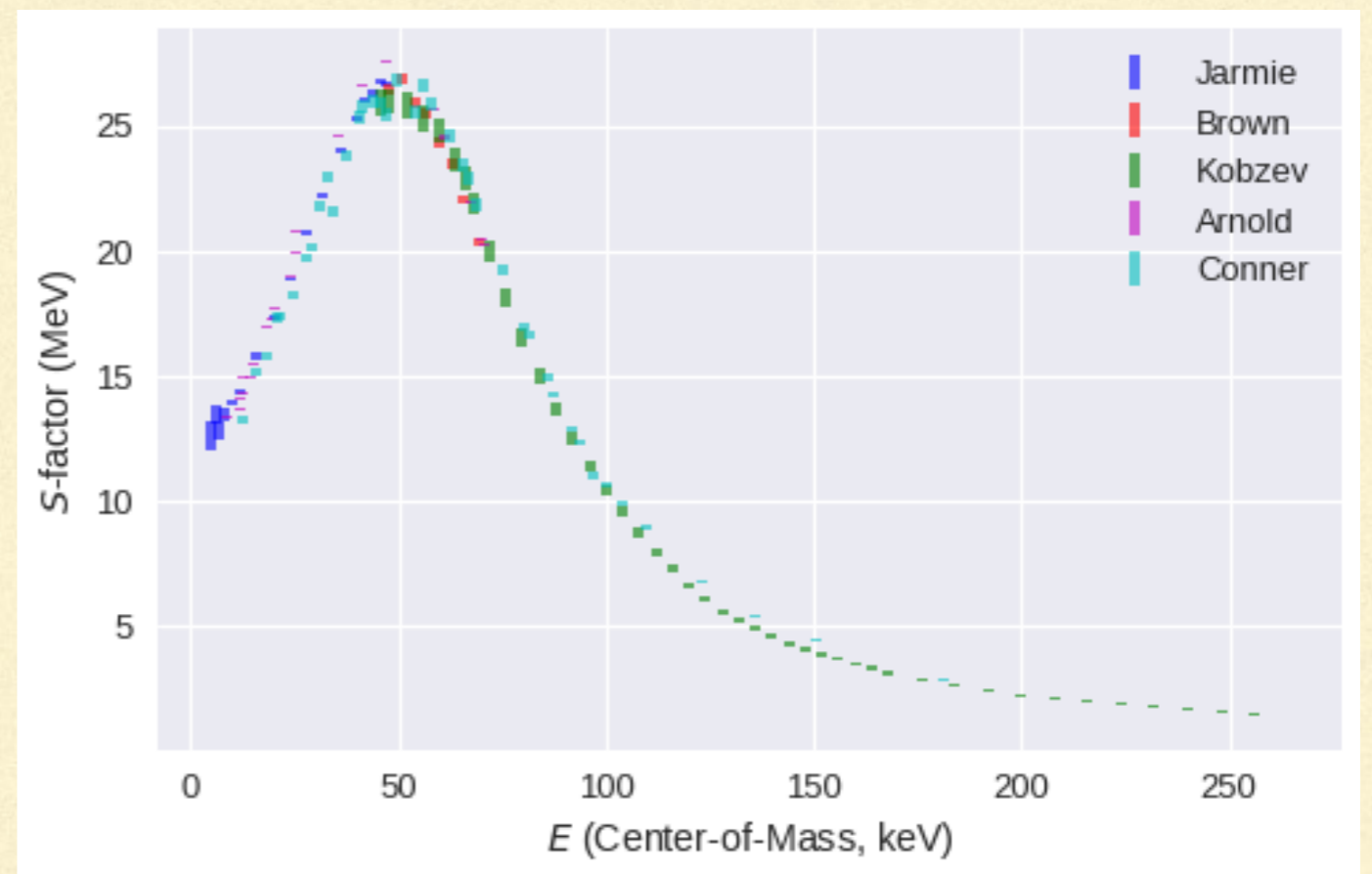
- ${}^2\text{H} + {}^3\text{H} \rightarrow {}^4\text{He} + \text{n}$ reaction important for applications; also part of BBN chain
- Here we will use five data sets: Jarmie, Brown, Kobzev, Arnold, Conner
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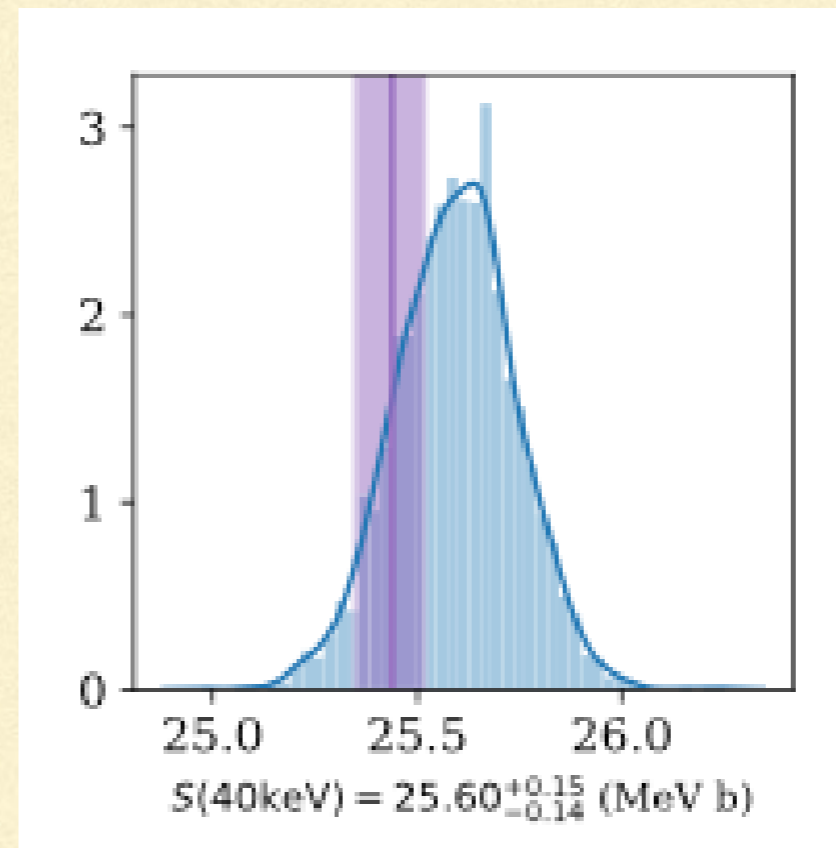
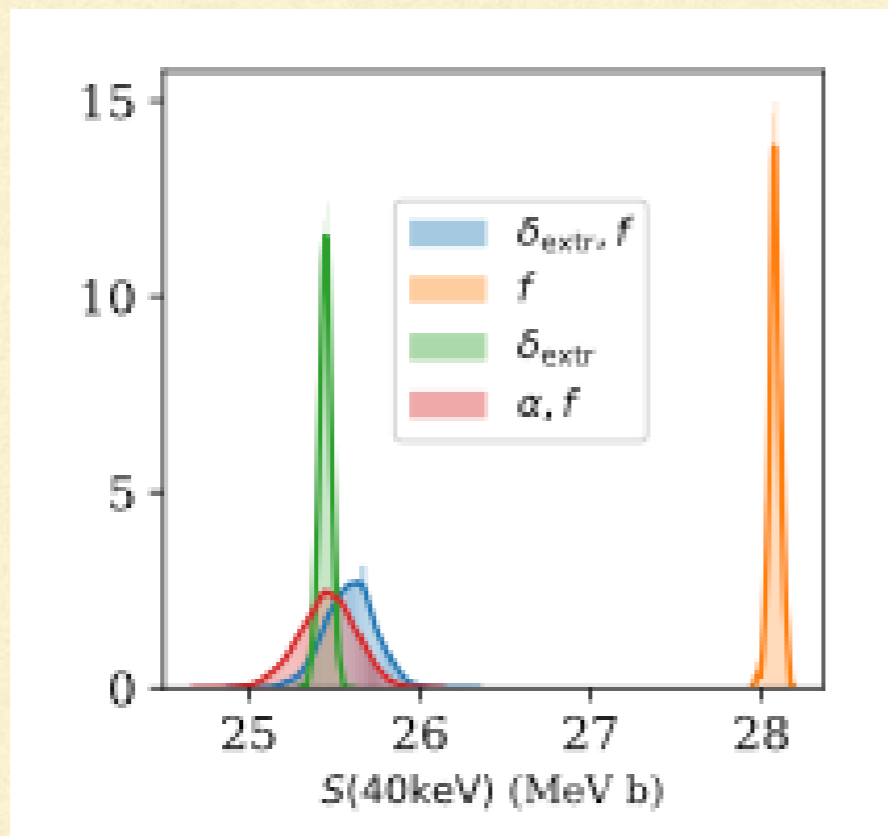
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- Precise data for $E_{\text{c.m.}}=5\text{-}260$ keV with stated normalization errors of 1.26-2.5%
- Goal: $S(40 \text{ keV})$ plus error bar



Improved statistical model

Odell, Brune, DP, Phys. Rev. C (2022) cf. de Souza et al., Phys. Rev. C (2019)

- A. Include “extrinsic errors”: additional point-to-point uncertainty, added in quadrature to nominal statistical error. Take $\sigma_{\text{extr},j}$ to be one number in MeV.b for each data set j and sample it to infer what it might be.
- B. Take $\sigma_{\text{point-to-point}} = \sqrt{\sigma_{i,j}^2 + \alpha_j^2 S_{i,j}^2}$ with α_j common to all points in data set j . (Relative extrinsic uncertainty rather than absolute extrinsic uncertainty.)

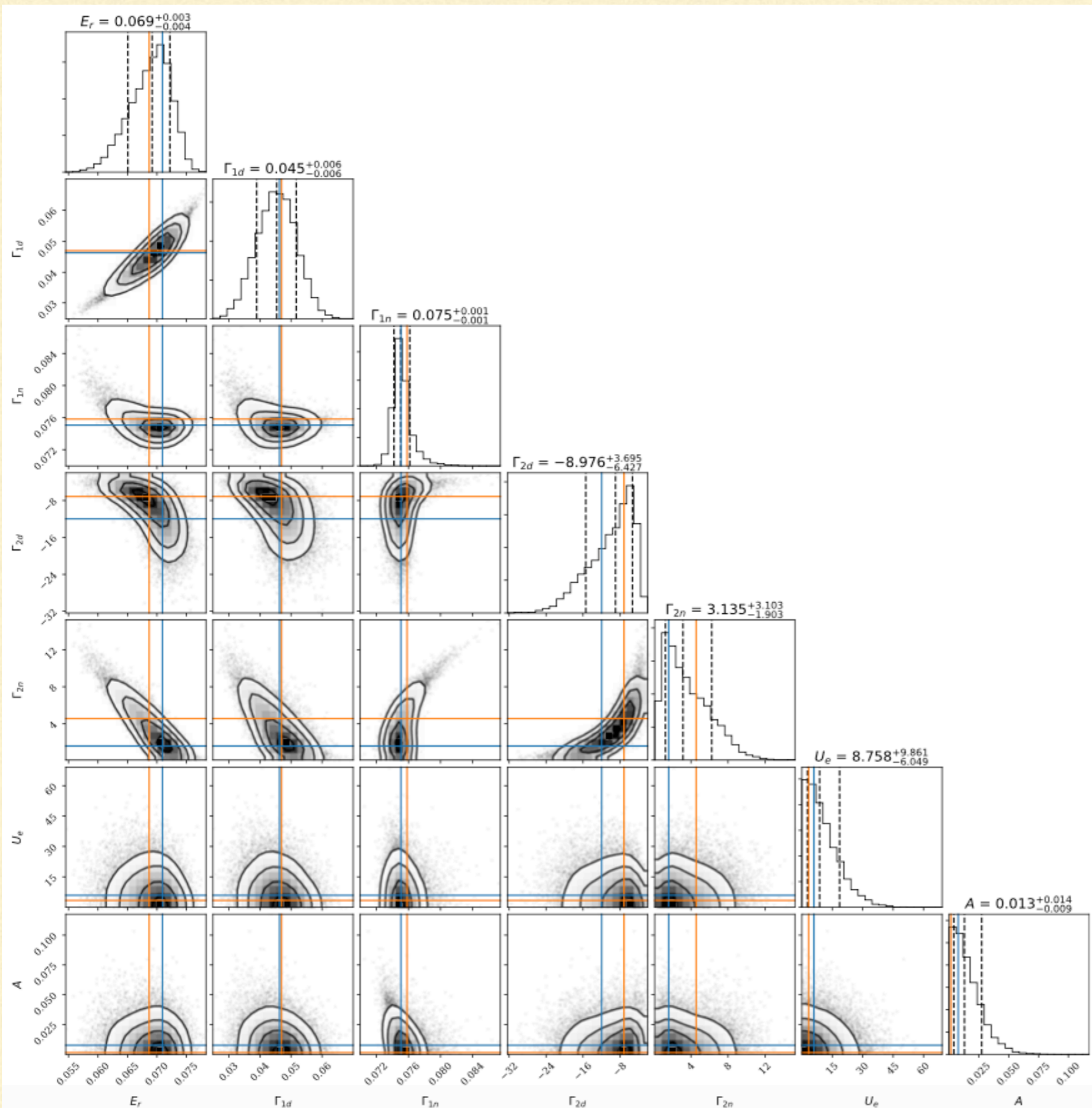


Purple:
de Souza et al.

Posteriors for our best model

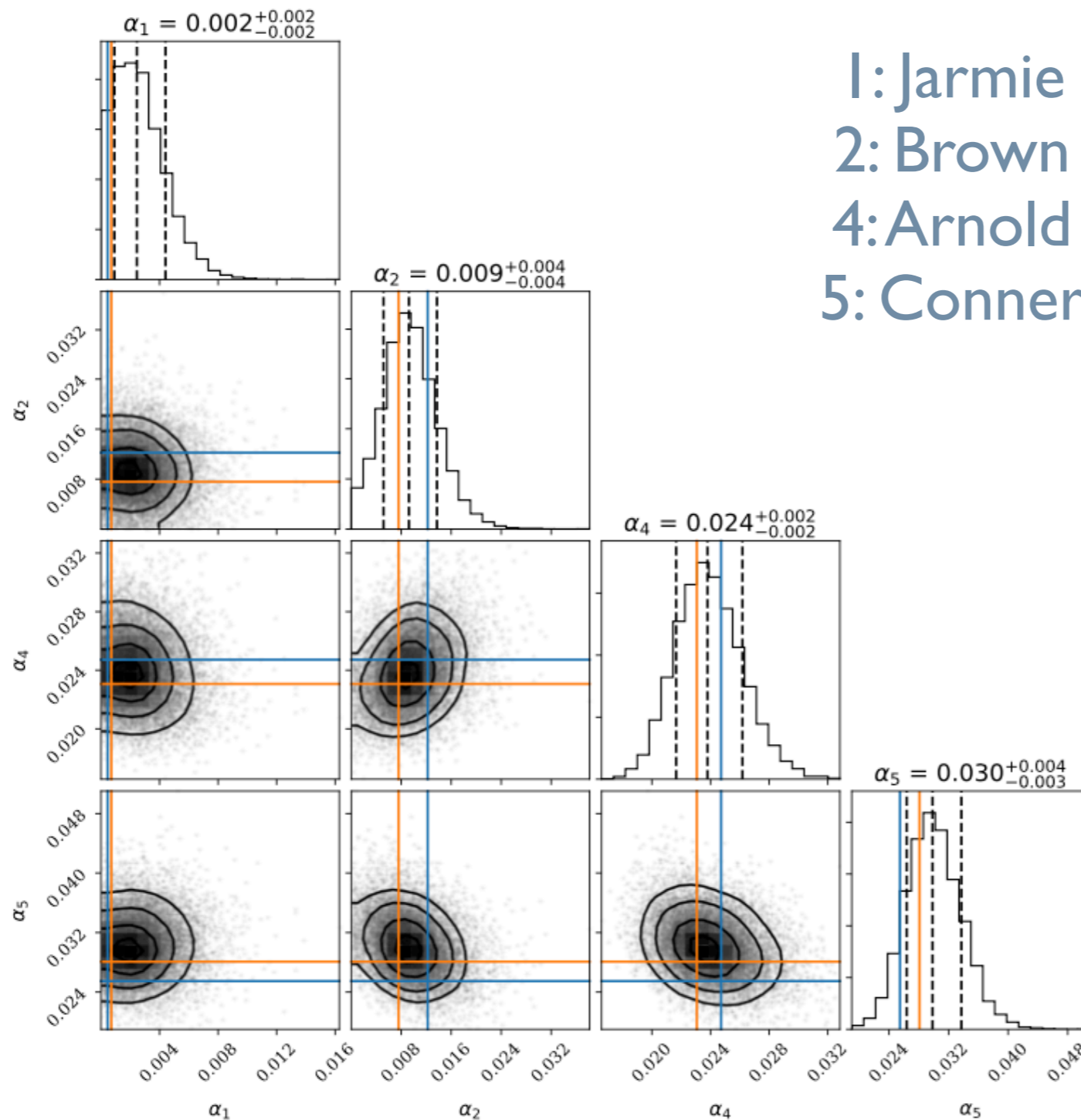
R-matrix parameters

- ${}^5\text{He}$ $3/2^+$ resonance energy
- ${}^5\text{He}$ $3/2^+$ resonance d width
- ${}^5\text{He}$ $3/2^+$ resonance n width
- ${}^5\text{He}$ $3/2^+$ background resonance parameters
- Screening potential
- ${}^5\text{He}$ $1/2^+$ background resonance parameter combination



Posteriors for our best model

1: Jarmie
2: Brown
4: Arnold
5: Conner

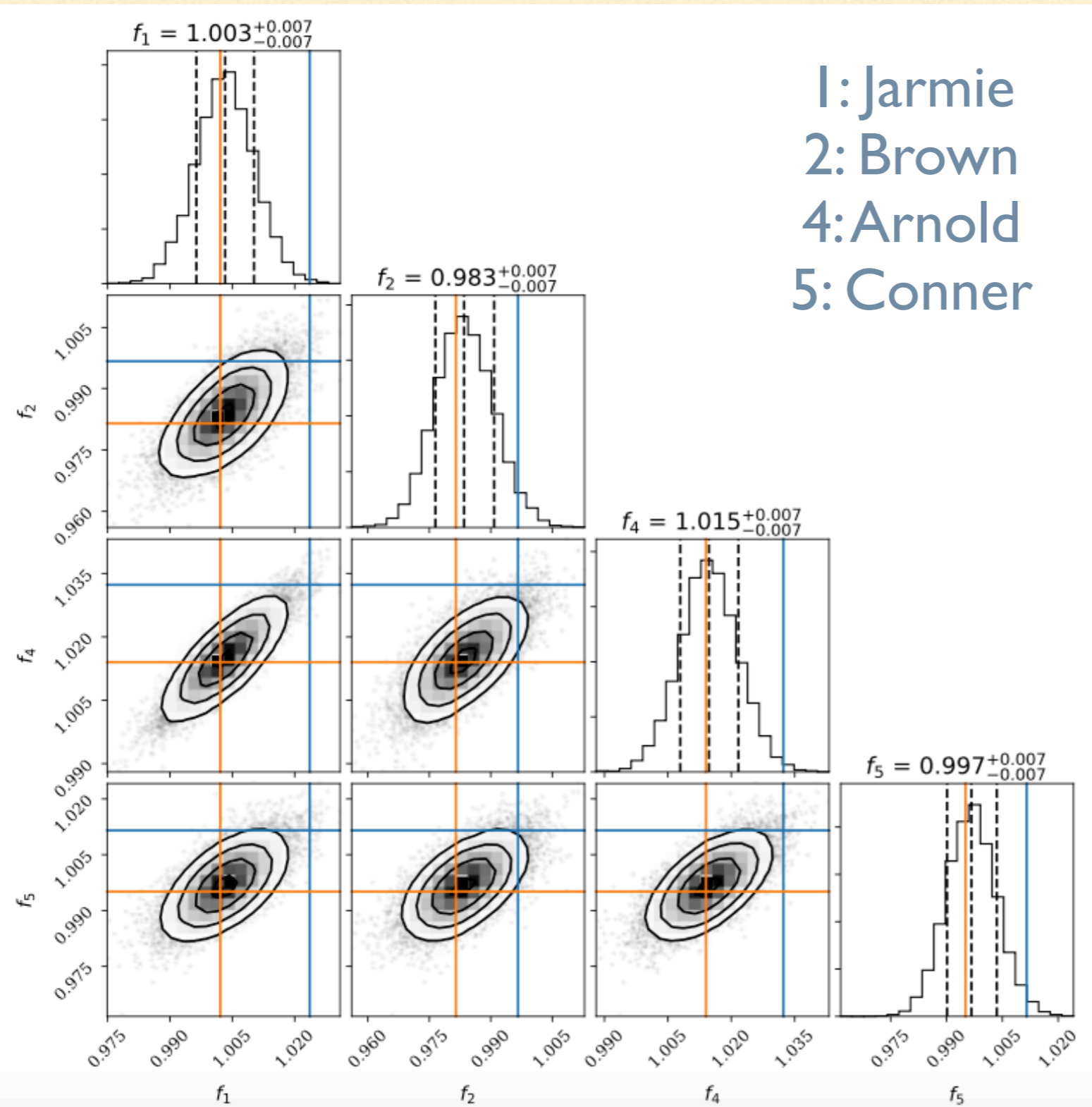


R-matrix parameters

- ^5He 3/2+ resonance energy
- ^5He 3/2+ resonance d width
- ^5He 3/2+ resonance n width
- ^5He 3/2+ background resonance parameters
- Screening potential
- ^5He 1/2+ background resonance parameter combination

Relative extrinsic uncertainty and normalization for each data set

Posteriors for our best model



1: Jarmie
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R-matrix parameters

- ^5He $3/2^+$ resonance energy
- ^5He $3/2^+$ resonance d width
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Relative extrinsic uncertainty and normalization for each data set

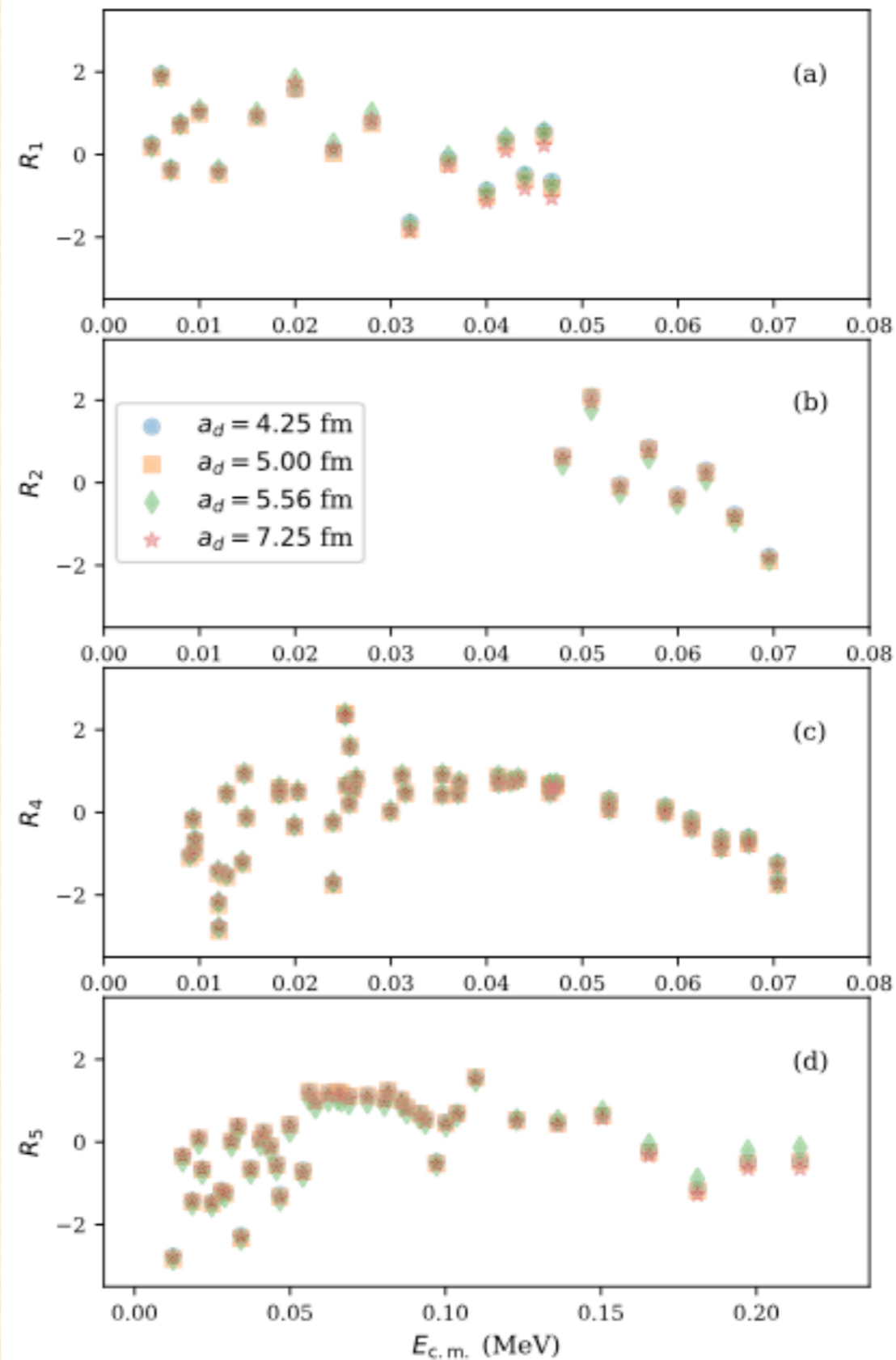
Residuals

Jarmie

Brown

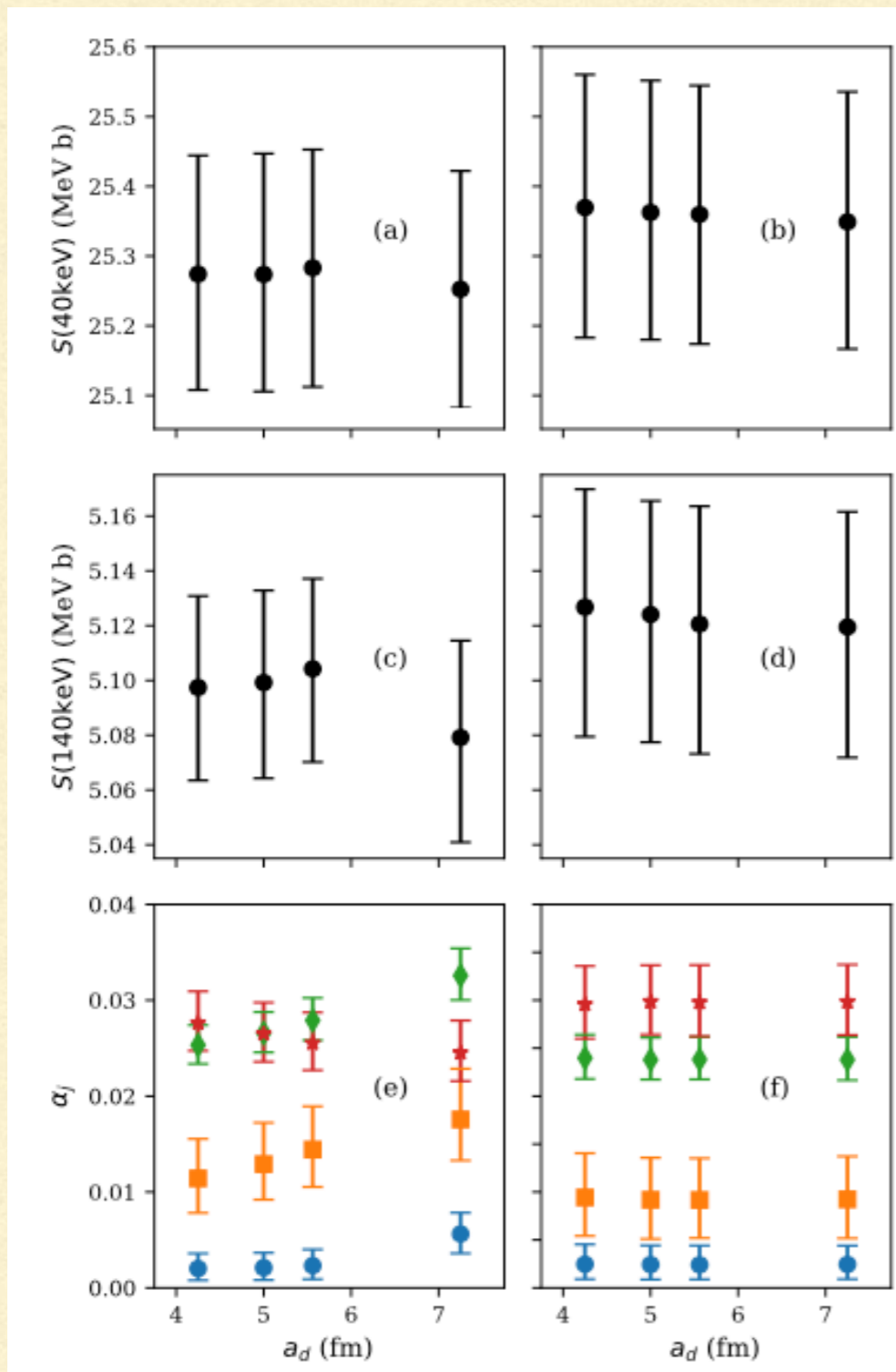
Arnold

Conner



cf. Kobzev
residuals

Result, comparisons, implications



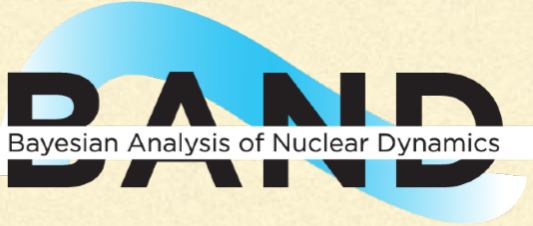


- Best R-matrix model also includes $1/2+$ background level, and $3/2+$ background level
- Physical results are then independent of channel radius
- Extrinsic errors needed for consistency, experimental norms. then more uncertain
- $S(40\text{ keV})=25.36 \pm 0.19\text{ MeV} \cdot b$
- Consistent with de Souza et al., but with error bar that is roughly 2 x bigger
- cf. Bosch & Hale (1993):
 $S(40\text{ keV})=25.87 \pm 0.49\text{ MeV} \cdot b$

Further applications of BRICK

- $^{19}\text{F}(p, \gamma)^{20}\text{Ne}$ – Zhang et al. (incl. deBoer, Odell) Nature **610**, 656-660 (2022)
 - Low-energy resonance opens up possibility of “warm” CNO breakout
 - $^{10}\text{B}(p, \alpha)^7\text{Be}$ – Van de Kolk et al. (incl. deBoer, Odell) PRC **105**, 055802 (2022)
 - possible temperature probe for $^{11}\text{B}(p, 2\alpha)^4\text{He}$ – aneutronic plasma fusion source
 - $^{23}\text{Na}(p, \gamma)^{24}\text{Mg}$ – Boeltzig et al. (incl. deBoer, Odell) PRC **106**, 045801 (2022)
 - breakout reaction linking NeNa and MgAl cycles
 - $^{13}\text{C}(\alpha, n_1)^{16}\text{O}$ – deBoer et al. (incl. deBoer, Odell) PRC **106**, 055808 (2022)
 - partial cross section measurement, improves BG modeling
-

Summary

- Parametric uncertainties in R-matrix analyses can be quantified by MCMC sampling of the Bayesian posterior and evaluating derived quantities
 -  <https://github.com/odell/brick>
 - 
 - 
 - Multiple examples of successful application to different reactions
 - Enables more sophisticated modeling of experimental imperfections
 - Knowledge of full posterior provides access to parameter correlations, allows diagnosis of which parameters are not needed, shows where there is multi-modality, non-Gaussianity, and more
 - Error propagation to derived quantities is straightforward with samples in hand
 - Model checking (residuals, coverage, etc.) needs to be done at end
 - Model uncertainties of R-matrix analysis? Comparison to EFT, ab initio, etc.
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