



DETAILED CHARGE EXCHANGE NEUTRAL DISTRIBUTION MODELLING FOR THE ITER MAIN WALL

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Rationale

- ERO2.0 requires detailed distribution of neutral particles impinging specific „diagnostic“ surfaces in ITER, i.e. mirrors
(c.f. THU talk Ch. Baumann, W-erosion in EU-DEMO)
- EIRENE post-processing of S. Lisgo's ITER „repository“ of SOLPS4.3 simulations with DivIMP extension of grids up to FW
- Set of simulations based on A. Khan NME 2019 WalldYN paper
- Some modifications in EIRENE to also have both energy and angular distributions
- Python tools for plotting & (statistical) data analysis after re-processing with EIRENE

Various extrapolation models for the ITER far-SOL

- d** L-mode pedestal, low density far-SOL ($v_{\perp} = 35$ m/s), low temperature far-SOL ($T_e = 10$ eV, $T_i = 20$ eV)
- o** H-mode pedestal, high density far-SOL ($v_{\perp} = 100$ m/s), high temperature far-SOL ($T_e = 20$ eV, $T_i = 40$ eV)
- g** H-mode, $v_{\perp} = 35$ m/s, $T_e = 20$, $T_i = 40$ eV
- m** H-mode, $v_{\perp} = 100$ m/s, $T_e = 10$, $T_i = 20$ eV

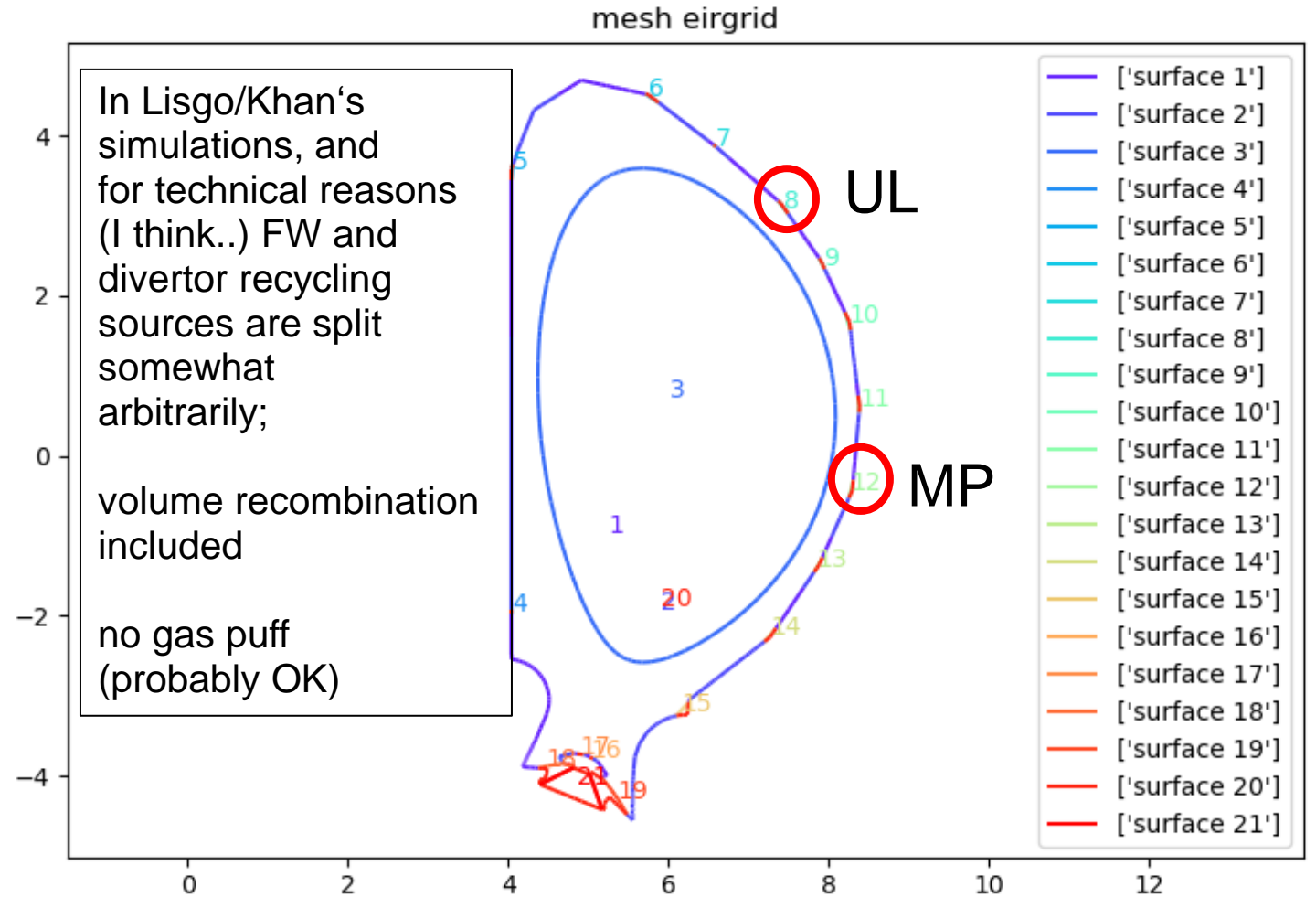
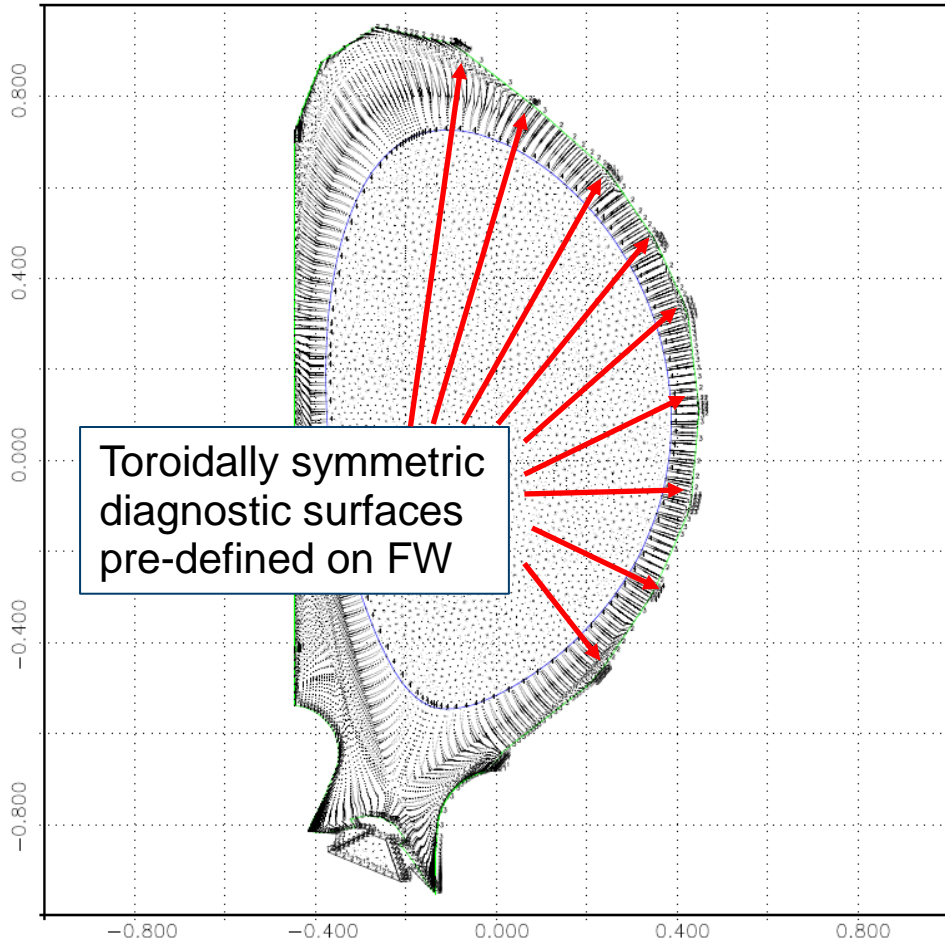
- 00** no near-SOL flow
- 01** near-SOL flow 0.5M

Name	Main SOL M_{\parallel}	Far-SOL v_{\perp} $m s^{-1}$	Far-SOL T_e eV	Far-SOL T_i eV
—	—	—	—	—
00d	0	30	10	20
00g	0	30	20	40
00k	0	65	20	40
00m	0	100	10	20
01d	0.5	30	10	20
01g	0.5	30	20	40
01k	0.5	65	20	40
01m	0.5	100	10	20

Concentrate on case series
2258_00*

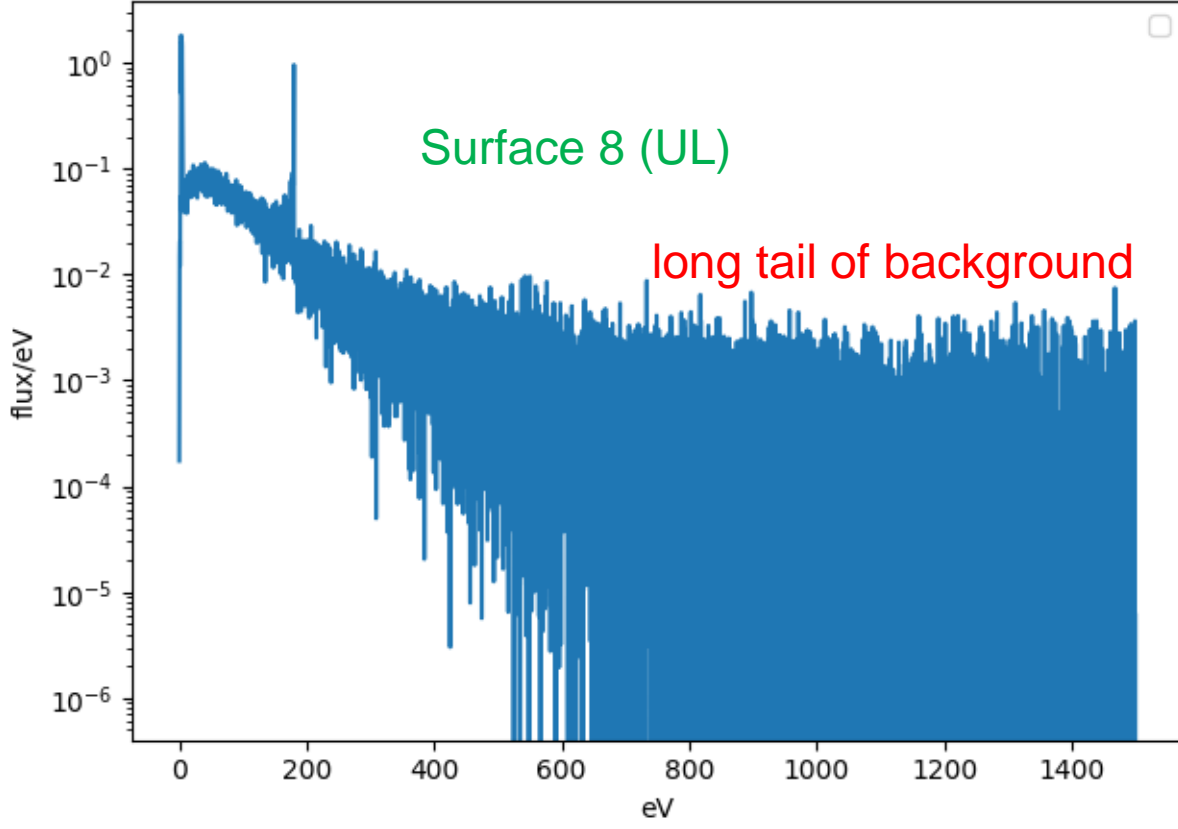
→ Semi-detached, high-power H-mode case

Collector surfaces (EIRENE tallies) for extended 2258 SOLPS4.3 geometry

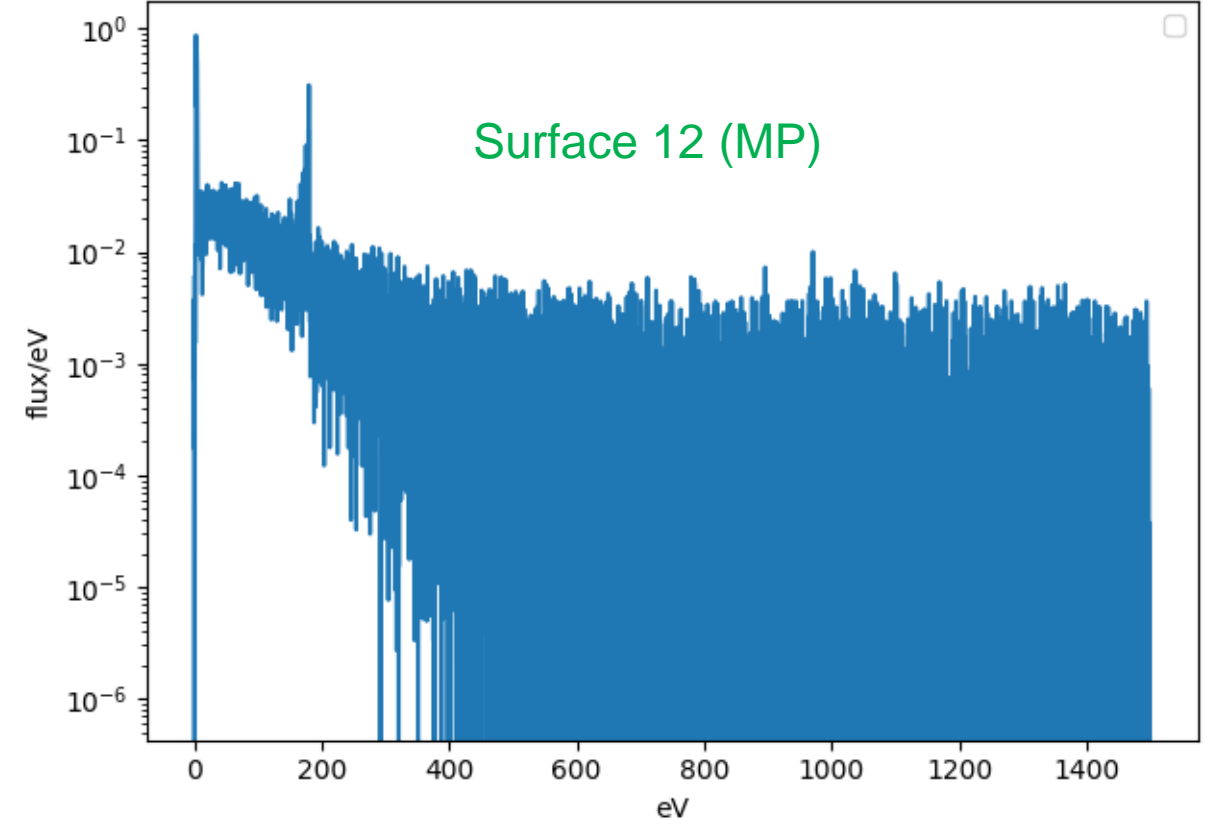


2258_00g D⁰ energy spectra

data/2258-00g_1e4/spec_e_8.dat



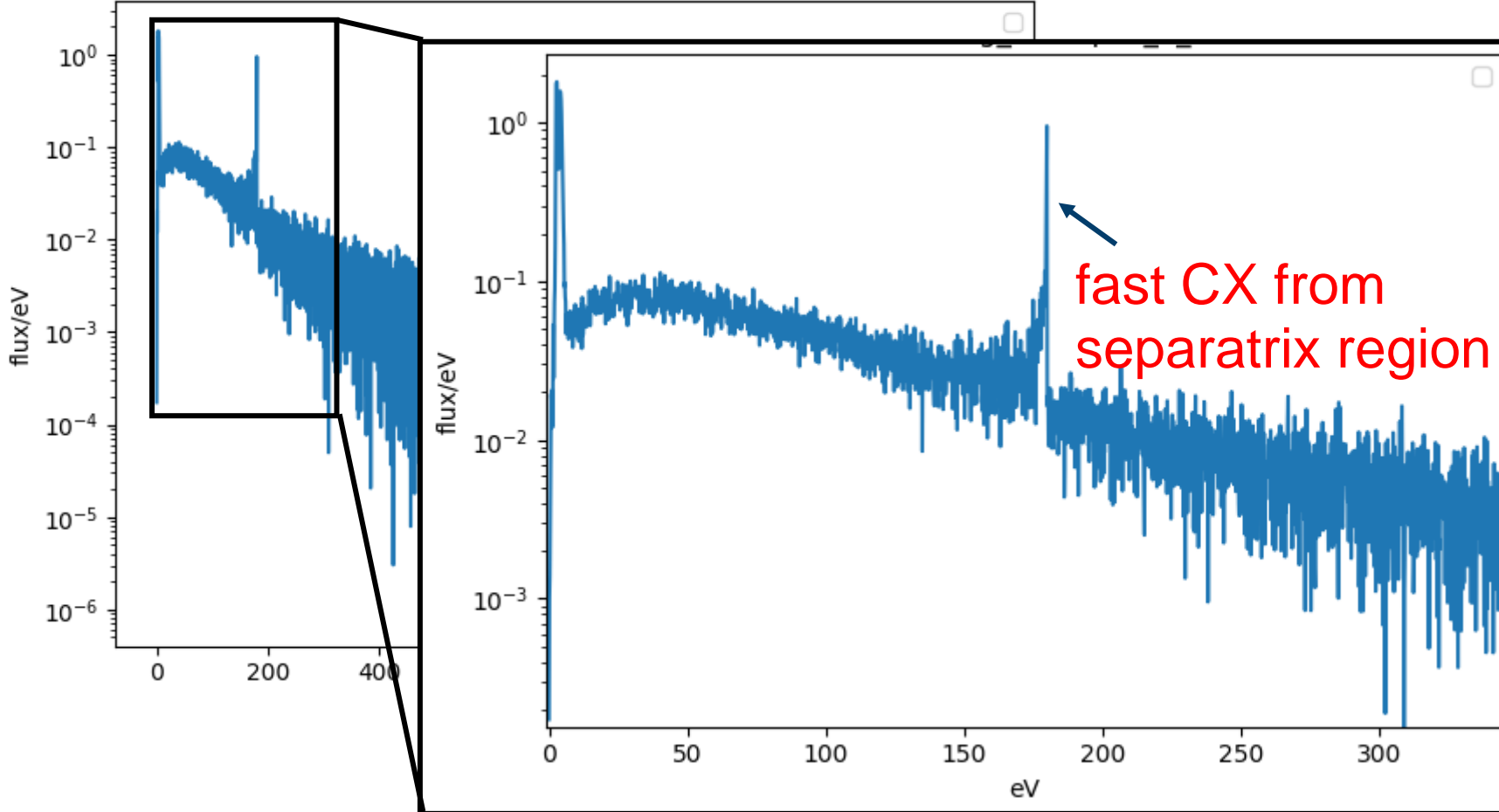
data/2258-00g_1e4/spec_e_12.dat



normalised „EIRENE-units“ Amp/eV

2258_00g D⁰ energy spectra

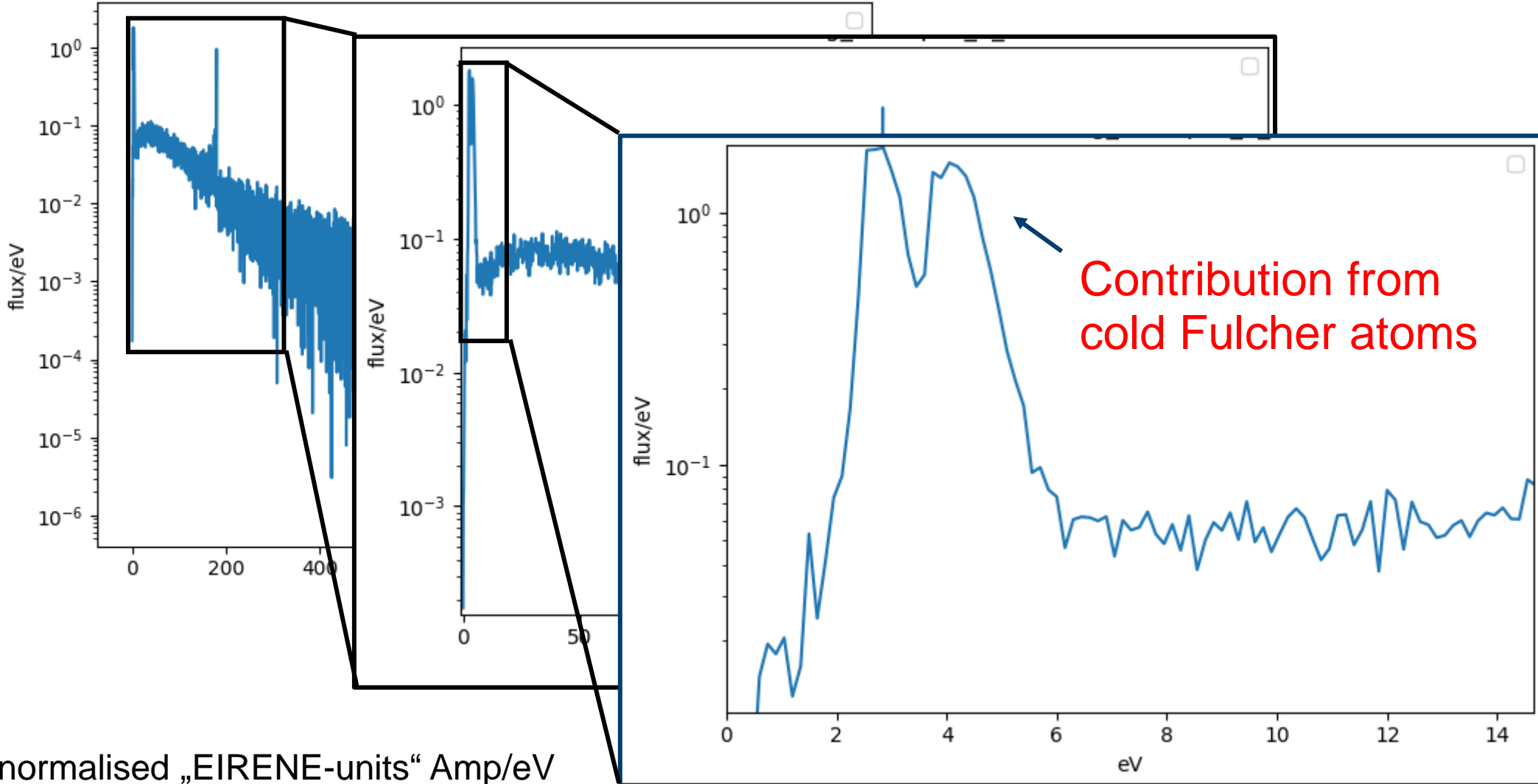
data/2258-00g_1e4/spec_e_8.dat



normalised „EIRENE-units“ Amp/eV

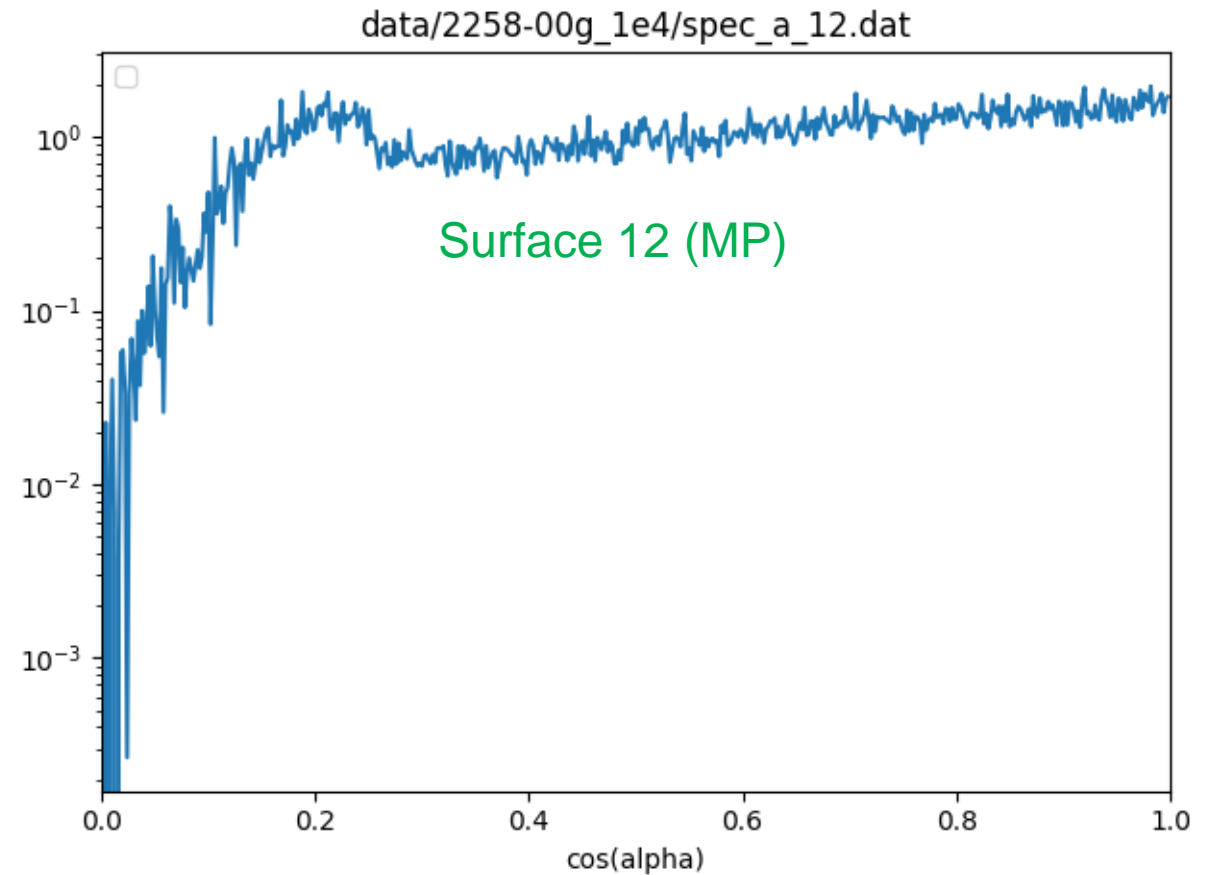
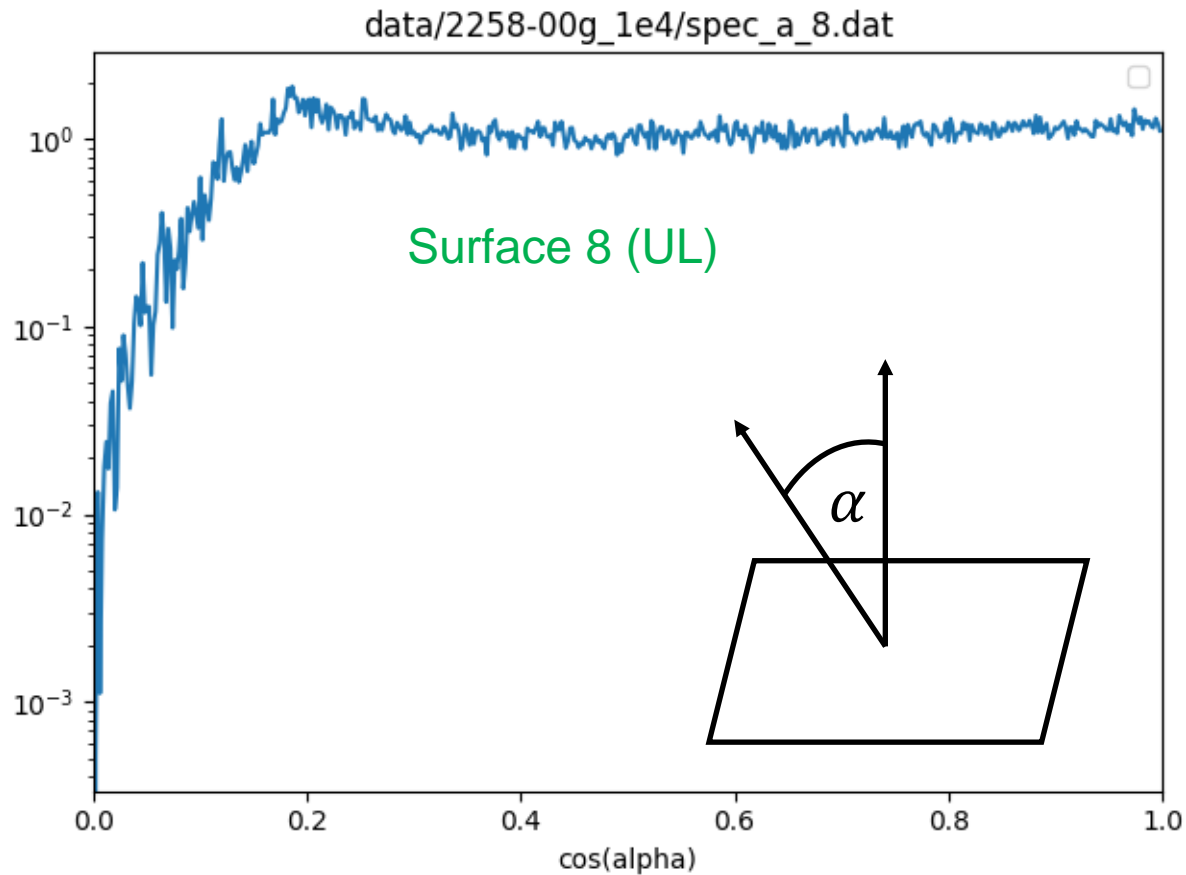
2258_00g D⁰ energy spectra

data/2258-00g_1e4/spec_e_8.dat



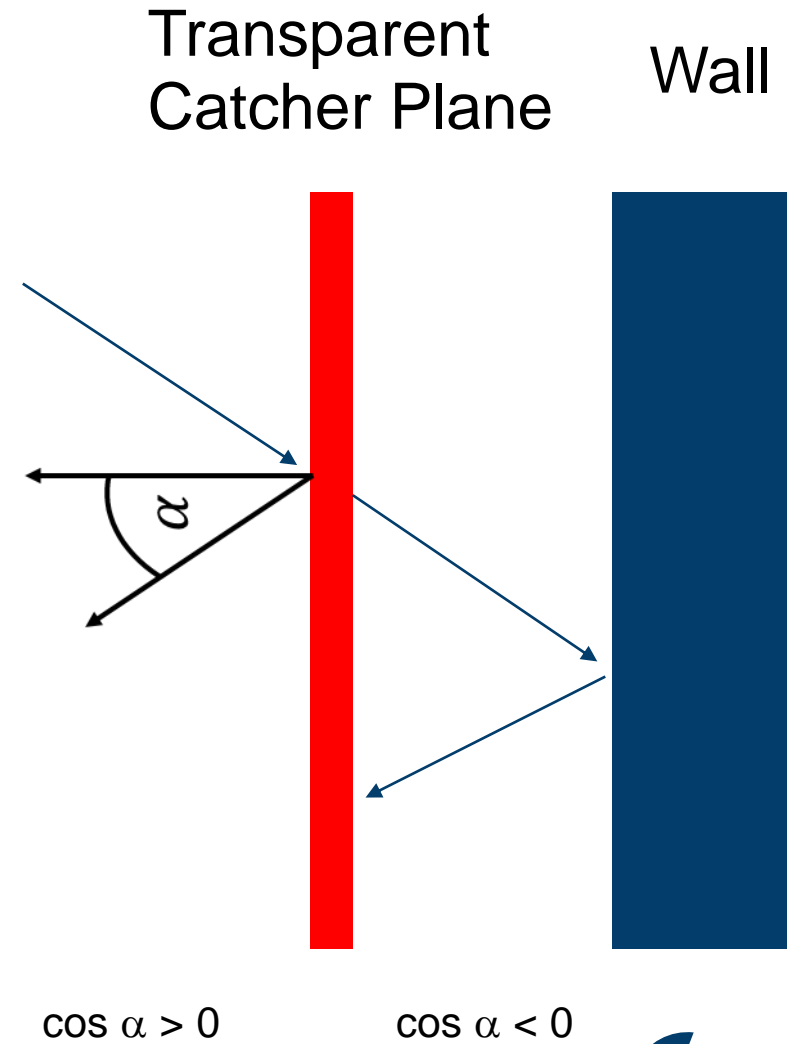
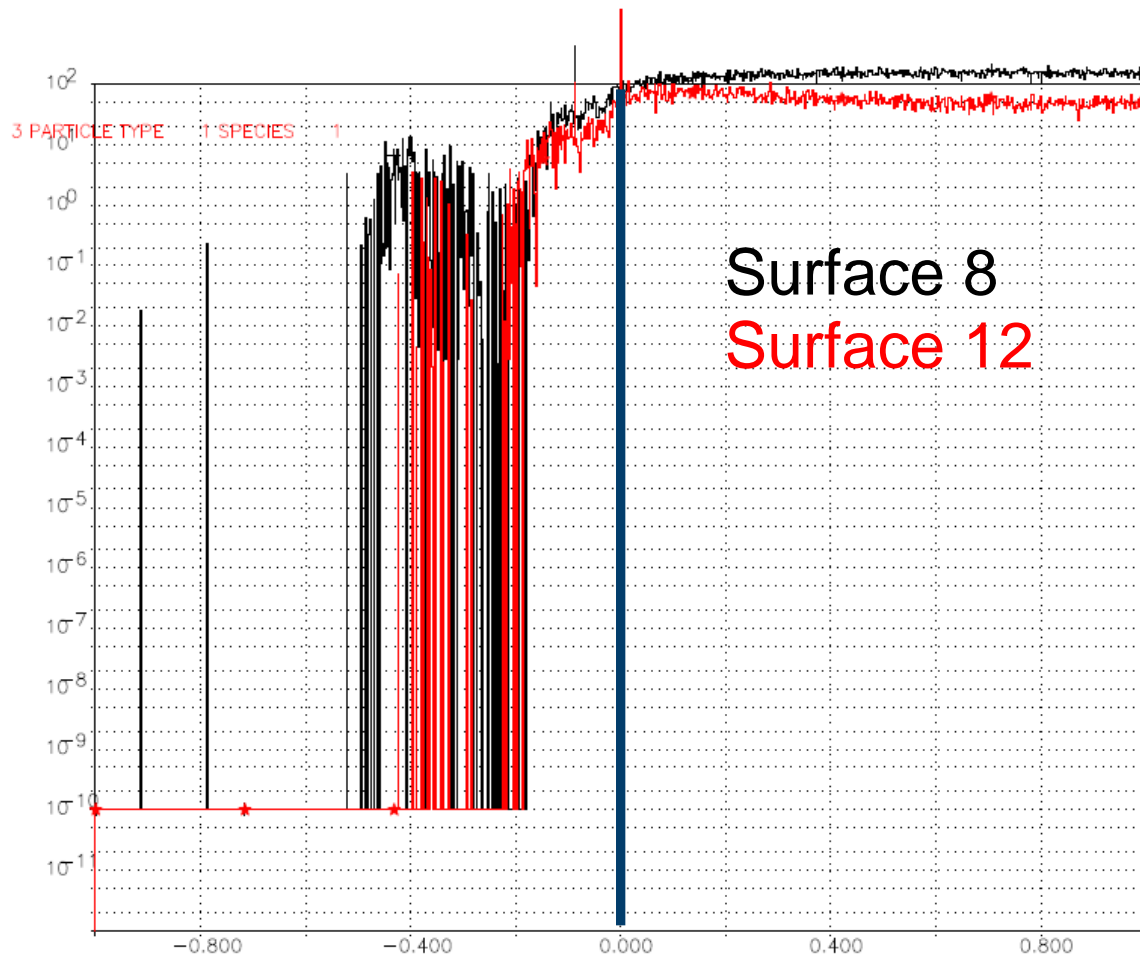
normalised „EIRENE-units“ Amp/eV

2258_00g D⁰ angular distribution (polar angle only, 2D-sym)

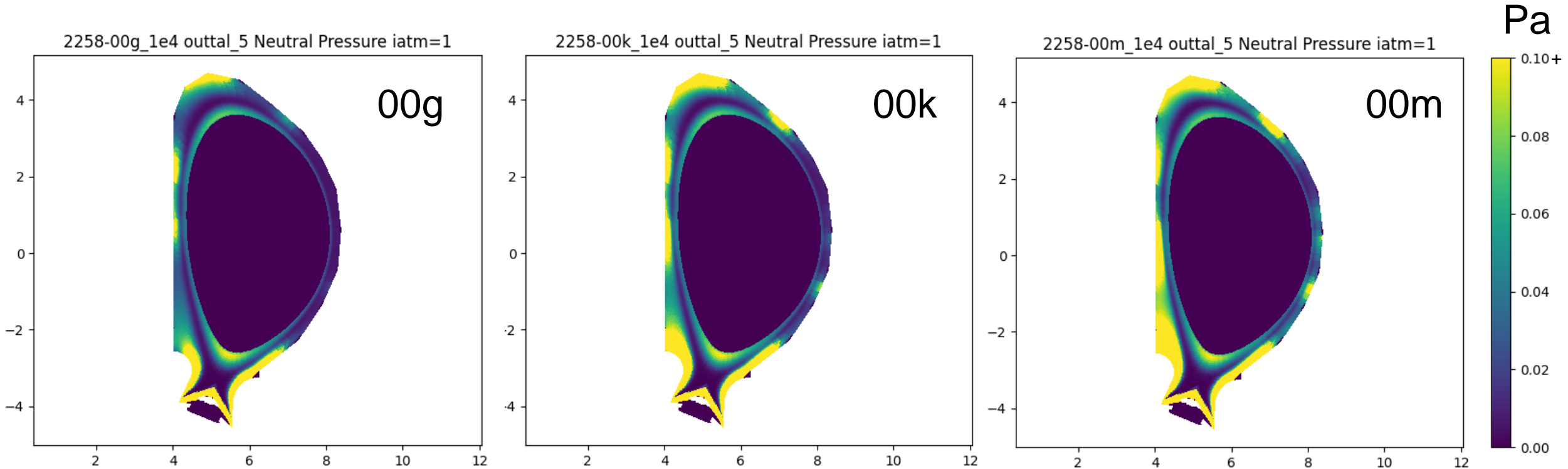


$f(\cos \alpha)$ normalised to 1

Expert's advice: don't tally wrong..

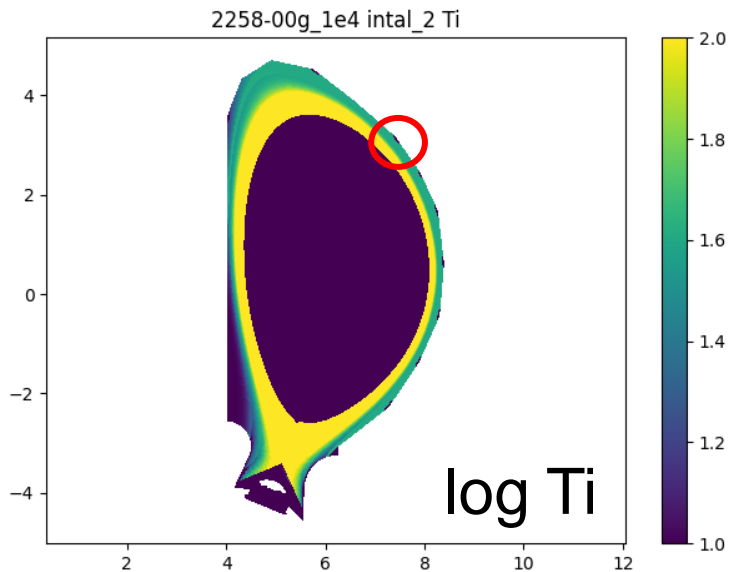
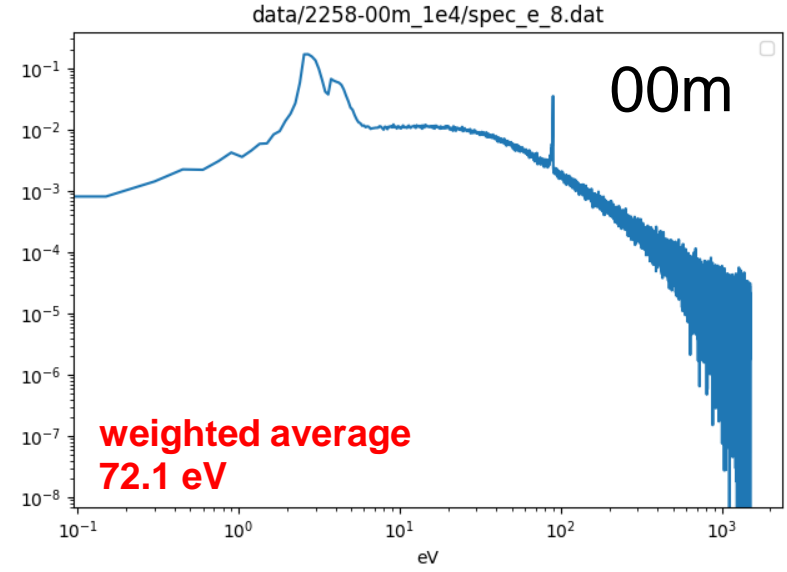
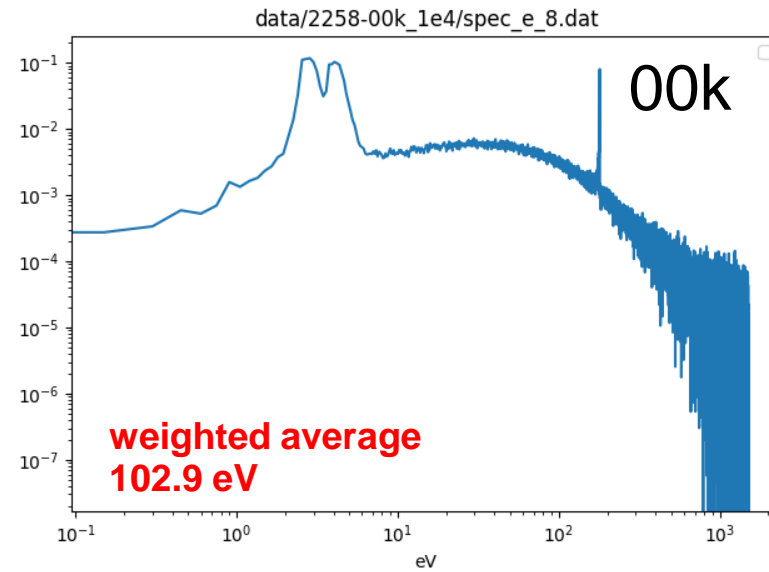
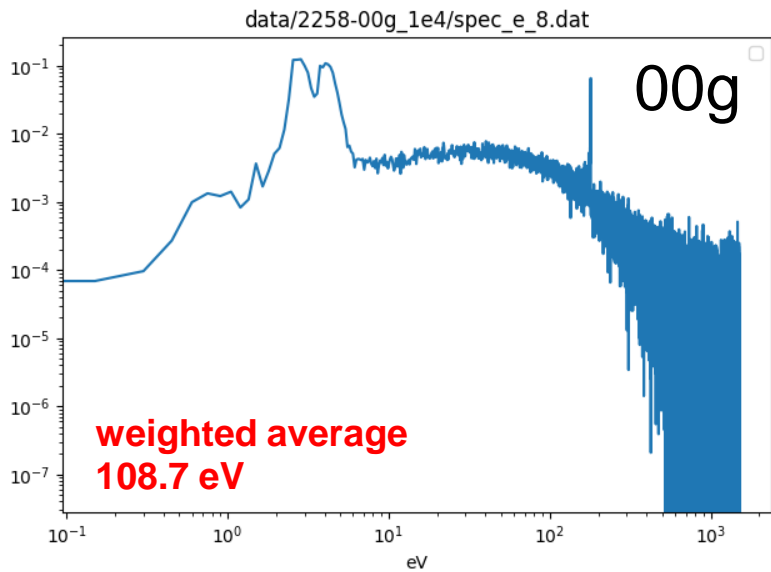


Comparison of far-SOL assumptions: D⁰ neutral pressure



Name	Main SOL M_{\parallel}	Far-SOL v_{\perp} $m s^{-1}$	Far-SOL T_e eV	Far-Sol T_i eV
—	—			
00g	0	30	20	40
00k	0	65	20	40
00m	0	100	10	20

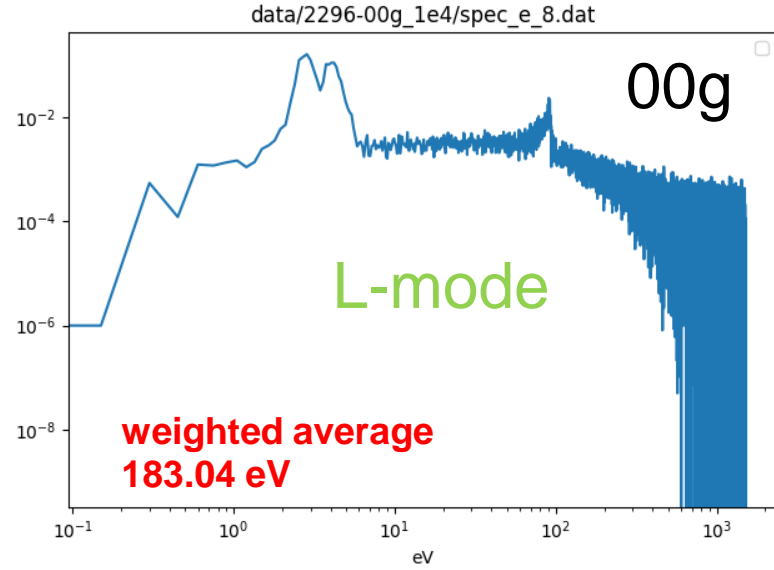
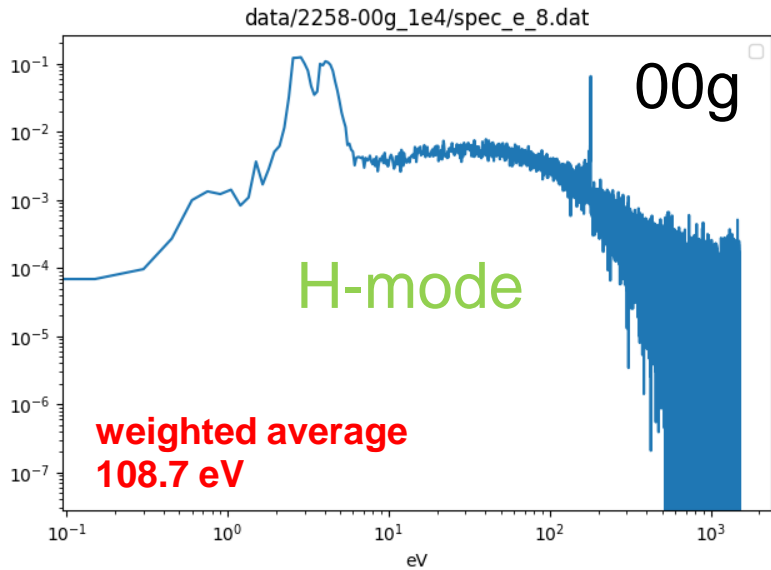
Comparison far-SOL assumptions: D⁰ spectra log-log (UL)



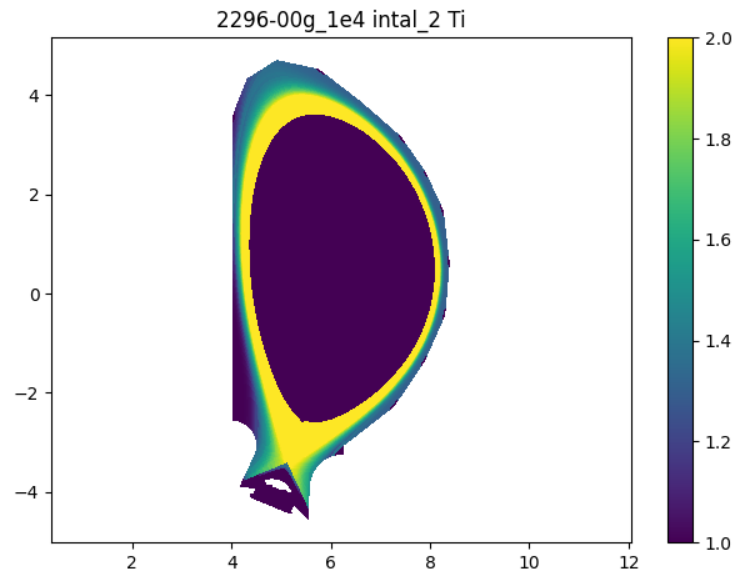
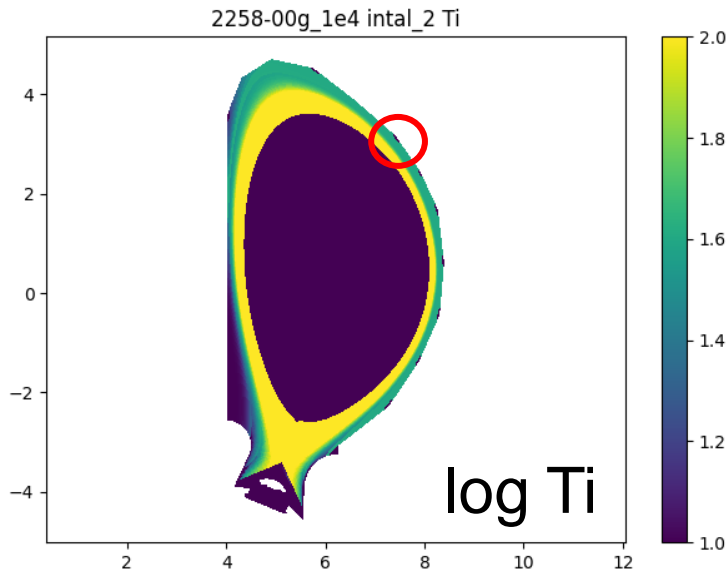
Name	Main SOL $M_{ }$	Far-SOL v_{\perp} $m s^{-1}$	Far-SOL T_e eV	Far-SOL T_i eV
—	—	—	—	—
00g	0	30	20	40
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00m	0	100	10	20

- The “spx region peak” (CX) shifts and weighted average $\langle E \rangle$ depend on far-SOL Ti-conditions

Comparison H-mode vs L-mode (00g)



- In L-mode the strong peak is less prominent
- However the average energy $\langle E \rangle$ is significantly higher than in H-mode (tail contribution)



Revised H-mode (L-mode) yield estimates (00g, UL)

	D on Be	D on W	Ne on W (*)
$Y(\langle E \rangle, \cos \alpha \equiv 1)$	0.025 (0.029)	0.0	0.00035
$\langle Y(E, \cos \alpha \equiv 1) \rangle$	0.015 (0.016)	0.0003 (0.0010)	0.00038
$\langle Y(E, \cos \alpha) \rangle$	0.031 (0.035)	0.0007 (0.0024)	0.00080

(*) wrongly (!) assuming same spectra for Ne as for D, and $\Gamma_{\text{Ne0}} = 1\% \Gamma_{\text{D0}}$
 → Need real MC calculations for Ne!

$$\langle Y(E, \cos \alpha) \rangle = \iint Y_{phys,93}^{EIRENE}(E, \cos \alpha) f(E, \cos \alpha) dE d \cos \alpha$$

Here: $f(E, \cos \alpha) = g(E) \cdot h(\cos \alpha)$

Roth, Bohdanský, Eckstein (SPUTER93):

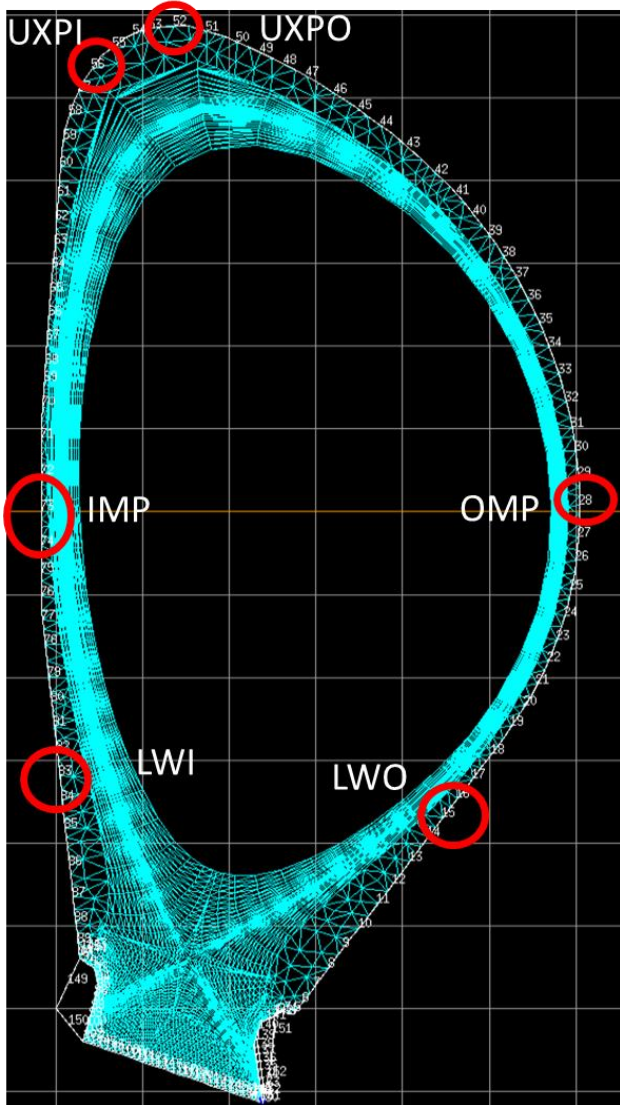
$$Y_{phys,93} = Q_p \left(1 - a^{2/3}\right) (1 - a)^2 \frac{\frac{1}{2} \ln(1 + 1.2288b)}{b + 0.1728\sqrt{b} + 0.008b^{0.1504}}$$

$$a = \frac{E_{th}}{E}, b = \frac{E}{E_{tf}}, Q_p = const$$

$$Y_{phys,93}^{EIRENE} = Y_{phys,93} \cdot (\cos \theta)^{-f} \exp\left(fc - \frac{fc}{\cos \theta}\right)$$

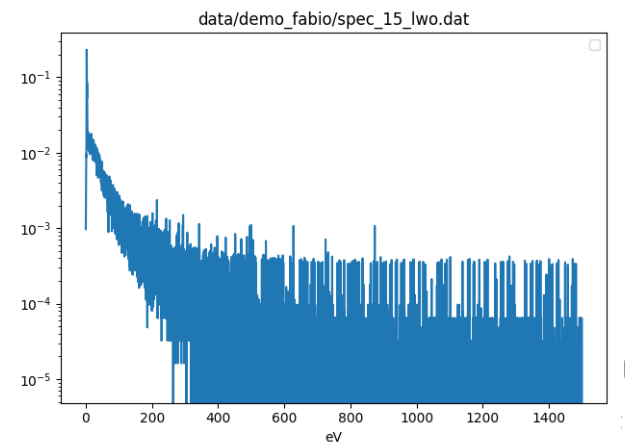
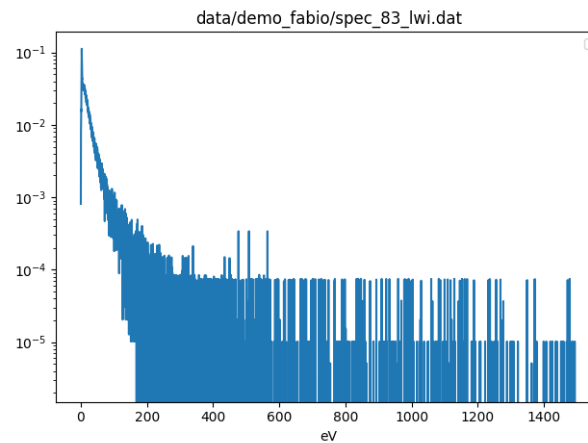
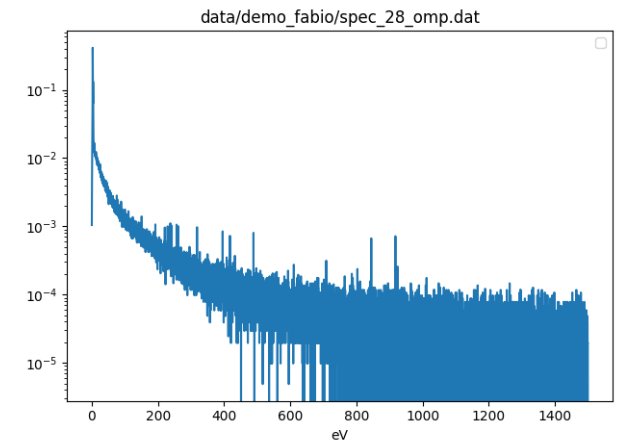
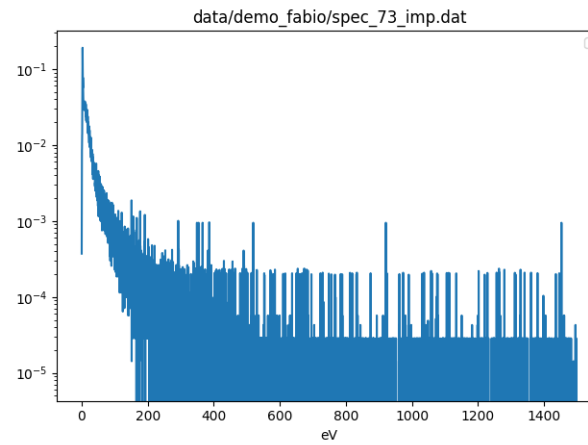
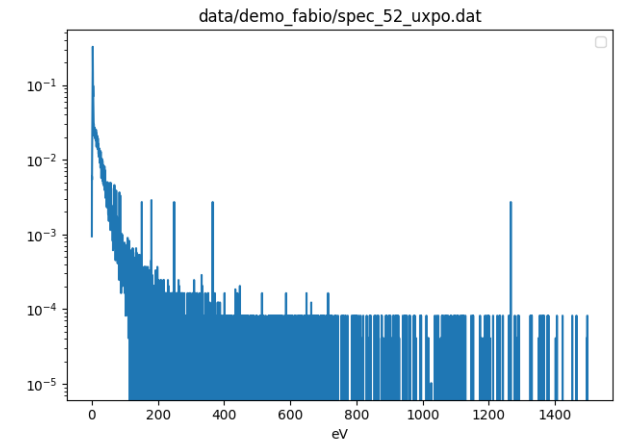
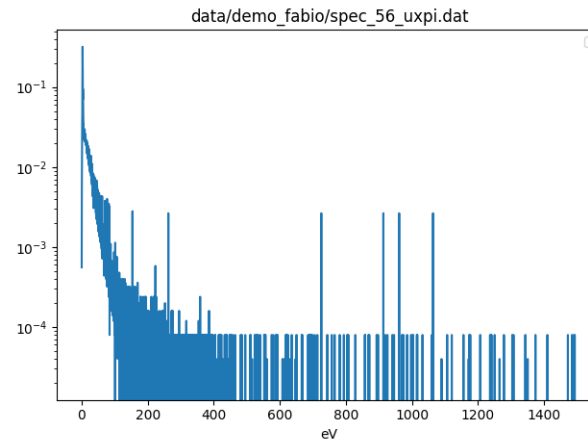
Yamamura-fit: $f = 2, c = 0.26$

EU-DEMO



Mitglied der Helmholtz-Gemeinschaft

28 November



Comments on spectral statistics

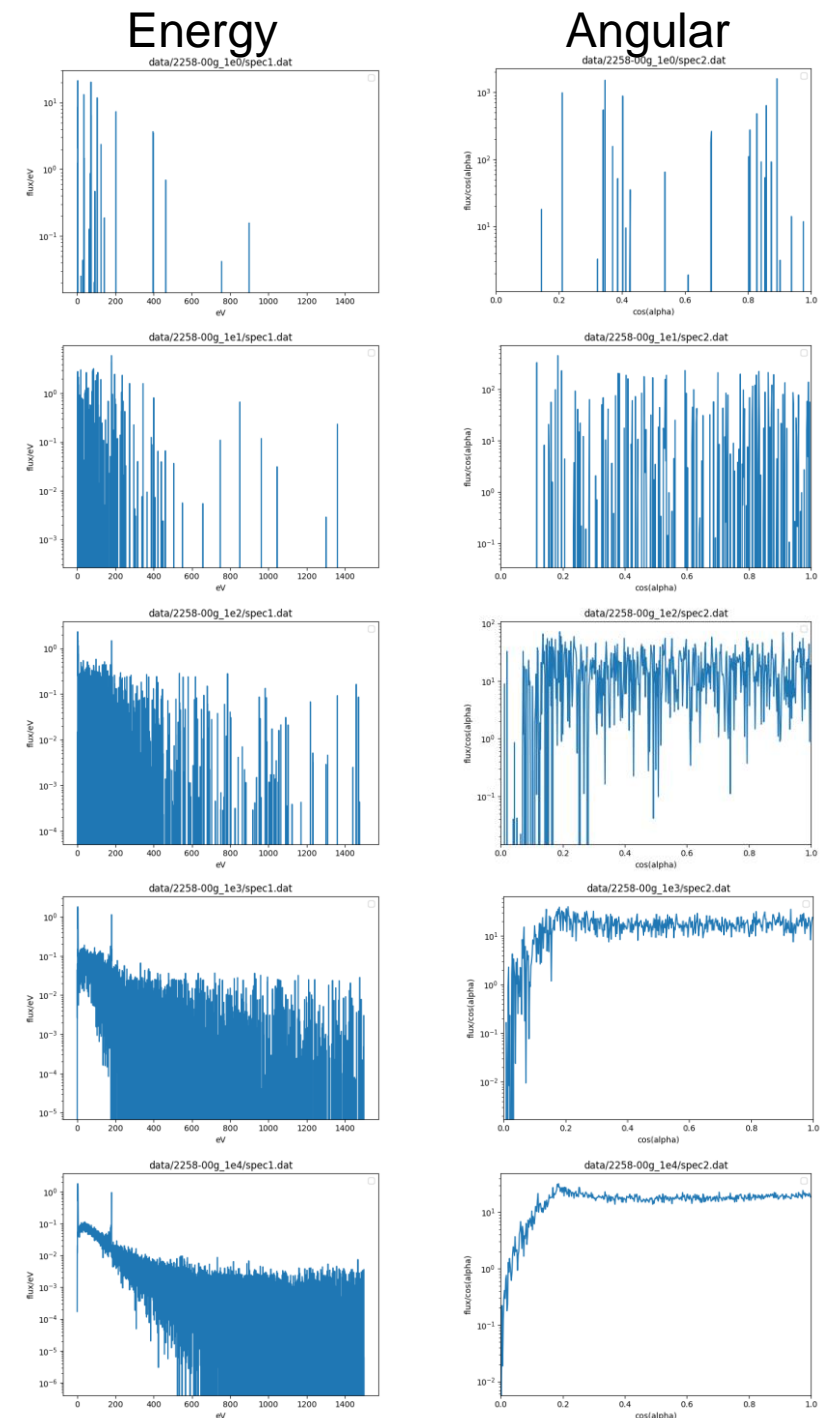
Shannon-Entropy

$$S = - \sum_k f_k \log f_k$$

MC histories	Wall-clock time (12 CPUs)	$S(g(E))$	$S(h(\cos \alpha))$
1.92e5	< 1min	2.409	2.409
1.92e6	7 min	4.097	4.143
1.92e7	71 min	5.926	5.792
1.92e8	12.5 hs	6.819	6.087
1.92e9	72 hs	7.083	6.134

OK, but Shannon-Entropy mixing invariant!
Need a better figure-of-merit for statistical accuracy

Increase in MC particle histories



Towards optimised statistics and data compression

MaxEnt-regularised approach

Maximise

$$\mathcal{L} = S + \sum_{l=1}^L \lambda_l \left(\int_{x_0}^{x_1} x^l \rho(x) dx - \mu_l \right).$$

- Requires calculation of higher moments in EIRENE

→ mean $l = 1$ & variance $l = 2$ (both done)

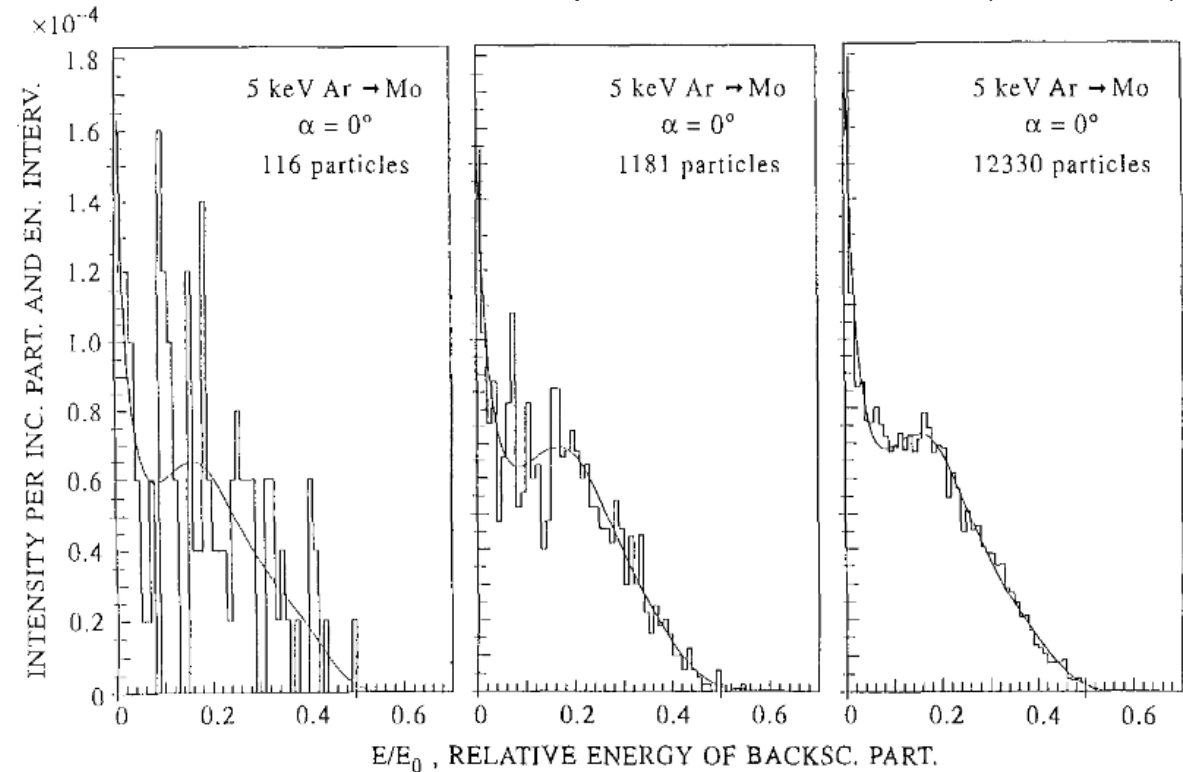
to be implemented $l > 2$:

→ skewness, kurtosis, hyper-skewness, hyper-tailness, ... → **recovers local/mixing structures**

- If statistical fluctuations < evolving structures
→ stop the EIRENE run
(implementation of a halt during run-time reqd.)

- **Bonus: moment-representation possible for spectra, data-compression rate > 100 possible**

W. Eckstein Nucl. Instr. And Meth. In Phys Res B 95 (1995)
“Efficient tabulation of plasma interaction data” (SDTrimSP)



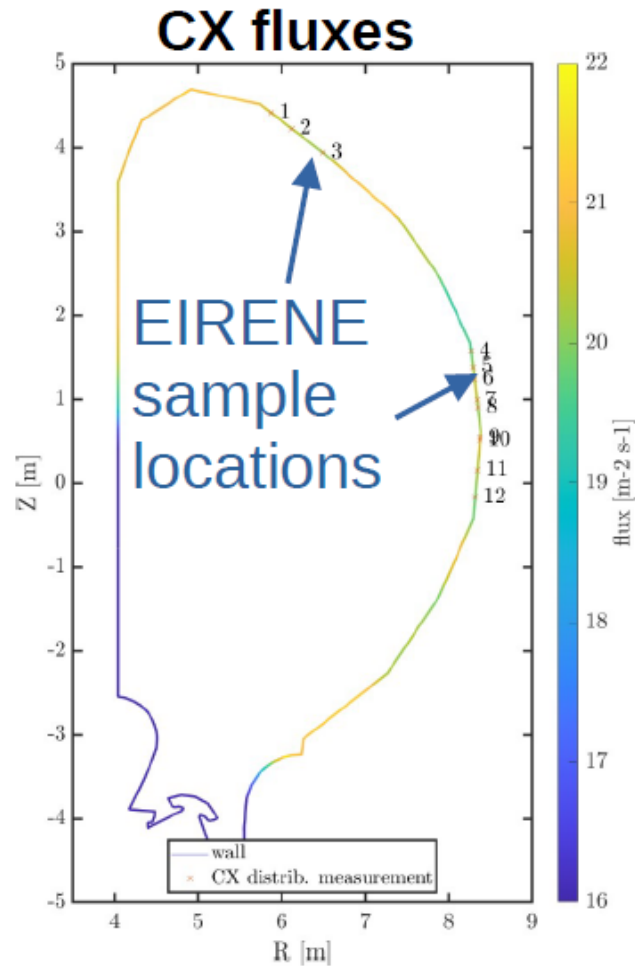
Summary / Conclusion

- Neutral particle energy and angular distributions $f(E, \cos \alpha)$ collected on diagnostic surfaces for ITER reference SOLPS-ITER plasmas with manually extended grid up to FW (A. Khan et al)
- **Result: detailed distributions give 2-3 larger D \rightarrow W sputter yields $\langle Y(E, \cos \alpha) \rangle$ compared to standard estimates $Y(\langle E \rangle)$, depends on far-SOL assumptions or H/L-mode, $\cos \alpha$ -dependence gives a factor 2, H-mode: main contribution from tail of distribution**
- Next Step: Ne \rightarrow W calculation, and compare relevance to D \rightarrow W
- SOLPS-ITER with wide-grid option should provide a better picture (IO task to provide data)
- Also: JET post-processing with EIRENE on-going (M. Groth et al), DEMO (Wiesen, Brenzke FZJ)
- So far only uncorrelated energy and angular distributions collected
 - \rightarrow extension to multi-variate distribution functions possible $f = f(E, \cos(\alpha))$
 - \rightarrow requires longer EIRENE run-times for improved statistics and requires large memory
 - \rightarrow data compression through MaxEnt regularization
- Only polar angles are collected (toroidally symmetric)
 - \rightarrow extension to full 3D possible (e.g. post-processing EMC3 plasma-backgrounds)

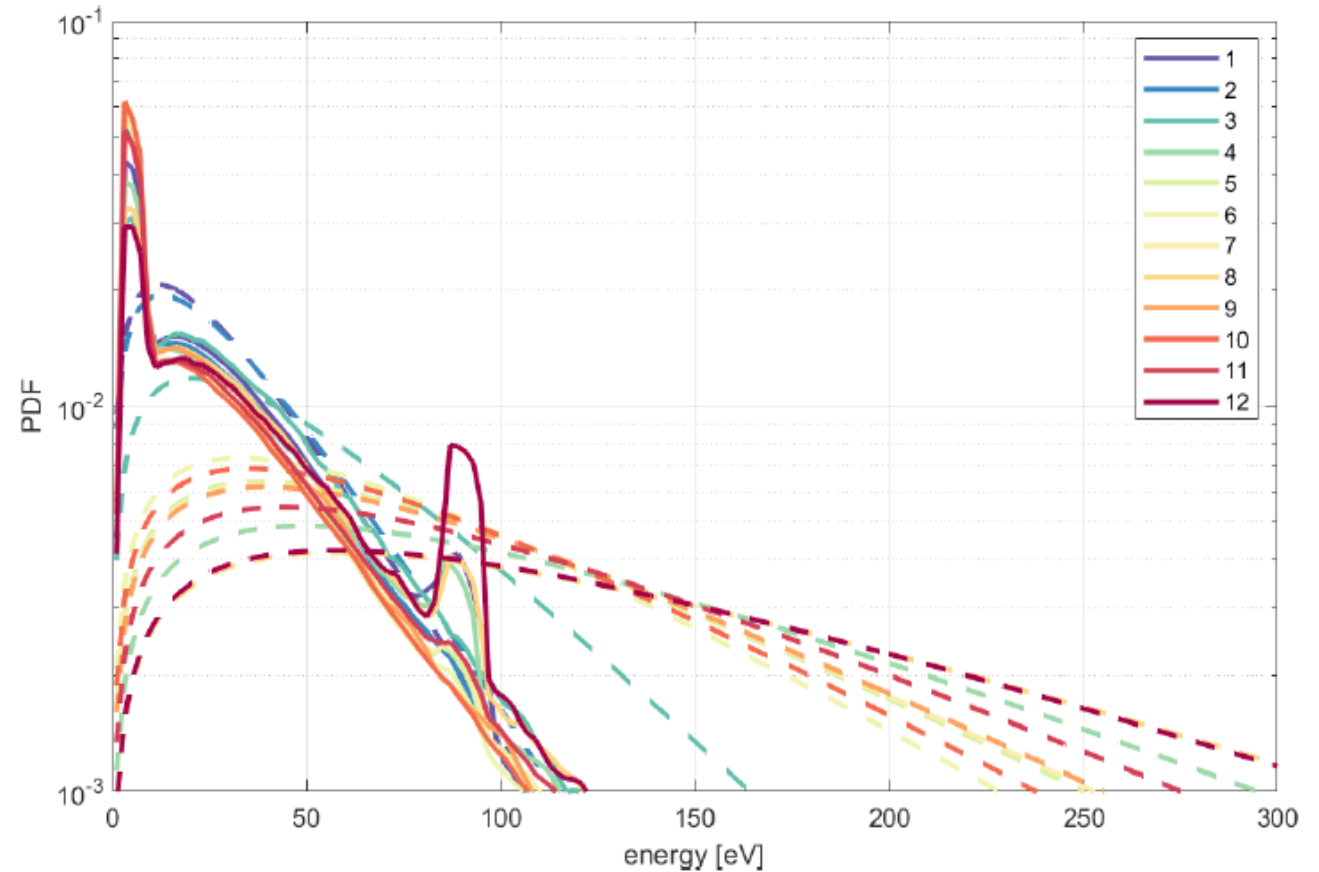
Backup

Example: CX distributions from Steve Lisgo for ITER (case #1514)

- File name “i-cxd-0003-1514-00g.wall_energy_spectra” -> I guess that’s not a standard EIRENE output…?
- Energy distributions at 12 locations, however no angular distributions.



CX energy distributions at the 12 sample locations



Modifications to the EIRENE code for angular distributions

- (Output/outspec)
- (Plotting/eirmod_plteir)
- Modules/eirmod_parmmod
Modules/eirmod_cestim → introduction of ISPCOPT flag (for angular spectra)
- Scoring/update_spectrum
- Startup_routines/input.f
Startup_routines/plasma.f
- Assistant/learca2
- Particle_tracing/folneut.f → bugfix in case of transparent srfs