Contribution ID: 23

Radiation burn-through measurements to infer opacity at conditions close to the solar radiative zone-convective zone boundary

Wednesday, 17 May 2023 15:40 (30 minutes)

Recent measurements at the Sandia National Laboratory of the x-ray transmission of iron plasma much higher than predicted by theory have cast doubt on modelling of iron x-ray radiative opacity at conditions close to the solar convective zone-radiative zone boundary. An increased radiative opacity of the solar mixture, in particular iron, is a possible explanation for the disagreement in the position of the solar convection zoneradiative zone boundary as measured by helioseismology and predicted by modelling using the currently accepted elemental composition based on photosphere analysis. Here we present data from radiation burnthrough experiments which do not support a large increase in the opacity of iron at conditions close to the base of the solar convection zone and provide a constraint on the possible values of both the mean opacity and the opacity in the x-ray range of the Sandia experiments. the data agree with opacity values from current state-of-the-art opacity modelling and hence do not support an increase in iron opacity as a partial explanation for the discrepancy in the predicted solar convective-radiative zone boundary position.

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Session Classification: Astrophysical Plasmas

Track Classification: Astrophysical Plasmas