

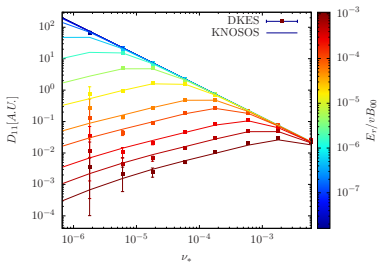
KNOSOS, a fast neoclassical code for three-dimensional magnetic configurations

New version of the code: the bounce-averaged drift kinetic equation implemented is

$$\int_{l_{b1}}^{l_{b2}} \frac{dl}{|v_{\parallel}|} \left(\mathbf{v}_{M,b} + \frac{B}{\langle B \rangle} \mathbf{v}_E \right) \cdot \nabla \alpha (\partial_{\alpha} + \partial_{\alpha} \lambda |J| \partial_{\lambda}) g_b - \int_{l_{b1}}^{l_{b2}} \frac{dl}{|v_{\parallel}|} C_b^{\text{lin}}[g_b] = - \int_{l_{b1}}^{l_{b2}} \frac{dl}{|v_{\parallel}|} (\mathbf{v}_{M,b} + \mathbf{v}_E) \cdot \nabla \psi \Upsilon_b F_{M,b}.$$

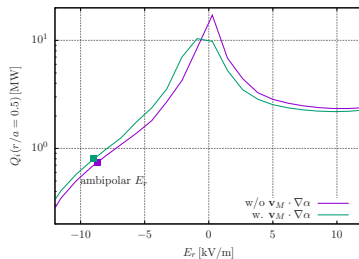
More accurate and/or faster than standard neoclassical codes in two limits:

(1) Large aspect ratio stellarators
(and standard values of E_r)



Benchmarked against DKES, orders of magnitude faster \Rightarrow KNOSOS will be used for stellarator optimization.

(2) Optimized stellarators
(relevant when E_r is small)



Tangential magnetic drift needed for correct calculation of energy flux \Rightarrow KNOSOS will be used for analyzing W7-X experiments.