



This project has received funding from the European Union's Horizon 2020 research and innovation programme under grant agreement No 101008324 (ChETEC-INFRA).



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Nuclear Data for Astrophysics - ChETEC-INFRA



IAEA Consultancy Meeting
April 25, 2023

Axel Boeltzig
HZDR Dresden, Germany

Institute of Radiation Physics · Nuclear Physics Division · Dr. Axel Boeltzig · a.boeltzig@hzdr.de / chetec-infra@hzdr.de · www.hzdr.de

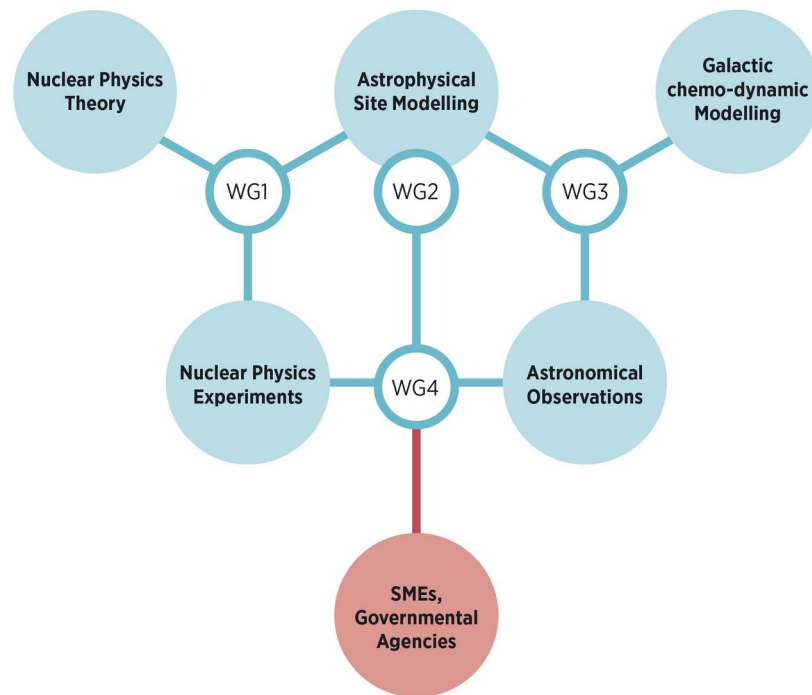
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concept



COST action ChETEC [ketek] on Nuclear Astrophysics

Chemical Elements as Tracers of the Evolution of the Cosmos

A network to bring European research, science and business together to further our understanding of the early universe



<http://www.chetec.eu>

<http://nebula.kis.keele.ac.uk:8080/chetec>

- ◆ ~160 k€/year 2017-2021
- ◆ 30 European countries

Chair:

- ◆ Raphael Hirschi, Keele/UK



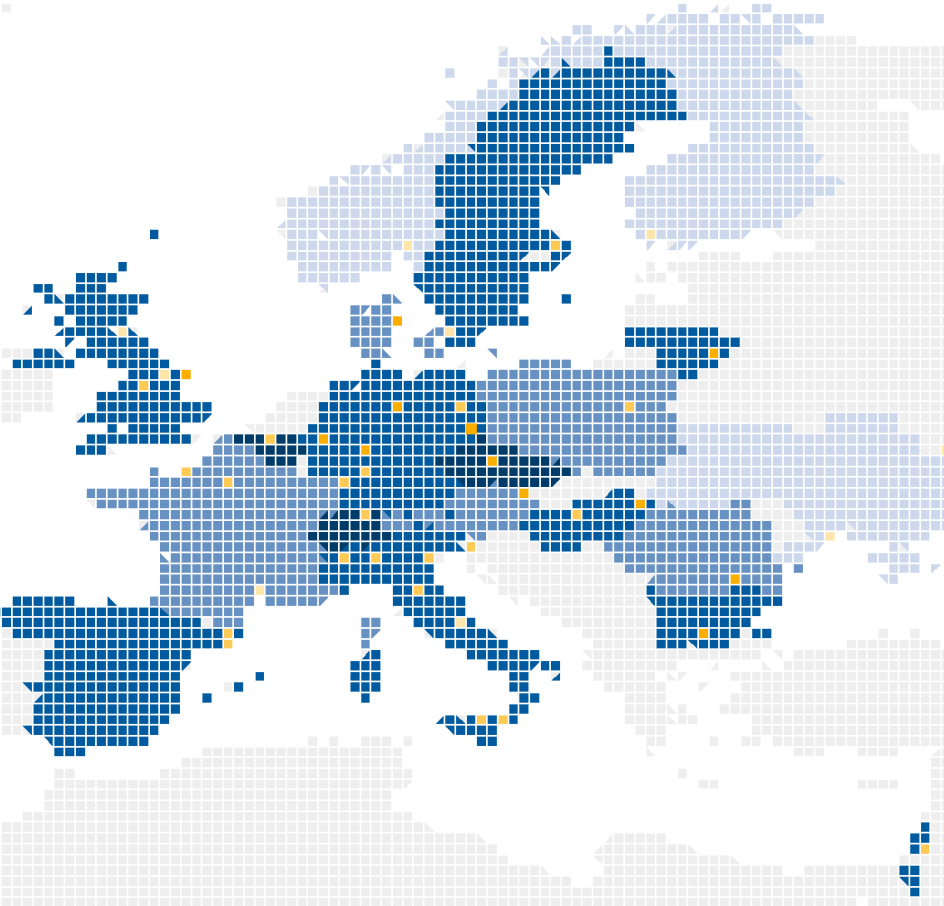
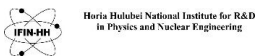
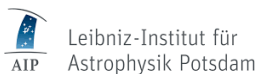
ChETEC-INFRA project for nuclear astrophysics [ketek-infra]

- ◆ EU Horizon 2020 **Starting Community** of research infrastructures to serve nuclear astrophysics
- ◆ H2020-INFRAIA-2020-1
- ◆ **32 partners** in 17 EU+ countries
- ◆ 1 May 2021 – 30 April 2025
- ◆ 5.0 M€ support by EU
- ◆ **13 research infrastructures** offer EU-supported transnational access, selection based on scientific merit



<https://www.chetec-infra.eu>

32 partners in ChETEC-INFRA



EU-supported Starting Community of Research Infrastructures for Nuclear Astrophysics

5.0 M€ EU HORIZON2020 support (2021-2025)

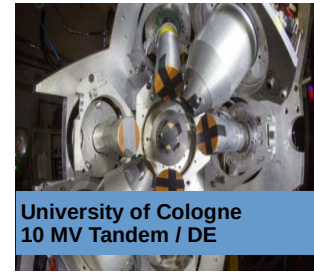
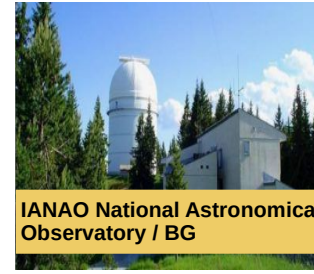
TA Transnational Access	JRA Joint Research Activities	NA Networking Activities
Infrastructure access <ul style="list-style-type: none">• 8 nuclear laboratories• 4 telescopes• 1 computing cluster	Infrastructure usability <ul style="list-style-type: none">• Targets• Neutron detectors• Abundance corrections• Analysis pipelines	Infrastructure networking <ul style="list-style-type: none">• Complementary Experiments• Solar fusion+model• Geochemistry/Astrophysics• Nuclear astrophysics schools• Outreach

32 partners, 17 countries, largest EU project for nuclear astrophysics yet

WP6 / NA1: “Comprehensive Nuclear Astrophysics”
PI: Jordi José, UPC

WP8 / NA3: “Astronuclear Library”
PI: Aldo Serenelli, ICE-CSIC

Transnational Access to 13 Infrastructures via ChETEC-INFRA



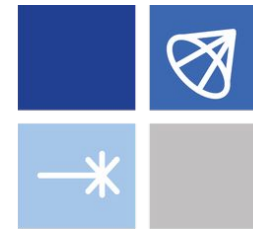
laboratories
supercomputers
telescopes



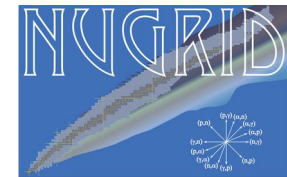
Nuclear Astrophysics Networks Worldwide



IRENA is a National Science Foundation AccelNet Network of Networks



JINA-CEE



Nuclear Astrophysics

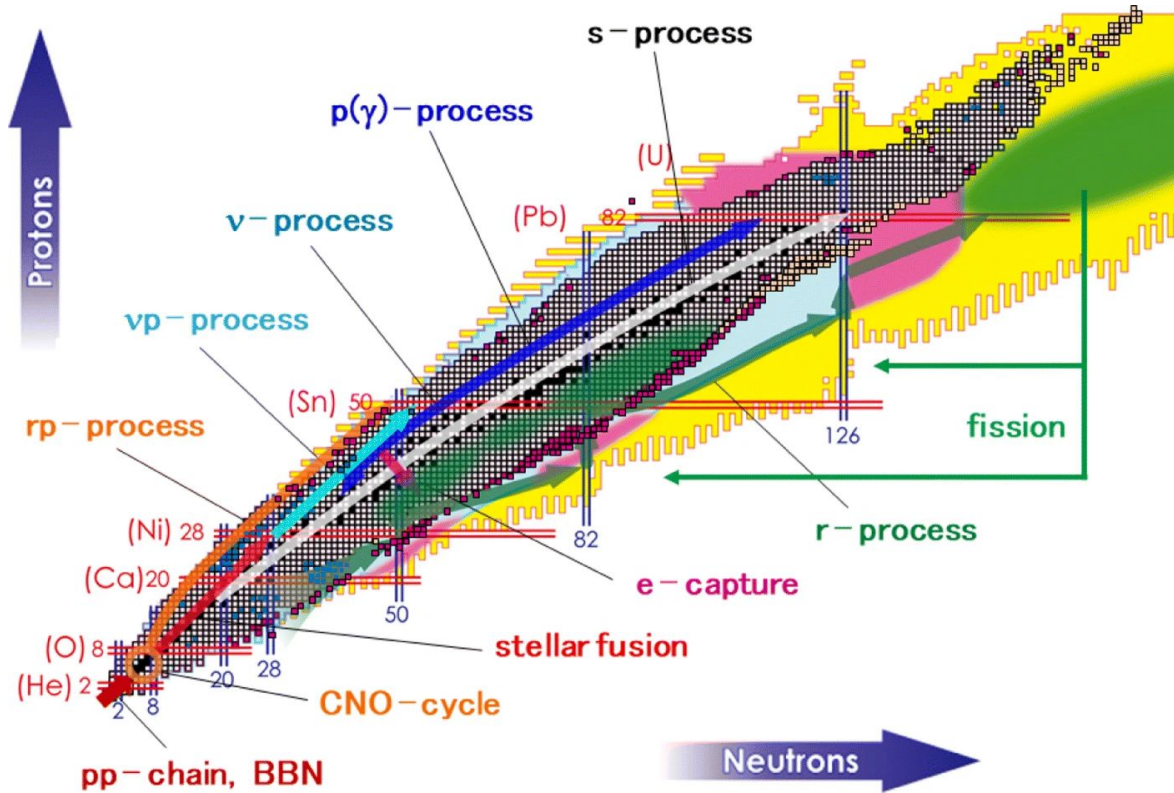
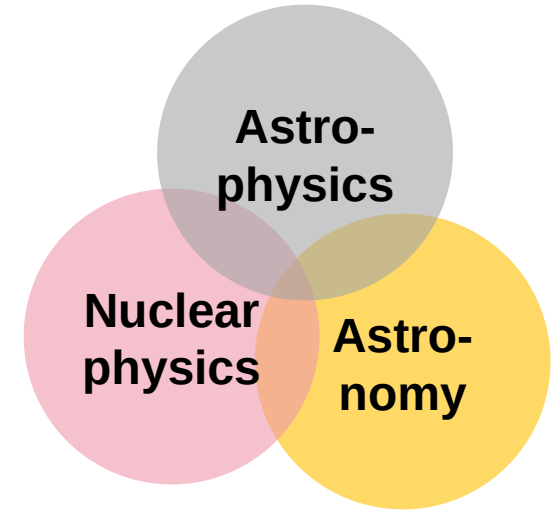
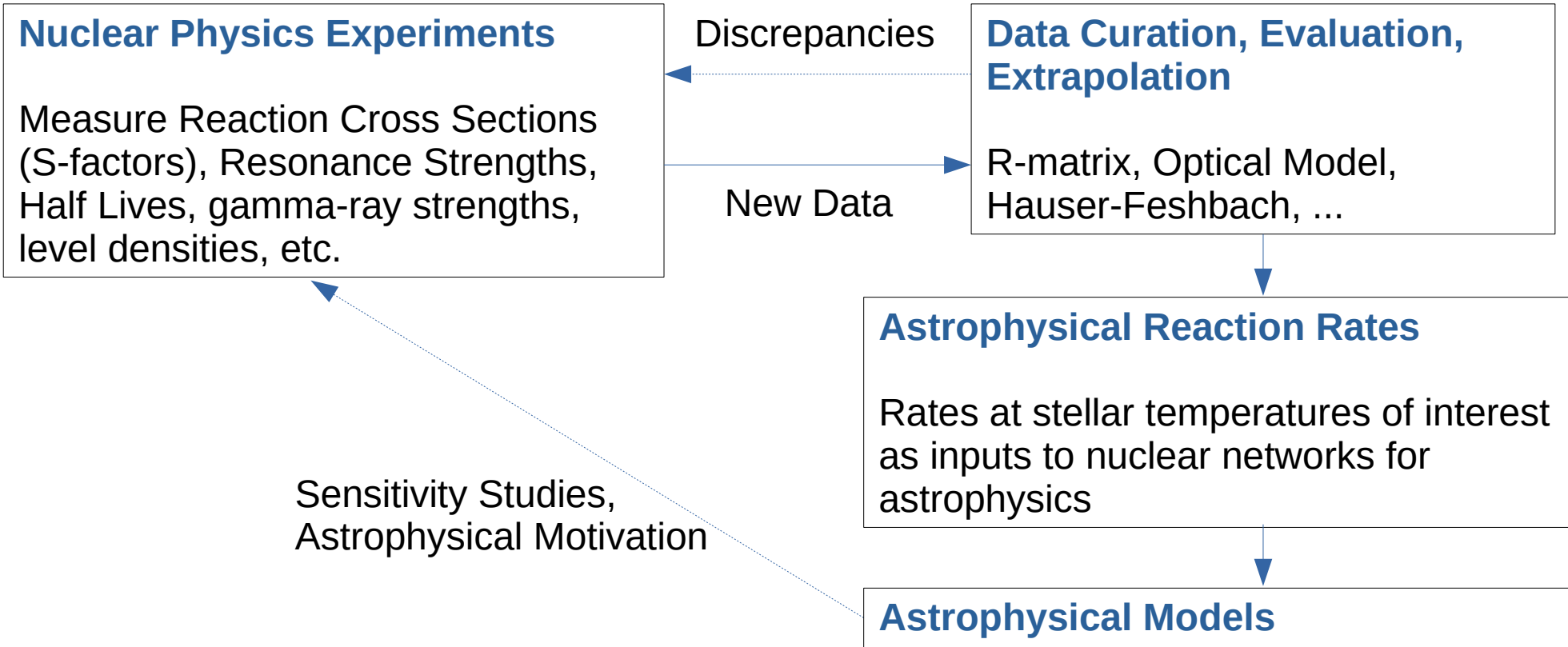


Figure: Aziz et al., 2022 (10.1007/s43673-021-00018-z)

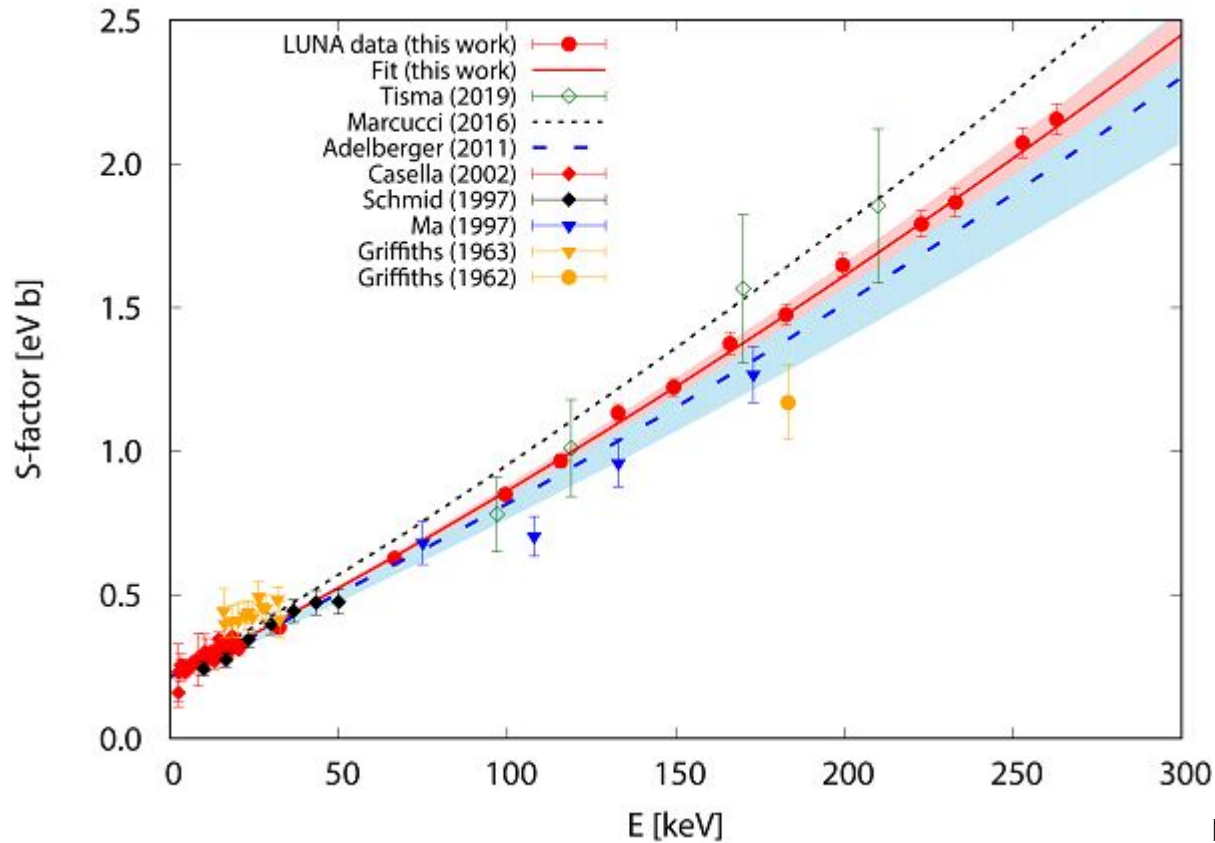


Interdisciplinary field to study the the chemical evolution of the elements in the Universe

Nuclear Astrophysics



Examples for Data Sets



${}^2\text{H}(p,g){}^3\text{He}$

Figure: Mossa et al., 2020

Examples for Data Sets

$^{12}\text{C}(\text{a},\text{g})^{16}\text{O}$

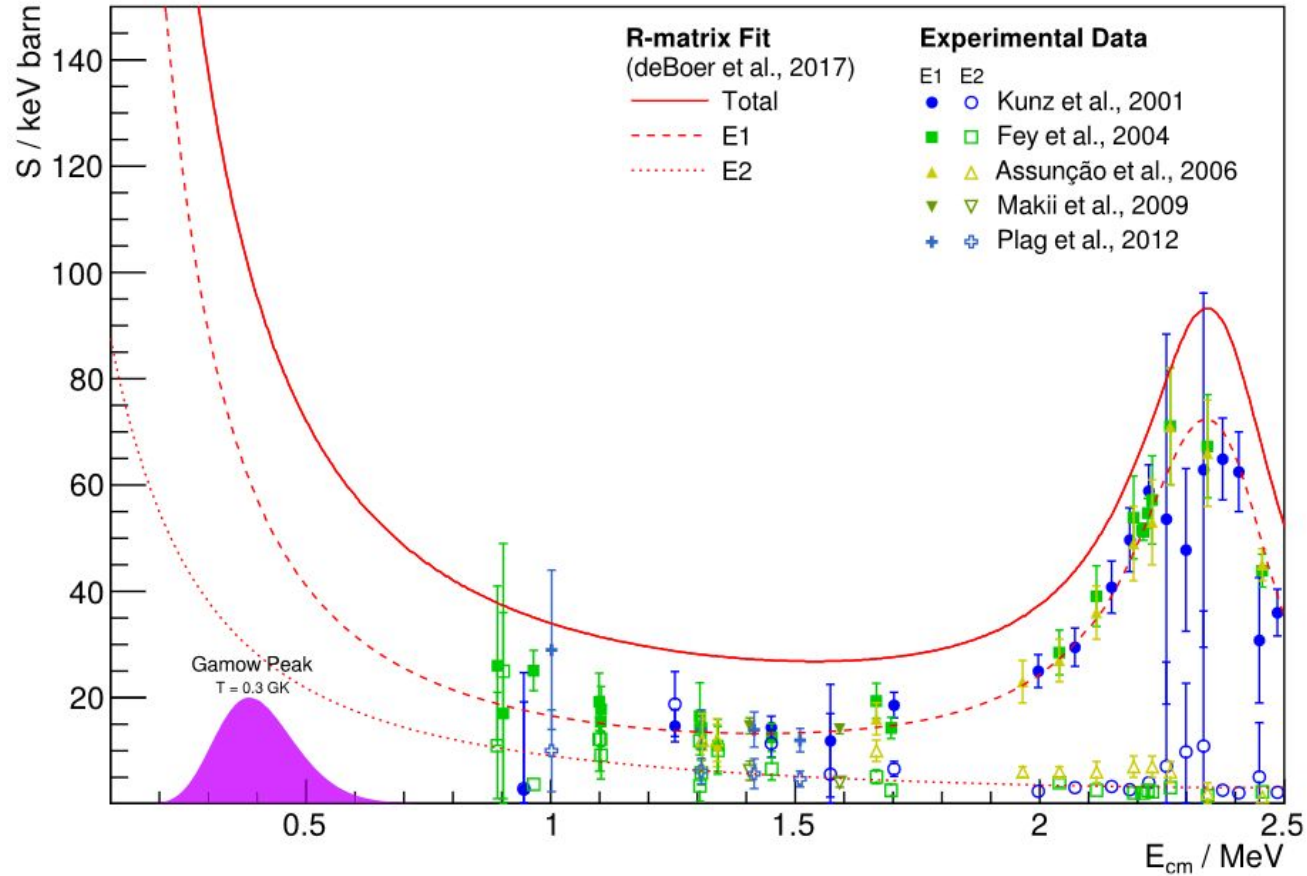


Figure: Ferraro et al., 2021

Examples for Data Sets

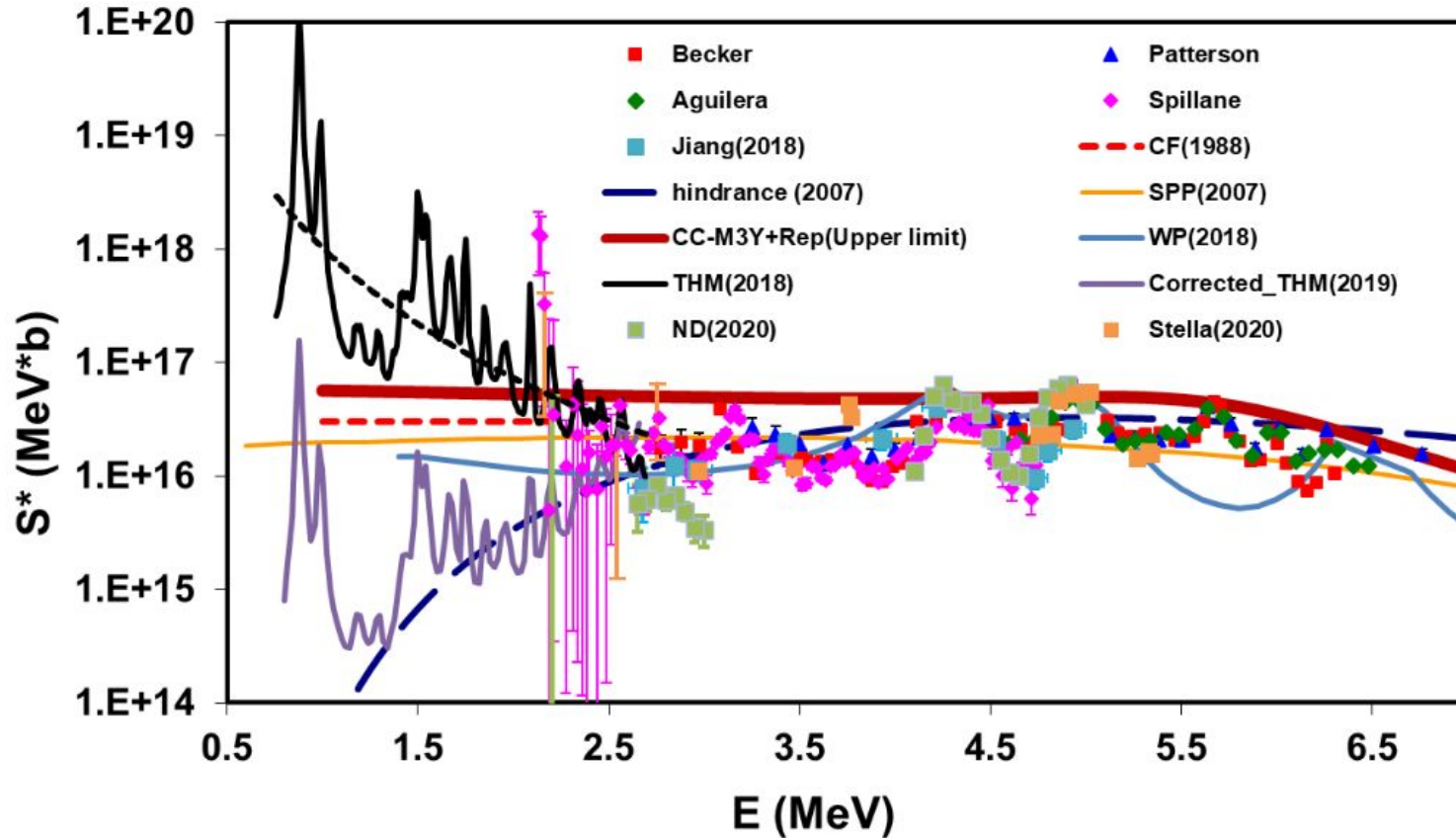


Figure: Tang et al., 2022

Nuclear Library for Astrophysics Applications



Energy Levels of Light Nuclei, A = 3 - 20

Nuclear Data Evaluation Project Triangle Universities Nuclear Laboratory

TUNL Nuclear Data
Evaluation Home
Page

Information on
mass
chains and
nuclides

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19	20

Group Info
Publications
HTML
General Tables
Level Diagrams
Tables of EL's
NSR Key# Retrieval
ENSDF
Excitation Functions

- [TUNL Nuclear Data Group](#): Who we are and what we do.

Our publications on Energy Levels of Light Nuclei, A = 5 - 20:



- **Publications**: TUNL evaluations of A = 3 - 20, and modified versions of Fay Ajzenberg-Selove's publications of A = 5 - 20, are available here in PDF format. The most recent HTML documents of A = 3 - 20, and EL diagrams of A = 4 - 20 are also available here. Some reprints and preprints may be requested by mail.
- **HTML for Nuclides**: HTML documents are available for individual nuclides found within the TUNL or FAS evaluations.

Resources relating to our publications:

- [Energy Level Diagrams](#) are available for A = 4 - 20 nuclides.
- [Tables of Energy Levels](#): a brief listing of tables of energy levels from the most recent publication for each nuclide A = 4 - 20.
- [SiteMap and Complete List of Available TUNL Documents](#): Trying to find a specific TUNL evaluation or preliminary report, HTML document, General Table, Update List or Energy Level Diagram? Click here for a complete list of what's available on our website.

Applications and databases relating to the A = 3 - 20 nuclides:

- [NSR Key Number Retrieval](#)
- [ENSDF](#): Information for A = 2 - 20 nuclides available through the National Nuclear Data Center (NNDC) site.
- [Excitation Functions](#): Compilation of the excitation functions for various (p, X) and (α , X) reactions.
- [Thermal Neutron Capture Data](#): Summary of level and branching intensity data measured in Thermal Neutron Capture.
- [Ground-State Decay Data](#): Summary of half-life, branching intensity, and mass excess data measured in ground state beta- and charged-particle-decay.
- [Half-Lives Table](#): List table of nuclear decay half-lives.
- [Live Chart of Nuclides - IAEA](#): Allows to search and plot nuclear structure and nuclear decay data interactively.
- [AME-Viewer](#): Java utility for interacting and comparing atomic mass data from the Atomic Mass Data Center 2020 mass evaluation.

Helpful links:

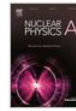
<https://nucldata.tunl.duke.edu/>

Nuclear Astrophysics Compilation of Reaction Rates (NACRE)



Nuclear Physics A

Volume 656, Issue 1, 23 August 1999, Pages 3-183



A compilation of charged-particle induced thermonuclear reaction rates

C. Angulo^{a,1}, M. Arnould^a, M. Rayet^{a,2}, P. Descouvemont^{b,3}, D. Baye^b, C. Leclercq-Willain^b, A. Coc^c, S. Barhoumi^{c,4}, P. Aguer^{c,5}, C. Rolfs^d, R. Kunz^e, J.W. Hammer^e, A. Mayer^e, T. Paradellis^f, S. Kossionides^f, C. Chronidou^f, K. Spyrou^f, S. Degl'Innocenti^{g,6}, G. Fiorentini^g, B. Ricci^{g...}, M. Laméhi Rachti^{j,10}

Show more ▾

(86 reactions)

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[https://doi.org/10.1016/S0375-9474\(99\)00030-5](https://doi.org/10.1016/S0375-9474(99)00030-5)

[Get rights and content](#)



Nuclear Physics A

Volume 918, 20 November 2013, Pages 61-169



NACRE II: an update of the NACRE compilation of charged-particle-induced thermonuclear reaction rates for nuclei with mass number $A < 16$

Y. Xu^{a,1}, K. Takahashi^{a,b}, S. Goriely^a, M. Arnould^a , M. Ohta^{c,d}, H. Utsunomiya^d

Show more ▾

(34 reactions)

+ Add to Mendeley Share Cite

<https://doi.org/10.1016/j.nuclphysa.2013.09.007>

[Get rights and content](#)

NACRE (1999) and NACRE II (2013) still widely used as baseline for reaction rates

<http://www.astro.ulb.ac.be/nacreii/>

BRUSLIB Database for Nuclear Astrophysics Applications

Free Access

Issue A&A
Volume 549, January 2013

Article Number A106

Number of page(s) 10

Section Atomic, molecular, and nuclear data

DOI <https://doi.org/10.1051/0004-6361/201220537>

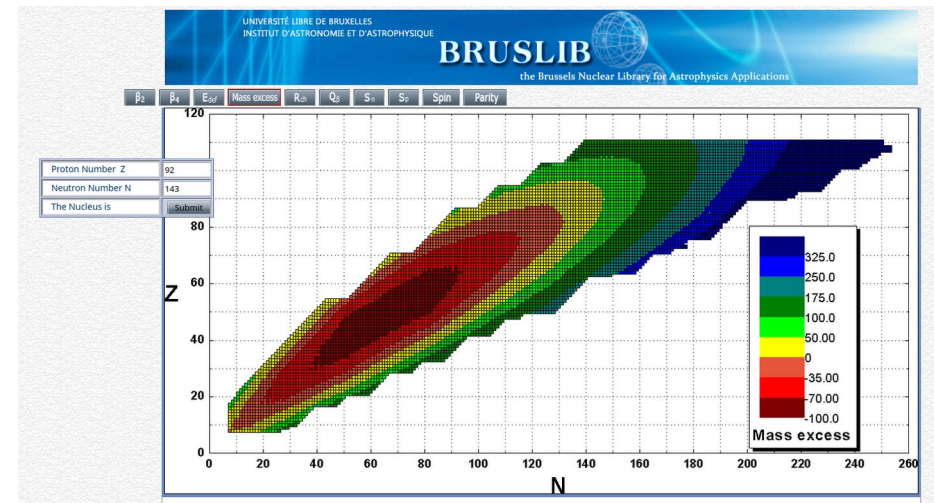
Published online 07 January 2013

A&A 549, A106 (2013)

Databases and tools for nuclear astrophysics applications

BRUSsels Nuclear LIBrary (BRUSLIB), Nuclear Astrophysics Compilation of REactions II (NACRE II) and Nuclear NETwork GENERator (NETGEN)

Y. Xu, S. Goriely, A. Jorissen, G. L. Chen and M. Arnould



<http://www.astro.ulb.ac.be/bruslib/>

Solar Fusion I & II

Solar fusion cross sections

Eric G. Adelberger *et al.*

Rev. Mod. Phys. **70**, 1265 – Published 1 October 1998

Article References Citing Articles (501) PDF Export Citation

>

ABSTRACT

We review and analyze the available information on the nuclear-fusion cross sections that are most important for solar energy generation and solar neutrino production. We provide best values for the low-energy cross-section factors and, wherever possible, estimates of the uncertainties. We also describe the most important experiments and calculations that are required in order to improve our knowledge of solar fusion rates.

DOI: <https://doi.org/10.1103/RevModPhys.70.1265>

Solar Fusion (1998)

Community-led evaluations, resulting in evaluated reaction rates

Solar fusion cross sections. II. The pp chain and CNO cycles

E. G. Adelberger *et al.*

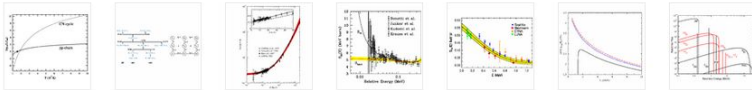
Rev. Mod. Phys. **83**, 195 – Published 12 April 2011

Article References Citing Articles (515) PDF HTML Export Citation

>

ABSTRACT

The available data on nuclear fusion cross sections important to energy generation in the Sun and other hydrogen-burning stars and to solar neutrino production are summarized and critically evaluated. Recommended values and uncertainties are provided for key cross sections, and a recommended spectrum is given for ${}^8\text{B}$ solar neutrinos. Opportunities for further increasing the precision of key rates are also discussed, including new facilities, new experimental techniques, and improvements in theory. This review, which summarizes the conclusions of a workshop held at the Institute for Nuclear Theory, Seattle, in January 2009, is intended as a 10-year update and supplement to 1998, *Rev. Mod. Phys.* **70**, 1265.



13 More
Received 21 April 2010

DOI: <https://doi.org/10.1103/RevModPhys.83.195>

Solar Fusion II (2011)

Solar Fusion III

Solar Fusion Cross Sections III (INT-22-82W)

26-29 July 2022

David Brower Center and UC Berkeley, Berkeley, CA, USA

US/Pacific timezone



* WG1 - S_{11} [$^1\text{H}(p, e^+ \nu)^2\text{H}$] and theory, extrapolations, radiative corrections, electron capture, neutrino spectra

Conveners: L. Marcucci (Univ. Pisa), J.-W. Chen (Natl. Univ. Taiwan)

* WG2 - Solar neutrino observations, helioseismology, asteroseismology

Conveners: F. Villante (Univ. L'Aquila), A. Serenelli (ICE), H. Robertson (Univ. Washington)

* WG3 - S_{114} [$^{14}\text{N}(p, \gamma)^{15}\text{O}$]

Conveners: R. James deBoer (Notre Dame U.), Tamás Szücs (ATOMKI)

* WG4 - S_{12} [$^2\text{H}(p, \gamma)^3\text{He}$]

Conveners: F. Cavanna (INFN Torino)

* WG5 - S_{17} [$^7\text{Be}(p, \gamma)^8\text{B}$]

Conveners: L. Gialanella (Univ. Studi Campania)

* WG6 - S_{112} [$^{12}\text{C}(p, \gamma)^{13}\text{N}$], other CNO, NeNa

Conveners: R. Depalo (Univ. Milano), Ch. Iliadis (Univ. North Carolina)

* WG7 - S_{34} [$^3\text{He}(^4\text{He}, \gamma)^7\text{Be}$] and S_{33} [$^3\text{He}(^3\text{He}, 2p)^4\text{He}$]

Conveners: D. Bemmerer (HZDR), K. Nollett (San Diego State)

* WG8 - New facilities: LUNA-MV, JUNA, NIF, etc

Conveners: M. Aliotta (Univ. Edinburgh), A. Tumino (Enna Univ.), M. Wiescher (Notre Dame Univ.)

* WG9 - Opacity, screening

Conveners: T. Nagayama (Sandia), J. Colgan (Los Alamos), B. Balantekin (UW-Madison)

STARLIB: Thermonuclear Rate Library

THE ASTROPHYSICAL JOURNAL
SUPPLEMENT SERIES



ARTICLES

STARLIB: A NEXT-GENERATION REACTION-RATE LIBRARY FOR NUCLEAR ASTROPHYSICS

A. L. Sallaska¹, C. Iliadis^{2,3}, A. E. Champagne^{2,3}, S. Goriely⁴, S. Starrfield⁵, and F. X. Timmes⁵

Published 2013 July 2 • © 2013. The American Astronomical Society. All rights reserved.

[The Astrophysical Journal Supplement Series, Volume 207, Number 1](#)

Citation A. L. Sallaska *et al* 2013 *ApJS* **207** 18

DOI 10.1088/0067-0049/207/1/18

<https://starlib.github.io/Rate-Library/>
<https://github.com/Starlib/Rate-Library>

The JINA Center for the Evolution of the Elements
REACLIB Database
you are not logged in | [login] [sign up]

- JINA-CEE
- Homepage
- About
- Status/Discussion
- Search/View
- Libraries
- Help
- Links

Welcome to the JINA ReacLib Database!

The reacLib data base has moved to a new home on jinaweb.org - Welcome! This move and significant recoding was necessary to satisfy security requirements. Thank you for your patience during the transition process. Most of the functionality has been restored except for the graphical display of reaction rates, which is still under construction. If you see any other issues please let us know.

This is a database for nuclear reaction rates to be used in astrophysical model calculations. To get help please visit the [help page](#).

For more details and REACLIB citation, see [Cyburt et al., ApJS 189 \(2010\) 240](#).

Database:

Contains multiple versions of each rate with one recommended rate. One can find rates using our [search engine](#) or by typing in the specific reaction URL (e.g. [http://jinaweb.org/reacLib/db/na21\(p.g\)](http://jinaweb.org/reacLib/db/na21(p.g)))

Libraries:

Libraries are collections of rates selected from the database. Public reacLib releases are available as libraries and represent snapshots of the database on a particular date. Users can also define their own libraries and make them public to document the rates they used.

Account:

You are not required to have an account to browse and download data. You will only need an account to create and manage your own libraries. To create your own account please click [sign up](#).

Contact: For further information, direct your inquiries to reacLib at nsc1 dot msu dot edu.

Notes:

- How are reverse rates calculated? or What does the "r" flag mean?
 - The "r" flag in REACLIB format, identifies "reverse" rates
 - "Reverse" rates are calculated via detailed balance (without partition functions)
 - Therefore, rates with this flag must be corrected to include partition function modifications.
- On Al26 ground and meta stable states in REACLIB.
 - Al-6 refers to the ground state of Al26.
 - Al*6 refers to the meta-stable state of Al26.
 - Al26 refers to the thermal population of Al-6 & Al*6.
- On the ground state weak decays under the labels wc07 & mo03.
 - these weak rates are the partial decay rates, not total rates.
 - beta-delayed particle emission is presented in the database.
- On the pp-d weak reactions.
 - the 2 rates presented are labeled "ec" and "bet+"
 - they must be accounted for specially (as they aren't beta-decays or electron captures).

News:

Updated default libra
2021-06-24 Submitted I Sneed
The default library was with 46 new rates today access the DAT file [here](#)

ReacLibV2.2
2016-11-14 Submitted I Sneed
In reacLib v2.2, we fix reverse rate issues. This is now available.

Server Maintenance
2015-02-20 Submitted By: Richard Cyburt
Expect some outages 2-20-2015 due to server maintenance.

REACLIB V2.0
2012-06-01 Submitted By Cyburt

After much effort, the R Team would like to announce the release of Snapshot library V2.0. The rates updated can be found on the Status page. Thank you for your helpful comments.

New Interface
2010-12-15 Submitted By Cyburt
The REACLIB Team is pleased to announce the release of the new web interface. If there are any questions/comments on the interface or database, please contact us via our [Email](#).

ReacLibV0.5

Viewing rates 1-50 of 76064

Page 1 of 1522. 1 2 3 4 5 6 7 ... 1522

Chap	Reaction	Type	Label	Modified	Last Change	Ver	Comments
1	n -> p	exp	wc07	2010-06-11 15:14:12	script	1	Added with weakrates.pl
1	t -> he3	exp	wc07	2010-06-11 15:14:12	script	1	Added with weakrates.pl
1	he3 -> t	exp	ec	2010-06-11 15:14:12	script	1	Updated Q value
1	he6 -> li6	exp	wc07	2010-06-11 15:14:12	script	1	Added with weakrates.pl
1	he8 -> li8	exp	wc07	2010-06-11 15:14:12	script	1	Added with weakrates.pl
1	li9 -> be9	exp	wc07	2010-06-11 15:14:12	script	1	Added with weakrates.pl
1	li11 -> be11	exp	wc07	2010-06-11 15:14:12	script	1	Added with weakrates.pl
1	be7 -> li7	exp	ec	2010-06-11 15:14:12	cyburt	2	

Original Publication 2010
Updates 2016, 2021

ChANUREPS

ChANUREPS

ChETEC AstroNUclear REPositories

ChANUREPS is a platform where members of the nuclear-astrophysics community provide new nuclear reaction rates to make them easily findable, open source and in a simple format. Such rates could be used for many research tasks, e.g nucleosynthesis calculations, nuclear sensitivity studies, comparison when new rates are becoming available and much more.

<http://chanureps.chetec-infra.eu>



Collection of reaction rates provided by the community.

No evaluated “best” rate, rather “pick & choose”.

Recently launched and growing (~20 rates).

Neutron-Induced Reactions: MACS – KaDONIS

Karlsruhe Astrophysical Database of Nucleosynthesis in Stars

s-process

[Standards]

[Logbook]

[FAQ]

[Links]

[Contact]

p-process

The new version KADoNIS v0.3 is finally online!

Version 0.3 provides data for 357 isotopes including 5 newly added isotopes, 42 updated MACS30, new stellar enhancement factors, and the MACS30 obtained from three different evaluated data libraries. More information below or in the logbook.



View Maxwellian-Averaged (n,g) Cross Section

Isotope

(Examples: Ba138, Ta180m, Se,)



KADoNIS v0.3

The KADoNIS project is an online database for cross sections relevant to the **s process** and **p process**. The respective s-process library provided on this webpage is an updated sequel of the well-established Bao et al. compilation [1].

<https://www.kadonis.org>


KADoNIS-p database

The KADoNIS project is an online database for cross sections relevant to the **s-process** and **p-process (y-process)**. The present p-process library includes all available experimental data from (p,y), (p,n), (α,y), (α,n), and (α,p) reactions between ^{70}Ge and ^{209}Bi in or close to the respective Gamow window.

View datasets

Isotope

(Ex.: ^{70}Ge , Ba138, Se;
or leave it empty for the whole list)



The basis of the new p-process database is the Experimental Nuclear Reaction Data (EXFOR) database, which has the advantage that it contains (almost) all available experimental cross sections and is regularly updated. However, it contains also a lot of irrelevant data for the p process, because most measurements were performed far away from the Gamow window and have no astrophysical importance. Therefore we restrict our database to energies of maximum 1.5 times the upper end of the Gamow window at 3 GK for the respective reactions. This upper cut-off energy was arbitrarily chosen because most available data was measured above the astrophysically important Gamow windows, which lie (for T=3 GK) between E= 3 and 6 MeV for proton-induced and beta and 14 MeV for α-induced reactions. In the present database we thus have 46 (p,y), >150 (p,n), 23 (α,y), 7 (α,p), but no (p,α) reaction.



Nuclear Data Sheets
Volume 120, June 2014, Pages 171-174



The Karlsruhe Astrophysical Database of Nucleosynthesis in Stars Project – Status and Prospects

L. Dillmann^{a,b}, I. Szücs^c, R. Plag^{d,b}, Z. Fülöp^c, F. Käppeler^e, A. Mengoni^f, T. Rauscher^{g,h}

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<https://doi.org/10.1016/j.nds.2014.07.038>

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Neutron-Induced Reactions: MACS – ASTRAL

ASTRAL

ASTrophysical Rate and rAw data Library

[Home](#)

[Internal](#)

View Maxwellian-Averaged Cross Section

Isotope

(Examples: Ba138, Ta180m, Se.)

175 isotopes found in database.

Download table of ASTRAL MACS (1 line per isotope)

Kind of reaction: Release version:

kT >= keV (leave open for full range)

kT <= keV (leave open for full range)

ASTRAL Releases

Version:

[Experimentelle Astrophysik](#) | [Goethe Universität Frankfurt](#) | [JAP](#) | [Datenschutz](#) | [Impressum](#) | [Kontakt](#)

Open Access

Issue EPJ Web Conf.
Volume 279, 2023
Nuclear Physics in Astrophysics – X (NPA-X 2022)

Article Number 11011

Number of page(s) 5

Section Nuclear Reactions

DOI <https://doi.org/10.1051/epjconf/202327911011>

Published online 22 March 2023

EPJ Web of Conferences 279, 11011 (2023)

<https://doi.org/10.1051/epjconf/202327911011>

The ASTRAL database for neutron-capture nucleosynthesis studies

Diego Vescovi*, René Reifarth, Enis Lorenz and Andreas Elbe

Goethe University Frankfurt, Max-von-Laue-Strasse 1, Frankfurt am Main 60438, Germany



<http://astral.chetec-infra.eu>

Summary / Conclusions

Nuclear astrophysics relies on nuclear input for modeling of the galactic chemical evolution - crucially **reaction cross sections** to calculate reaction rates.

Measured reaction cross section data typically need to be **compiled, evaluated, and often extrapolated** to provide the necessary reaction rates.

Evaluation and extrapolation requires modeling, hence **nuclear input parameters**.

Mix of established (literature) compilations and newer (online) databases, between stable references and quick incorporation of new experimental results.

– Require **continuous efforts**, to provide up-to-date data to the community.

Periodic efforts for reaction rate evaluations from **within the nuclear astrophysics community**, supported by different institutions / projects.

Open data in nuclear experiments still in early phase, but could help to investigate / reconcile contradicting data sets found in evaluations.